

2016 Annual Progress Report and 2017 Program Plan of the Gemini Observatory

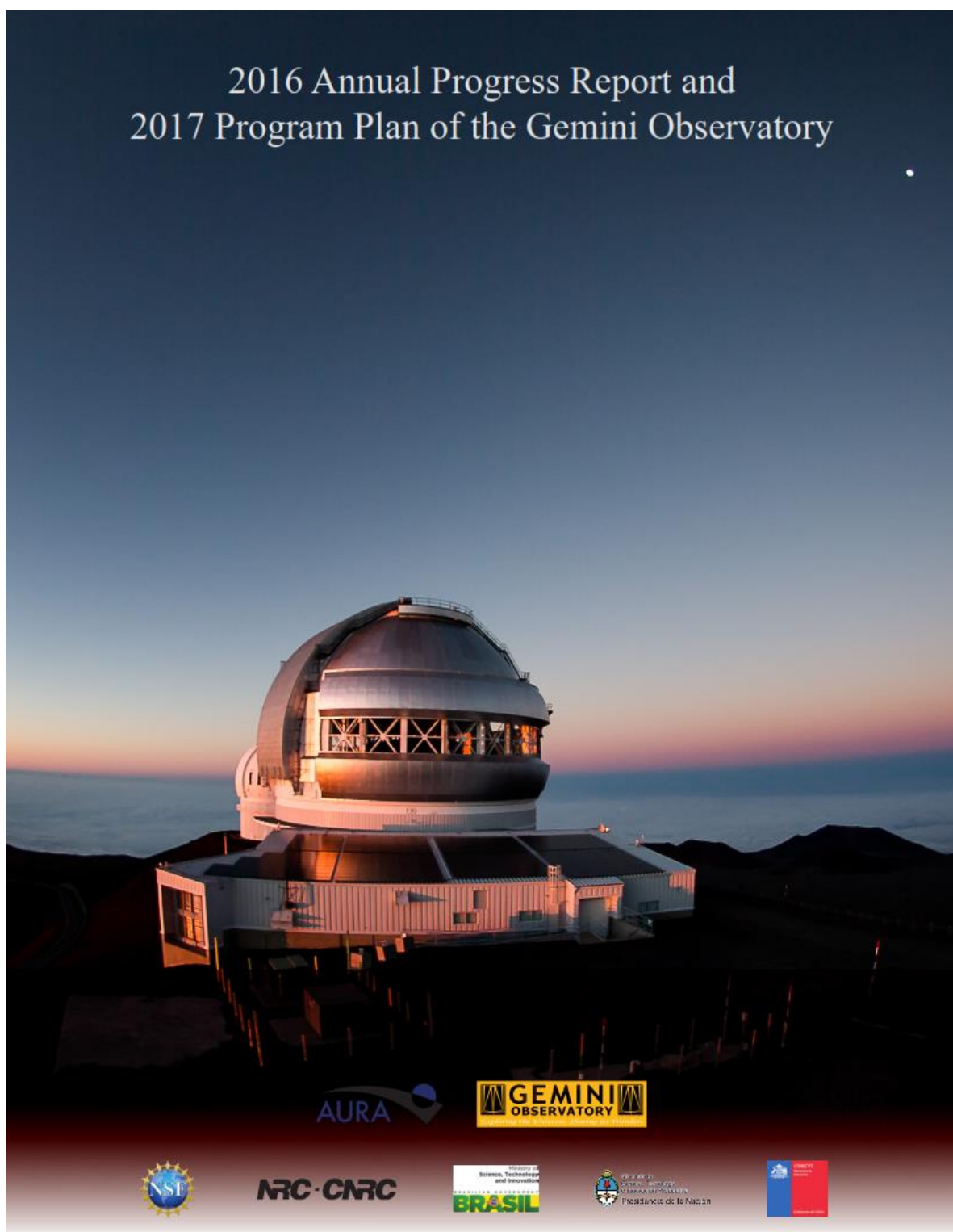


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Guide to Locating Material According to Cooperative Agreement Terms and Conditions

| Requirement | Description | Fulfillment |
|-------------|--|-----------------------|
| a.i | Summary of achievements, including a comparison of actual accomplishments versus goals | 4,5; Appendix I |
| a.ii | Identification of problems faced, their solutions and impact on observatory operations | 4,5 |
| a.iii | List of observing programs, with their investigators, site visitors, observers, and hours devoted to each | Appendix D |
| a.iv | Report on the education and public outreach activities including non-scientific visitor statistics, press releases, etc. | 6, Table 6-1 |
| b. | Scientific accomplishments of the scientific staff, as well as their activities and expectations for the calendar year | Appendices B, E |
| c. | Technical accomplishments of each technical department, as well as the Departments' expectations for the calendar year | 4, 5, 8 |
| d. | Listing of publications and reports produced by staff and, as far as possible, the users of the observatory | Appendices B, C |
| e. | Table showing the division of effort, adding up to 100 percent, for all scientific staff and/or Key Personnel among major activities, such as administration, visiting observer support, scientific research, etc. | Appendix E, Table E-1 |
| f. | Chart or other description of Gemini's organization during the new program year together with an explanation for any changes from the previous year's organization | Appendix F |
| g. | Status report and plan for the new program year on the state of the Observatory | 8 |
| g.i | Scientific plans | 8.2, 8.3, 8.4 |
| g.ii | Detailed technical plans and the potential impact on the observatory of these plans | 8.2, 8.3, 8.4 |
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1 Executive Summary

Gemini Observatory operates twin 8-meter diameter optical/infrared telescopes located on two of the best observing sites on Earth. Locations in the northern and southern hemispheres (Hawai'i and Chile) provide access to the entire sky. A range of instrumentation provides for visual and infrared imaging and spectroscopy, with enhancements from adaptive optics and specialized instrumentation.

Two primary activities dominate the work of the Observatory. The first is in maintaining and supporting operations on behalf of the international scientific community of the Gemini Partnership. The Observatory's goal is to enable our users' scientific progress by being an efficient, nimble, and responsive observatory. We offer a variety of observing and proposing modes to suit the varying needs of our individual Principal Investigators (PIs). Astronomers may visit and conduct their own observations, or have staff execute their observations in a "queue" mode. The queue matches observing conditions to science programs, and provides access to the time domain. Now well-established, the Fast Turnaround program allows monthly proposals; combined with rapid peer review by other PIs, successful proposers can obtain data as early as a month after having their scientific idea.

Gemini's second primary activity is the development of instrumentation and facilities. Development projects provide novel capabilities for users through new facility instruments and upgrades to existing ones, on both small and large scales. The facility instruments and adaptive optics systems remain stable on both telescopes, with improvements to the popular GMOS (Gemini Multi-Object Spectrograph) instruments, including a detector upgrade to increase sensitivity. Visiting instruments increasingly expand the range of capabilities for all users; with a well-established, demand-driven program in the North, 2016 saw the first run of the speckle imager DSSI, and the return (as a visitor) of the high-resolution infrared spectrometer Phoenix to Gemini South.

The Transition Program (TP) is wrapping up with a total realized savings of more than \$1M per year, leaving Gemini sustainable for the longer term. Key ingredients of the TP include the Base Facility Operations (BFO) project, which concluded in 2016 and enabled us to operate both Gemini North (GN) and Gemini South (GS) from their respective base facilities at night. Our efforts in energy savings continued with the installation of solar arrays at both the Hilo Base Facility and on Cerro Pachón. With the end of the transition we are ramping up on activities such as planning for obsolescence mitigation, which will form a key part of our engineering operation over the coming five years.

In addition to these primary activities, we are concentrating increasing effort on the provision of post-observing support for scientists across the Partnership, via the Science User Support Department, while continuing our regular outreach activities. 2016 saw considerable technical and structural improvements to the Data Reduction Users' Forum, and the introduction of non-staff contributors who have reinvigorated the flow of information there. Local outreach programming, including the annual Journey Through the Universe and Viaje al Universo events, kept our local host communities actively engaged. Meanwhile, coordination with the new Science User Support Department (SUSD) expanded dramatically and communications with our scientific users continued to evolve as service to users became an ever more important aspect of all Observatory

staff's core responsibilities. Several members of the Directorate attended the American Astronomical Society meeting in January 2016. The Deputy Director attended La Sociedad Chilena de Astronomía (SOCHIAS) meeting in March and Sociedade Astronômica Brasileira (SAB) meeting in August; the Director attended the Canadian Astronomical Society (CASCA) meeting in June; and the Associate Director for Operations attended the La Asociación Argentina de Astronomía (AAA) meeting in September.

Gemini users continued to generate impressive science results; over the full range of astronomical study enabled by the Observatory. 2016 saw the first publications of data from GRACES, the novel spectrograph fed by fiber from Gemini to CFHT, and an example of the sort of science being done is included in the Science Highlights section of this report. The rate of peer-reviewed publications based on Gemini observations continues to be competitive, and adaptive optics results remain more than 20% of the Observatory's output.

The Development Division focused its efforts on the next facility-class instruments and a new laser for Gemini South. GHOST, the high-resolution spectrometer, proceeded into its build phase in 2016, and the Request for Proposals (RfP) for the next-generation instrument Gen 4#3 was released in May. We received proposals for that instrument in August, and made a selection recommendation to the Board in November. We are now working on contract negotiations. In addition, we signed a contract for a new laser for the Gemini Multi-Conjugate Adaptive Optics System (GeMS) at Gemini South. We expect delivery of the laser by the end of 2016 with installation and integration complete by mid-2017.

We continue to manage the Gemini budget and hit targets as set by the Observatory Governance. The 2016 budget is expected to be spent within the target tolerance range. Tables 1-1 and 1-2 summarize the spending and FTE levels of Gemini Observatory in 2016. Gemini's staffing remains at the post-transition level, with approximately 168 FTEs distributed evenly between the two sites. The principal sources of budget variance were the devaluation of the Chilean Peso, a net decrease of \$1M in labor costs (partly offset by hiring into extra positions), reduced electricity and indirect costs, increasing project costs, and the purchase of a new laser for Gemini North.

| Fund | 2016 Actual |
|-----------------------------------|--------------------|
| Operations and Maintenance (O&M) | \$27,800 |
| Instrument Development Fund (IDF) | \$3,506 |
| Facilities Development Fund (FDF) | \$31,306 |

Table 1-1: 2016 spending by fund, values in US\$1,000

| Division | FTE |
|-----------------|------------|
| Administration | 17 |
| Development | 18 |
| Operations | 113 |
| Deputy Director | 13 |
| Directorate | 7 |

Table 1-2: 2016 Staffing

2017 will bring the full commencement of the Obsolescence Mitigation ("Regeneration") program, new data reduction cookbooks for GNIRS and FLAMINGOS-2, the installation and commissioning of new detectors into GMOS-North, the new GeMS laser at Gemini South, and the completion of the build phase of GHOST.

2 Introduction and Overview

Gemini Observatory's mission is:

To advance our knowledge of the Universe by providing the international Gemini Community with forefront access to the entire sky.

Gemini's foundation is the twin 8.1-meter telescopes on Maunakea, Hawai'i, and Cerro Pachón, Chile, with their complement of instrumentation that provides a range of broad and specialized capability across the optical–infrared bandpass, including adaptive optics. That foundation supports our multi-national astronomical user community, which includes Gemini's own scientific staff. Our users are directly responsible for Gemini's scientific achievements, pursuing interests that range from the Solar System to the most distant galaxies and structure of the Universe. Gemini Observatory provides the flexible, responsive platform that enables these advancements.

The Fast Turnaround program, which allows a wide range of scientific ideas to be turned into results very quickly, is now fully established at both Gemini North and South. It has proved very popular, and the novel peer review mechanism has been appreciated by the community. The second phase of the Base Facility Operations (BFO) project is about to conclude, and we are now operating the Gemini South telescope from the control center in La Serena, Chile, with no staff at the summit at night. BFO is part of Gemini's Transition Program, in which we restructured the Observatory to function robustly despite a budget reduction of roughly 25%. The final report of the Transition Program was delivered to the Gemini Board in its meeting in November 2016. The TP has realized a total of \$1.2M in savings annually. Finally, the Development Division continued the strong progress of GHOST, the high-resolution optical spectrograph destined for Gemini South in 2018, and received proposals for the next instrument (Gen 4#3).

Section 3 describes scientific highlights from 2016. These include results from facility instruments, novel modes such as GRACES, and visiting instruments. Section 4 reports on Operations, including demand, usage and productivity. We also present summaries of user interactions, proposing and observing modes, and Observatory metrics. Section 5 discusses accomplishments in instrumentation and facility development, including progress on new facility-class instrumentation and upgrades of existing instruments. In Section 6, we review our educational programs, our efforts toward broadening participation, and our communications for users and the general public. Section 7 summarizes administration and finance. Section 8 contains the program plan for 2017.

3 Science Highlights

Resolving the Stellar Mass-Size Relationship at $z=1$

Sarah Sweet (Australian National University) and collaborators observed the distant cluster of galaxies SPT-CL J0546-5345 using the Gemini Multi-conjugate Adaptive Optics System (GeMS) with the Gemini South Adaptive Optics Imager (GSAOI). They found a mass-size relationship which is offset from that at the present day, consistent with size evolution proportional to $(1 + z)^{-1.25}$, as has been found in field galaxies. However, the size evolution does not appear to be a function of mass, so neither

mergers nor adiabatic expansion are responsible for this evolution. Sweet et al. demonstrate that to achieve reliable results at this distance requires infrared observations which trace the older stellar populations, and the sub-kpc spatial resolution provided at $z=1$ by the GeMS adaptive optics system. (Sweet et al. 2016 MNRAS, in press)¹

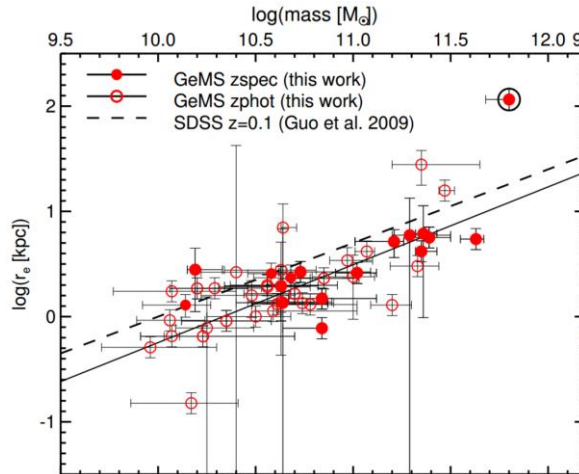


Figure 3-1: Stellar-mass, size relation for the cluster SPT-CL J0546-5345. The relation —where mass is measured in units of $\log(\text{mass } [M_{\text{Sun}}])$ —for this $z=1$ cluster is offset from the local relation (dotted) by 0.21 dex.

A Dark Matter Galaxy

The very-low surface brightness galaxy Dragonfly 44 was discovered in 2014. Fast-Turnaround observations using the Gemini Multi-Object Spectrograph (GMOS) on Gemini North, along with spectroscopic data from Keck II, reveal the galaxy's physical properties. They show that it is like a “failed” Milky Way, having similar mass, size, and globular cluster population, but lacking stars. The Keck spectroscopy enabled Pieter van Dokkum (Yale University) and collaborators to measure the mass of Dragonfly 44. Deep imaging from Gemini then yielded the galaxy's mass-to-light ratio, and also showed the large population of globular clusters in the halo. The team concludes that the galaxy's mass is approximately $10^{12} M_{\text{Sun}}$, and that the total galaxy is 99.99% dark matter. (Van Dokkum *et al.* 2016 *ApJ* 828 L6)²

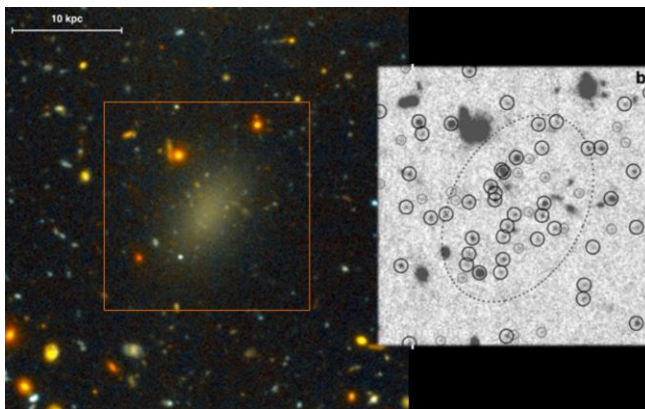


Figure 3-2: Narrow-field (left) view of Dragonfly 44, obtained with GMOS-North. The inset (right) shows the derived distribution of globular clusters.

¹ <https://arxiv.org/abs/1609.06054>

² <http://iopscience.iop.org/article/10.3847/2041-8205/828/1/L6/meta>

The Deepest Ground-Based NIR Color-Magnitude Diagram

Adaptive optics enables accurate and extremely deep photometry of crowded fields. Sara Saracino (University of Bologna) and colleagues measured thousands of stars in the globular cluster NGC 6624 using the Gemini Multi-conjugate adaptive optics System (GeMS) and the Gemini South Adaptive Optics Imager (GSAOI), finding an age of 12.0 ± 0.5 gigayears. The near-infrared color-magnitude diagrams (using the J and K_s photometry) each span more than 8 magnitudes. The team detects the main sequence “knee” for the first time in a purely near-infrared color-magnitude diagram, at $K_s \sim 20$. They find clear evidence for mass segregation, which confirms that NGC 6624 is at an advanced stage of dynamical evolution. (Saracino *et al.* 2016 *ApJ*, in press)³

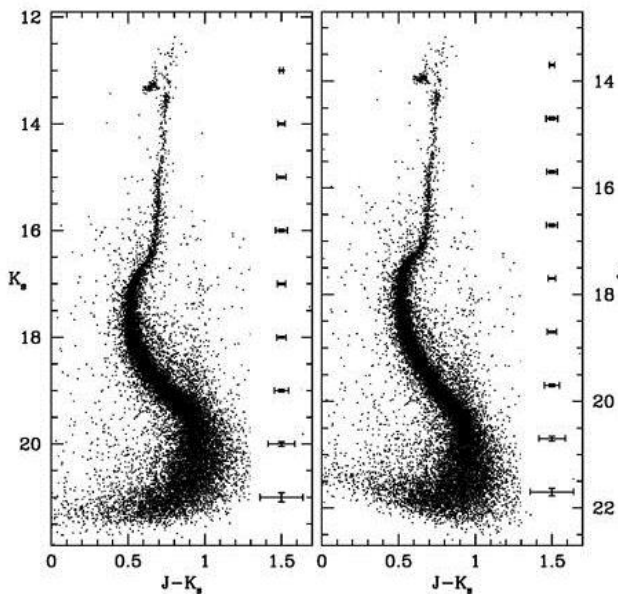


Figure 3-3: Color-magnitude diagrams of NGC 6624 span over 8 magnitudes, and show all the main evolutionary sequences. Photometric errors for each magnitude bin are shown on the right side of each diagram.

A Spectrum of the Coldest Brown Dwarf

Gemini has made spectral observations of the coldest known brown dwarf, which has a temperature of only about 250K. The object, WISE J085510.83071442.5 (WISE 0855), looks as much like Jupiter as any exoplanet discovered to date, and there is evidence for water clouds in its atmosphere. Andy Skemer (University of California Santa Cruz) led this work, using the Gemini Near-Infrared Spectrograph (GNIRS) on Gemini North. The best conditions were required to catch the weak emission, which is five times fainter than any other object detected in ground-based spectroscopy in the $5\mu\text{m}$ atmospheric window: observations were carried out in queue mode over some 13 nights spanning more than a month (Skemer *et al.* 2016 *ApJ* 826 L17)⁴.

³ <https://arxiv.org/abs/1609.02152>

⁴ <http://iopscience.iop.org/article/10.3847/2041-8205/826/2/L17/meta>

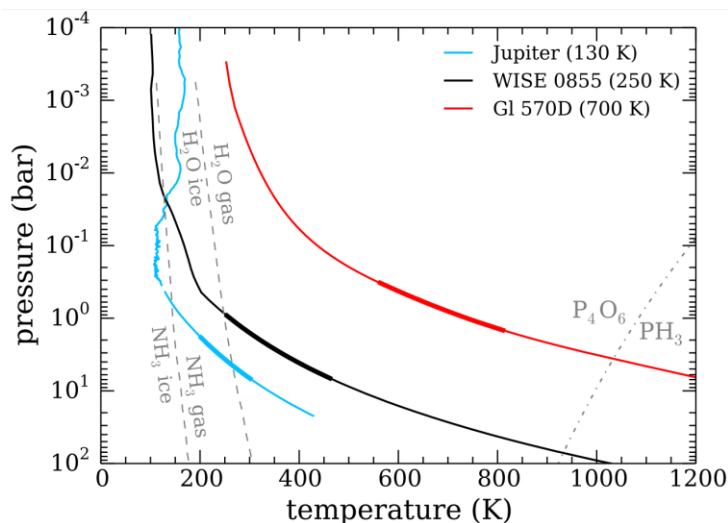


Figure 3-4:
Temperature-pressure profiles of Jupiter, the cold Brown Dwarf WISE 0855, and a somewhat warmer object - GI 570D. Thick lines mark the regions responsible for the 5 μ m emission.

Confirming Nearby Exo-Earths

The Differential Speckle Survey Instrument (DSSI) visited Gemini South for the first time in June 2016 and is already delivering exciting results, including the validation of nearby Earth-like exoplanets. Previous observations using the TRAnsiting Planets and Planetesimals Small Telescope (TRAPPIST) had shown variations in the light curve of the star TRAPPIST-1, implying the presence of several Earth-sized planets. Steve Howell (NASA Ames Research Center) and colleagues used the high-resolution images from Gemini to confirm the small size and mass of these suggested planets by ruling out the presence of a very nearby companion. DSSI on Gemini provides the highest resolution optical images available to astronomers anywhere, and here achieved a resolution of 27 milli-arcseconds, or 0.32 astronomical units at the 12-parsec distance of TRAPPIST-1. (Howell *et al.* 2016 *ApJ* 829 L2).⁵

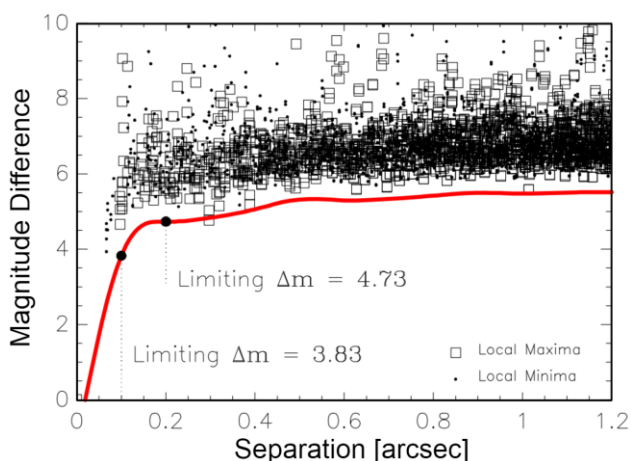


Figure 3-5: Detection limit analysis for the Gemini-South 2016 June 22 observation of TRAPPIST-1 at 883 nm. The red line represents the relative 5 σ limiting magnitude as a function of separation from 0.027 to 1.2 arcsec. At the distance of TRAPPIST-1, these limits correspond to 0.32–14.5 AU. The two listed limiting magnitudes given for reference are for angular separations of 0.1 and 0.2 arcsec.

Measuring a Hot Jupiter with GRACES

Many gas-giant planets lie very close to their host stars. As they could not have formed in their present locations, questions remain: do these giants move close-in when the system is young, after interacting with the protoplanetary disk, or do they only move

⁵ <http://iopscience.iop.org/article/10.3847/2041-8205/829/1/L2/meta>

later, following interaction with multiple planets? The discovery of a $0.77 M_{\text{Jupiter}}$ exoplanet located within 0.06 astronomical units of the young star V830 Tauri confirms both rapid planet formation and early migration. Such early-forming “hot Jupiters” probably play a key role in shaping planetary systems overall. Jean-François Donati (Observatoire Midi-Pyrénées, France) led a program, which included Director’s Discretionary Time observations using Gemini North and the visitor instrument GRACES (Gemini Remote Access to CFHT ESPaDOnS Spectrograph) in collaboration with the Canada-France-Hawai’i Telescope (CFHT). The work also used ESPaDOnS on CFHT and the spectropolarimeter NARVAL on the 2-meter Telescope Bernard Lyot. (Donati *et al.* 2016 *Nature* 634 662)⁶

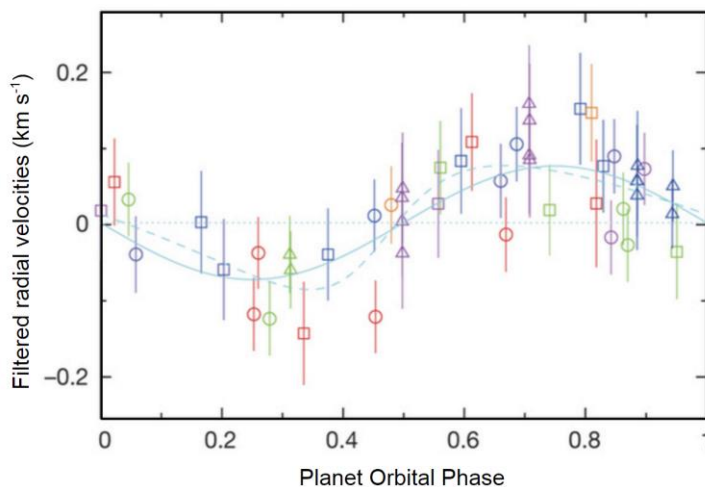


Figure 3-6: Radial velocity measurements of V830 Tauri, after filtering out stellar activity. Data from ESPaDOnS / Gemini (triangles) ESPaDOnS / CFHT (circles); and NARVAL / TBL (squares) are color-coded by rotation cycle. The best fit circular (solid line) and eccentric (dashed line) orbits are shown, with the zero radial velocity (dotted line).

3.1 Proposal and Publication Statistics

Gemini serves hundreds of Principal Investigators (PIs) each year, offering a variety of different program types and observing modes. We list the total numbers of approved programs by proposal mode for semesters 2015B and 2016A in Table 3-1. National Time Allocation Committees evaluate “Semester” programs, which once accepted, may be executed in queue or classical mode. Project titles and PIs in these completed semesters are listed in Appendix D.

As described above, Partner community astronomers use the Gemini facilities effectively to make scientific discoveries. Overall, the publication rate based on Gemini observations continues to be strong, although we may have reached a plateau. A total of 188 refereed publications were published in 2016 (Figure 3-7). 92 of these (49%) resulted from observations with Gemini South; 72 (38%) from Gemini North, and 24 (13%) from observations with both telescopes. Approximately 40% of all publications included data from more than one Gemini observing program.

The use of adaptive optics remains strong, with more than 20% of these publications utilizing Gemini adaptive optics capabilities. 15% of these 2016 publications have a Gemini Observatory staff member as lead or co-author. All Gemini staff publications are listed in Appendix B, and all publications based on Gemini data are listed in Appendix C.

⁶ <http://www.nature.com/nature/journal/v534/n7609/full/nature18305.html>

| Proposal Mode | N |
|-----------------|-----|
| 2015B | |
| Semester | 166 |
| Large and Long | 10 |
| Fast Turnaround | 37 |
| Discretionary | 10 |
| 2016A | |
| Semester | 181 |
| Large and Long | 11 |
| Fast Turnaround | 29 |
| Discretionary | 16 |

Table 3-1: Number of programs per proposal mode for completed Semesters

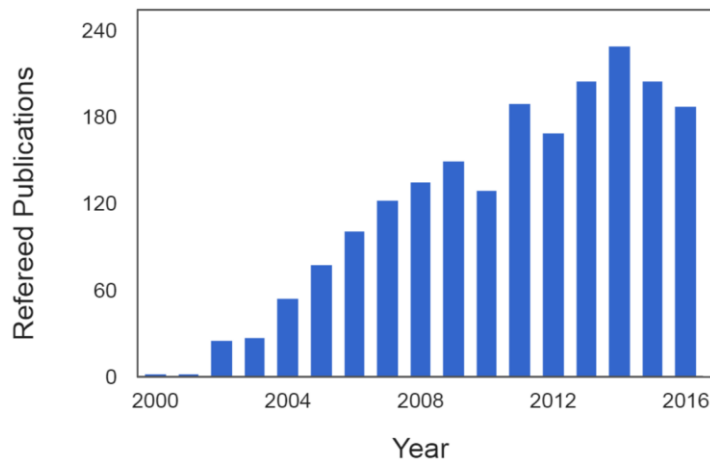


Figure 3-7: publications per year, to date

4 Operations

4.1 Operations Summary

The 2016 program plan provided a list of “regular operations” activities and tabulated specific additional major tasks scheduled for 2016. Tables 4-1 and 4-2 show those major items; the main tasks are complete. Throughout this section we describe progress with these activities, list changes and decisions taken within the year, and give highlights of the regular operations activities.

| Title | Completion | Section |
|--|------------|---------|
| Develop an Observatory-wide Obsolescence Mitigation Plan | 90% | 4.8 |
| Operations Software Upgrades and Obsolescence Management | 70% | 4.8 |
| Science Operations Model Upgrade Prioritization | 100% | 4.8 |
| Upgrade the Cerro Pachón Network Link | 100% | 4.8 |
| Commission the FLAMINGOS-2 MOS mode | 50% | 4.2.2 |
| Implement new post-observing communication strategy | 75% | 4.3 |
| Improve data reduction software documentation and cookbooks | 100% | 4.3 |
| Complete final imaging mode of QAP and release to public | 90% | 4.3 |
| Quick-look tool for quality assessment of spectroscopic observations | 100% | 4.3 |
| Repackage Ureka using Conda | 100% | 4.3 |
| Update Gemini IRAF for new GMOS-N CCDs and F2 MOS mode | 50% | 4.3 |

Table 4-1: Major Operations Activities in 2016

| Title | Completion | Section |
|--|------------|-------------|
| Remedial action on NIRI and F-2 cold heads | 100% | 4.2.1,4.2.2 |
| Support DSSI and Phoenix on Gemini South | 100% | 4.8 |
| Port existing vibration monitoring system to Linux | 100% | 4.8 |

Table 4-2: Other Operations Activities in 2016

4.2 Instrumentation

Here we summarize operational performance of facility instrumentation over the last two complete semesters (2015B and 2016A). See section 7 for information on instrument upgrades.

4.2.1 Gemini North

GRACES

GRACES, a fiber feed from Gemini North to Espadons at CFHT, has proven to be more sensitive than HiRes at Keck for observations between 600nm and 980nm. In 2015B the instrument had two successful 10-night runs. Three successful runs from this popular instrument in 2016A resulted in 81% of the allocated band 1-3 GRACES program time being executed. Improvements to operational efficiency were made by the GRACES team.

Gemini North Laser Guide Star

Faults associated with the laser increased through 2015B, and there were serious technical issues throughout 2016A. The effects of these problems were compounded by poor weather to significantly impact the completion of science programs. We are discussing with the Operations Working Group the possibility of restricting laser programs to Band 1 until the existing laser is replaced with a Toptica system in 2018.

GMOS-N

GMOS-N has been performing well without any major technical issues. The instrument has operated without the R150 grating since June 2016, when the old grating with degraded throughput was removed. The replacement grating is scheduled to be installed in November 2016. In the second half of 2016, the incidence of “high read-noise” events increased somewhat; these had already been noted in 2015B and caused the loss of about 8 hours accumulated over the semester. We expect these issues to be remedied by the Hamamatsu CCD upgrade in early 2017.

NIRI

The facility near-infrared imager, NIRI, has suffered progressive failure of mechanisms over the past few years, resulting in the decommissioning of a number of modes. After a further failure of the aging cryocoolers in October 2015, it was determined that the standard practice of replacing only the cooler inserts was no longer sufficient and in March 2016 we opened the instrument to replace the coolers completely. In the process, two mechanisms which had failed in 2010 were restored to free motion, but these once again proved unreliable in subsequent operation. In August the NIRI pupil viewer primary position sensor failed, and this mechanism is now relying on the backup sensor.

NIFS

NIFS, the facility Near-IR Integral-Field Spectrometer, performed well throughout the reporting period. The main fault on sky was an invalid positioning of the focal plane mask in 2016A, which occasionally produced significant loss of counts in science data and calibrations. Little time was lost to this fault.

4.2.2 Gemini South

GMOS-S

The instrument operated reliably in 2015B and 2016A; though within 2016B significant bias structure was seen in one of the Hamamatsu detectors. This is still being worked on. The known problem with oil leaking from the collimator optics is still present and will require a major intervention. Inspection of science data in 2016 showed that the problem is at least not deteriorating and we now plan to carry out the work in 2017.

FLAMINGOS-2 (F-2)

The instrument went through a major and very successful repair and maintenance stand-down in May 2016. The failing cold heads were replaced and the On-Instrument Wavefront Sensor (OIWFS) was brought back to operability, although weather prevented much progress in its characterization and a later fault (after the end of the semester) returned it to remedial work. Commissioning of the MOS mode depends on the OIWFS, and is thus delayed to the end of Semester 2016B/start of 2017A.

GPI

GPI (Gemini Planet Imager) saw fairly reliable operation throughout 2015B and 2016A. Active vibration cancellation at the closed-cycle refrigerators (CCRs) has been working very well. We operated the unreliable pupil-plane mask (PPM) mechanism on a few nights in May to allow use of the Non-Redundant Mask (NRM) mode. The PPM mechanism has since been fully restored. With the practically complete removal of the vibrations from the CCR's, we have been able to test the High Order Wavefront Sensor (HOWFS). We have found that the fringes in the HOWFS are being smeared by a 100 Hz vibration originating outside of the GPI; this has been narrowed down to either the F-2 CCRs or the helium supply lines to them. More work is needed to fully understand and mitigate this problem.

GeMS/GSAOI

We executed none of three scheduled GeMS blocks in 2015B due to the major earthquake of September 2016. This seriously misaligned the GeMS laser and caused some damage in the Canopus AO bench. By 2016A, we were still suffering from variable and low laser power but attempted to resume science operations; the weather then became the limiting factor and very little science was actually carried out. The first observing block in 2016B had much better technical performance, but again was affected by weather. We look forward to 2017, which will bring with it a more reliable laser system and the potential for more regular scheduling of GeMS.

4.3 Science User Support

The Science User Support Department (SUSD) is tasked with creating a collaborative community of users and staff and consolidating post-observing support. Areas of attention include data archiving and reduction, interactions with the National Gemini Offices (NGOs), and oversight of the HelpDesk and Data Reduction User Forum systems. Increased communication and interaction is at the core of creating a collaborative community. We have been working to implement a strategy to improve communications with our users. Two projects that form the basis for this strategy were started this year:

1. We have begun to restructure the Gemini public website, making it more user-centric. The first stage of the work consists of research and pre-assessment. A

workgroup identified our audiences, elaborated a global vision statement, and drafted a website policy document. We conducted a series of usability tests on the current website, to identify the main improvements and features required to address our users' needs.

2. We plan to collect trackable and actionable feedback from Gemini users through the use of regular surveys. A short survey format (2-3 questions) will motivate a higher level of participation. The questions will be invariant each semester, so answers can be compared over time. The surveys will be timed to coincide with the four phases of the science program lifecycle (proposals, preparation, execution, data reduction) and repeated each semester. The results will be analyzed every semester beginning 2017, and monitored long term.

We have also concluded a series of technical improvements to the Data Reduction User Forum, and begun a program of care and feeding using specially-designated user contributors drawn from the community. The number of posts, and the interaction between users, continues to grow. Finally, requirements have been set for a replacement Helpdesk system. The current antiquated system is awkward to use, lacks important features (e.g. search), and is difficult to maintain.

With the completion of the Quality Assessment Pipeline (QAP) Transition Project, we have turned our primary focus to development of data reduction software for users. IRAF is growing obsolete and we are working to extend the Python infrastructure developed for QAP to meet the requirements of science-quality data reduction. Additionally, a standalone Python utility for GSAOI distortion correction and stacking (Disco-Stu) was released in August. This partially resolves pent-up demand from our user community for better GSAOI data reduction tools that specifically handle the distortion correction.

We are collaborating with STScI to repackage the STScI- and Gemini-developed software within Ureka using Conda. Use of Conda will allow the user to stay more up-to-date with third-party Python packages such as numpy and matplotlib (among hundreds of others) and decrease effort to release and maintain subsequent versions of our software. The first Gemini Conda release is scheduled before the end of the year.

In addition to improving our data reduction tools, users also request improved data reduction documentation. While addressing this in the web page project by making existing documentation easier to find, we are also working to make additional documentation available via staff authors or in collaboration with the NGOs. As part of our agreement with the US NGO to provide data reduction support in trade for phase II support, they have recently made a GMOS cookbook, FAQ, and observation preparation guide available to the Gemini community.

4.4 Storage and Archiving

The Gemini Observatory Archive (GOA) operated with essentially zero unplanned downtime since its release December 2, 2015. There has been one planned interruption of service of two hours on August 10, 2016 to release a new version that included support for GRACES data, miscellaneous file type support, some bug fixes, and minor improvements. As of late November, we have 520 registered users (it is not necessary to register to access non-proprietary data). With close to 12 million hits to date this year, typical usage is approximately 450 searches and 5GB of data downloaded per day.

The next update, to occur late in 2016, will include new features like PI name searching, publication tracking, and better support in the user interface for reduced data products.

| Quarter | Queries | Site Hits | Downloads (GB) | Files ingested |
|--------------|---------|-----------|----------------|----------------|
| Dec 2015 | 10941 | 217434 | 109 | 62765 |
| Jan-Mar 2016 | 23039 | 495332 | 373 | 141815 |
| Apr-Jun 2016 | 30376 | 3019684 | 756 | 116968 |
| Jul-Sep 2016 | 67283 | 4871447 | 528 | 100855 |

Table 4-3: Statistics for the Gemini Archive in 2016

4.5 Special Observing Modes

4.5.1 Fast Turnaround

2016 marks the first full year in which the Fast Turnaround (FT) program operated at both sites. The program has now been used by PIs from all partner countries, and for all of the purposes anticipated when the program was originally set up. Furthermore its novel “peer review” process has given many graduate students their first experience of proposal assessment. Proposal loads, quite variable in 2015, have now settled to a level around 15 per month. Oversubscription rates are typically between 1.5 and 2. These vary because the requested conditions cannot always be accommodated and so we do not always fill the allocated 20 hours per month.

4.5.2 Large and Long Programs

In 2016A, eleven Large/Long Programs (LLPs) were active. Five more were added via the 2016 call for proposals, covering topics as diverse as main-belt comets and supernova cosmology. The oversubscription rate for observations requested by LLPs in Semester 2016B was ~6.1, comparable to ~5.5 in 2015B.

All active LLPs submitted annual reports to the LLP time allocation committee (LPTAC) in late April. The LPTAC recommended that two programs be granted extensions through 2018A; these were accepted by the Gemini Director. In Semester 2015B, 93% of the allocated LLP time at Gemini North was executed (excluding targets of opportunity). 60% of the allocated LLP time at Gemini South was executed, consistent with completion rates in the regular queue and reflecting a very difficult semester with instrumentation problems and the impact of the September 2016 earthquake. In Semester 2016A, 91% of the allocated LLP time at Gemini North was executed, excluding ToOs. 57% of the allocated LLP time at Gemini South was executed, quite a good return in a semester in which fully 55% of the observing time was lost to weather. In the second half of 2016, a quarterly LLP newsletter was launched to provide an additional, regular, means of communication to our existing LLPs. The newsletter reminds LLP PIs of upcoming deadlines, offers information on Gemini processes and procedures impacting LLP programs, etc.

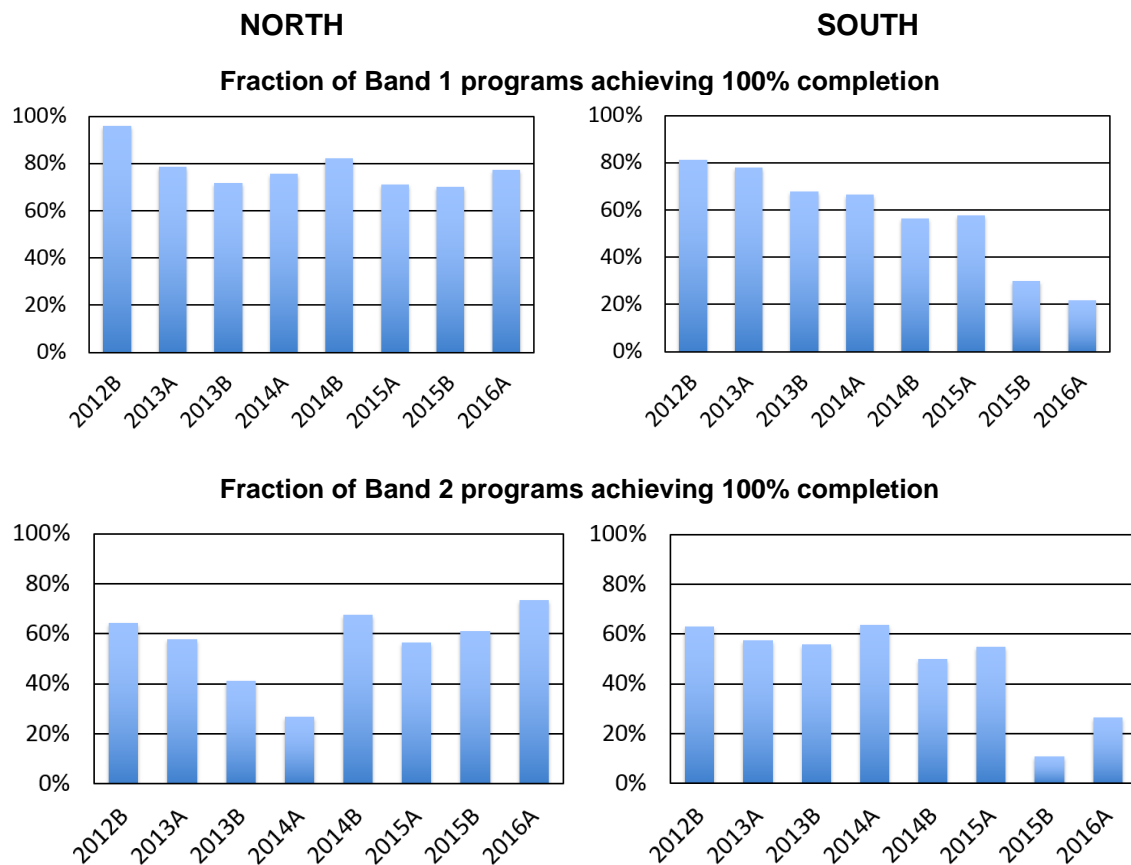
4.6 Telescopes and Enclosures

Gemini’s domes have two shutter panels: the larger, upper panel which covers the majority of the dome slit when closed, and the smaller, lower panel which covers the lower part and pulls up the wind blind when moved upwards. In August 2016, a drive box bearing on one side of the Gemini North lower shutter failed. It was repaired over a 3 week period, using a spare drive box. No science time was lost to the repair, as we

brought forward work from the scheduled October maintenance shutdown. The drive box on the opposite side of the shutter was inspected, and its bearings were also showing signs of premature wear. On the basis of calculations of expected lost time, we determined that it would be better to pin the bottom shutter in a parked position and defer the long shutdown required to effect a complete fix until mid-2017. Until then, we will operate without wind blind protection. After the upper shutter drive box failures of 2014, we redesigned the torque arms which relieve stress on the bearings; this design was prototyped at Gemini South and has been installed and trialed there. In mid-2017 we will also install these new torque arms at Gemini North.

4.7 Operations Metrics

The statistics in this section refer to demand and performance in the last two complete semesters: 2015B and 2016A, given in some cases in context of recent years. First, Figure 4-1 shows program completion rates in the regular semester queue. The major problems encountered in Chile due to the 2015 earthquake and the loss of more than half of the 2016A semester to weather caused very low completion rates at Gemini South. In the North, Band 1 completion rates are stable and the completion rate in Band 2 has rebounded well from its low in 2014 (due to the impact of shutter failures).



Fraction of Band 3 *started* programs achieving 75% completion

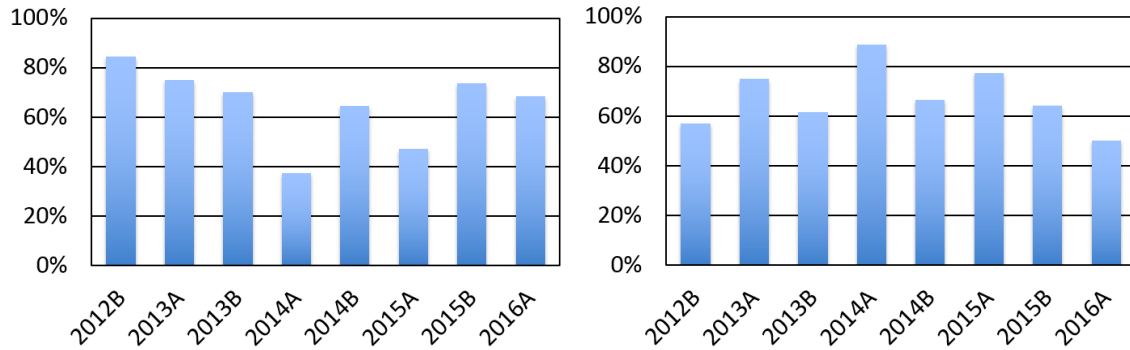


Figure 4-1: Completion statistics for GN and GS. 2015B and 2016A are represented by the bars at the right of each group. Band 1 completion rates in the final two semesters will increase as some Band 1 programs have rollover status.

4.7.1 Telescope Time Usage

Tables 4-4 and Table 4-5 show top-level time and fault distributions and science usage in the most recent two complete semesters.

| Semester | Site | Science | Engineering & Commissioning | Weather loss | Fault Loss | Shutdown |
|----------|-------|---------|-----------------------------|--------------|-------------------|----------|
| 2015B | North | 66% | 2% | 24% | 3% | 5% |
| | South | 40% | 2% | 40% | 3% | 15% |
| 2016A | North | 77% | 1.9% | 18% | 2.8% | 0.0% |
| | South | 41% | 1.9% | 55% | 1.4% ^a | 0.6% |

^a This number may be artificially low due to the cancellation and weathering-out of observing with more recent and more challenging instrumentation

Table 4-4 : Overall operational statistics, semesters 2015B and 2016A

| Semester | Category | North | South |
|----------|-----------------------------|-------|-------|
| 2015B | Computer/Software | 13% | 7% |
| | Instruments & AO Facilities | 45% | 26% |
| | Telescope and enclosure | 42% | 67% |
| 2016A | Computer/Software | 13% | 15% |
| | Instruments & AO Facilities | 58% | 52% |
| | Telescope and enclosure | 29% | 33% |

Table 4-5: Categorized fault distribution, semesters 2015B and 2016A

4.7.2 Regular Semester: Oversubscription and Demand

Oversubscription rates in the regular Semester process are shown by Partner in Figure 4-2. These rates are calculated as the ratio of the total time PIs request to the total advertised available science time. These rates do not account for the time required for standard calibrations. Such calibrations vary by instrument, resulting in an effective factor of between 5% and 10% that is not part of the oversubscription rates shown. (The difference between the telescopes is a consequence of the different instruments used, with the infrared instruments generally having larger overheads for calibration.) In addition, because programs are approved to fill 80% of the queue time available, these values further underestimate the effective oversubscription (*i.e.*, a proposer's success

rate) by a factor of 25%. The distribution of time requested by instrument is shown in Figure 4-3. Note the significant fraction of Gemini North taken by visiting instruments (DSSI and TEXES) in the B Semester.

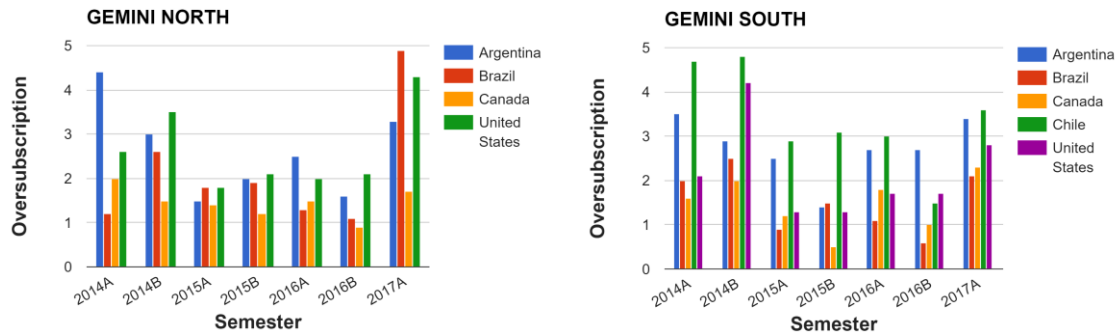


Figure 4-2: Oversubscription by partner over the last three years. 2017A values are provisional as submissions have only just been made. Note that Chile has access only to the South.

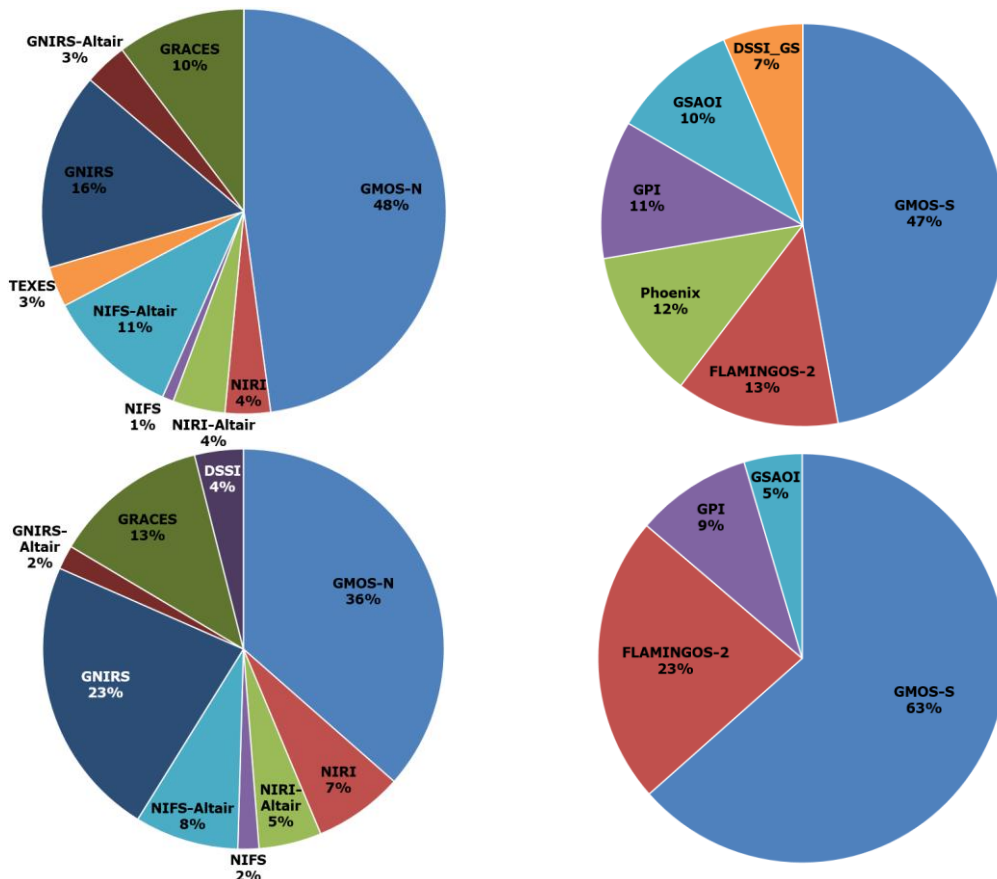


Figure 4-3: Distribution of demand by instrument, for Semester 2015B (upper) and 2016A (lower).

4.8 Other Operations Activities

Visiting instruments

We continued to welcome visiting instrumentation to both telescopes. In 2016, Phoenix had its first run as a visitor at Gemini South, and DSSI, used successfully on Gemini North since 2012, made its first visit to Gemini South in 2016A. In November, a new instrument (POLISH-2) visited Gemini North. This instrument is designed to enable parts-per-million polarimetry, with the science goal of detecting scattered light from exoplanetary atmospheres.

Science Operations review

A review of Science Operations was carried out in 2015 and early 2016, with all reports from the six working groups completed by mid-year. The Directorate then carried out a process of prioritization, resulting in a first round of changes and improvements. These are in most cases already either under way, or in preparation for activity in 2017.

BFO

Operations from the base in Hilo continued throughout 2016. No significant time has been lost. Operations from the base in La Serena commenced in November 2016. In preparation for the start of Base Operations in Chile, we installed a new microwave link between La Serena and Cerro Pachón; this is preliminary to the installation of the LSST fiber link which we now expect to be able to access in 2017.

Visiting observers

In 2016 a total of about one month per semester has been carried out by visiting observers, either Large/Long program PIs in Priority Visitor (PV) mode or visiting instrument teams) at both telescopes. In PV mode, there does not appear to be any significant difference in data quality between visiting observers and staff.

Obsolescence mitigation

After more than 15 years of night-to-night operation, Gemini's telescopes and associated systems face numerous obsolescence issues which must be dealt with to ensure future efficient and fault-free observing. In the first half of 2016 Operations and Development produced a plan which laid out the currently-known obsolescence issues, outlined and costed the work to be done to address them, and briefly described how a future rolling obsolescence program will work. This plan was described in a report to NSF in July 2016. In an ongoing project dealing with obsolescence in the Real-Time systems, we completed the consolidation of common software libraries for the real-time systems, with a final review in October. With this in place, the upgrade stage has started. We plan to complete the upgrade of the simpler systems (Weather System, Acquisition and Guidance Sequencer, Calibration Unit) by end of 2016, as originally planned.

Vibration Monitoring

Monitoring equipment similar to that employed in the North has been procured, readout of the accelerometers under Linux has been completed. The system is being readied for installation on Gemini South in 2017. Once that is done, we will retrofit the computer and software to the North, where the initial prototype computer system has proved somewhat unreliable.

4.9 Transition Program

The objective of the Transition Program (TP) is to produce a sustainable, scientifically competitive observatory running sustainably on an operations budget that has had its buying power reduced by nearly 25%. Since 2013, we have been implementing spending reductions of \$6.5M compared with the 2012 budget. Gemini management examined the Observatory's core mission, identified the required principal services, and restructured the operations accordingly to define the program.

Financial savings result from three TP activities:

1. Reducing staff ("Labor" in Fig. 4-4);
2. Generally reducing non-labor expenses where possible ("General non-labor" in Fig. 4-4)
3. Implementing projects either to support operations with a smaller staff or to achieve additional yearly savings ("TP non-labor" in Fig. 4-4)

During 2016, four Transition Program projects remained active: *Base Facility Operations* implementation at Gemini South, *Gemini North Energy Savings*, *Gemini South Energy Savings*, and *Reduce Base Facility Expenses*. As of the end of 2016 only the *Gemini North Energy Savings* remains incomplete. We closed the Transition Program at the end of 2016, and delivered a closure report to the AOC-G and the Gemini Board at their November 2016 meetings. In the following we briefly summarize the deliverables from the program, as well as its cost and lessons learned.

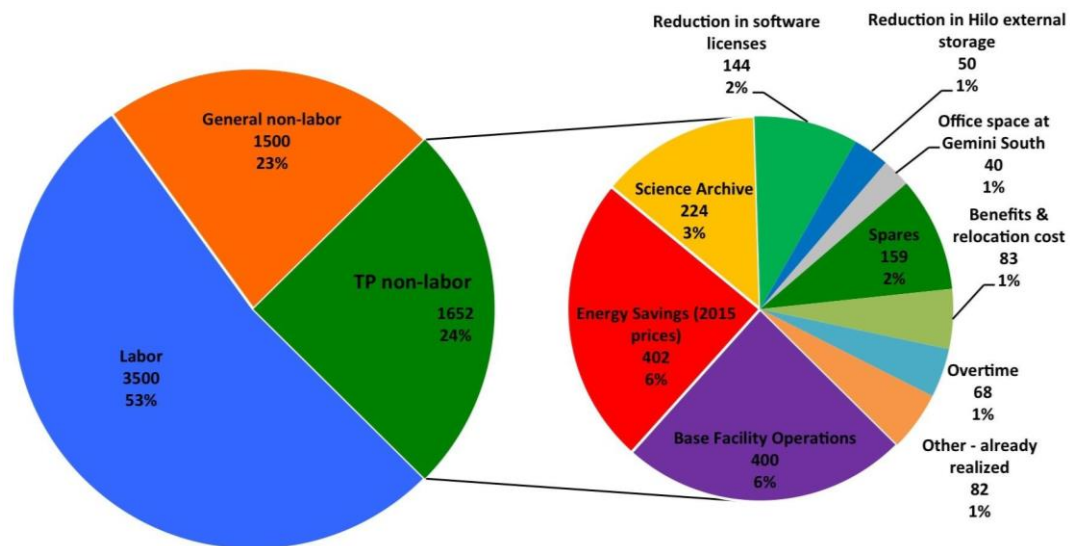


Figure 4-4: Distribution of the O&M budget reductions; amounts are in \$1000. The savings from the Transition Program projects ("TP non-labor" on the left) are detailed on right.

4.9.1 Staffing Reductions

We are currently very close to the post-TP staffing. The total staffing is planned to be 174 FTE, while the actual staffing before the start of the TP (2011) was 201 FTE. The reduction results in operational changes in all divisions. The staffing will support only four facility instruments plus an AO facility at each site, and we will not have the capacity for major instrument refurbishments or development in the O&M budget. Non-research staff carries out most of the queue observing. Support for development of data reduction

software is reduced. Administrative support is reduced at the sites, while procurement, accounting and human resources have been moved to AURA Centralized Administrative Services.

4.9.2 General Non-Labor Budget

Reductions in non-labor budgets not related to Transition Program projects are enabled primarily through two channels and were required to result in savings of about \$1.5M annually:

1. Reductions in staff led to reductions in spending on computers, and supplies & materials directly related to staff numbers.
2. Clear principles for budgeting non-labor expenses as well as top-level reductions were put in place in 2013.

The difference between 2013 actual expenses and the 2016 forecast is ~\$2.8M. This includes the savings from the TP projects of \$1.1M, showing that the general non-labor budget has been reduced by ~\$1.7M, more than meeting the requirement of \$1.5M reduction. A similar calculation for the 2017 budget shows a reduction of \$1.5M again meeting the required reduction on the general non-labor O&M budget.

4.9.3 Transition Program Projects

In Figure 4-5 we show the status of the annual non-labor savings as of the end of 2016Q3. The savings are at \$1.2M annually (red line). By the end of 2016Q4 we expect another \$325k in savings from a combination of GS BFO, Gemini North (GN) energy savings project, and the Southern Base Facility (SBF) lab-to-office conversion.

The Transition Program strategic map (Figure 4-6) shows the status of the projects as of December 2016. Only the GN energy savings project will run into 2017 as we establish contracts for the installation of the fluid cooler and chiller at the Maunakea summit facility and the upgrade of the Hilo Base Facility A/C units.

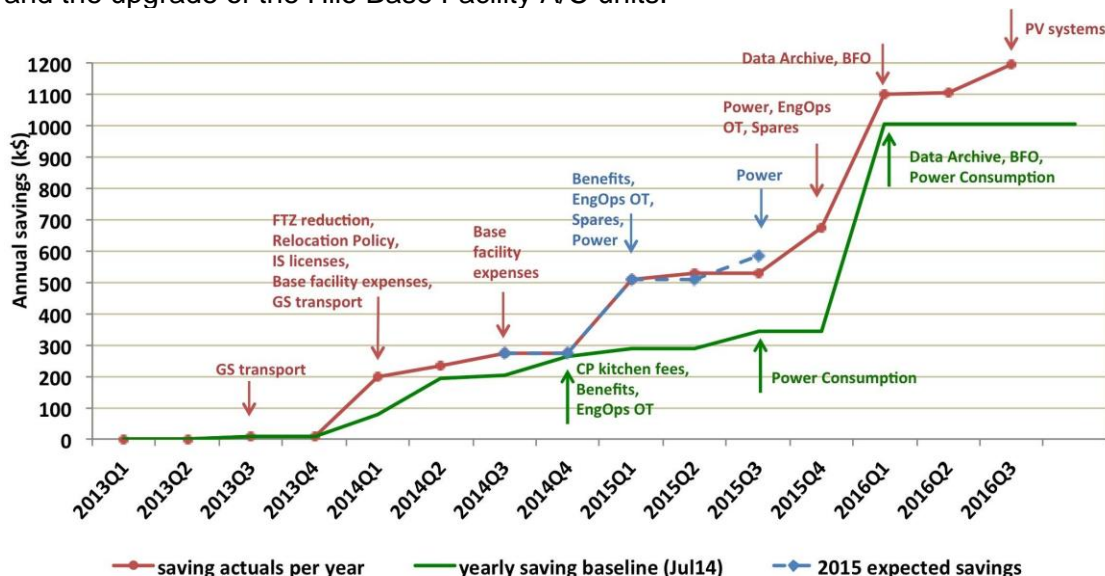


Figure 4-5: Transition Program non-labor savings. Savings at end of 2016Q3 is \$1.2M annually (red solid line). The green line shows baseline savings planned, including discounts for confidence at the time of planning.

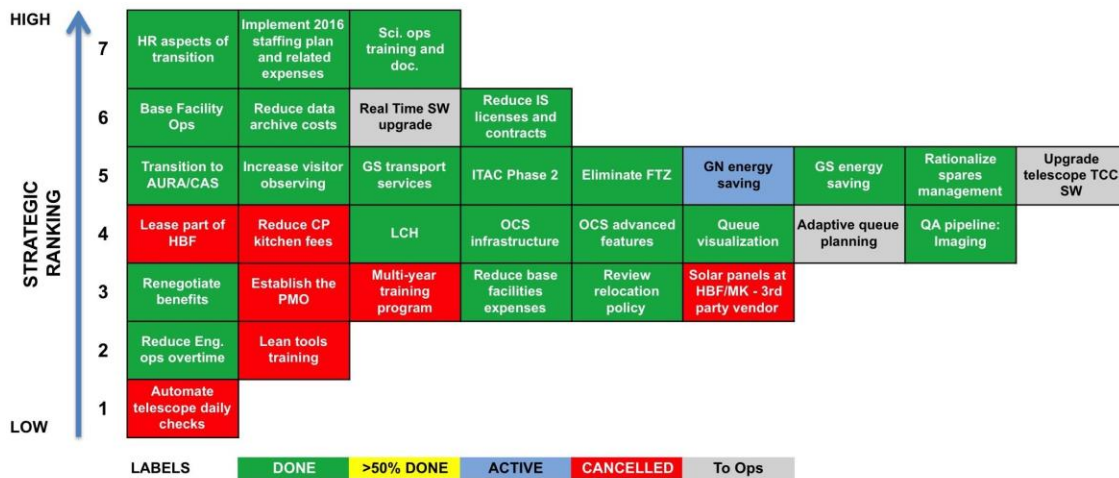


Figure 4-6: Transition Program strategic map as of December 2016.

Base Facility Operations

The implementation at Gemini South is essentially complete. Base facility operations started in mid-November. Formal closure of the project will happen in January 2017.

GN/GS Energy Savings

We have completed the installation of the photovoltaic panels at Cerro Pachón and at the Hilo Base Facility (HBF). At Gemini North, the lighting has been changed from CFL to LED lighting, and all transformers have been replaced as planned. We have the complete engineering design for the work on the Gemini North cooling equipment (summit chiller replacement and HBF air conditioning system). The installation of this equipment will be done under separate contract during 2017.

Reduce Base Facilities Expenses – Gemini South lab-to-office conversion

We have converted the instrument lab at the SBF to offices, enabling us to vacate Casa 8 and Casa Verde and save the related rent and utilities on these buildings.

4.9.4 Cost of the Transition Program

The Transition Program total cost was tracked as non-labor and labor effort. Full tracking using dedicated accounts and timecard categories was put in place by 2013 Q4. Table 4-6 summarizes the total cost in each the six accounts and timecard categories. The 2013 labor effort accounting is incomplete as the tracking was put in place late in the year. We estimated for the 2013 annual report that a total of 10 FTE were used in 2013 on TP projects (2013 Annual Report, Table 7.2). The 2013 effort was primarily spent on BFO, QAP, and software TP projects, though we do not have detailed information on the distribution.

| TP Account | Non-Labor / \$k | | | | | Labor / FTE | | | | | Total Cost / \$k |
|-------------|-----------------|------|------|------|-------|-------------|------|------|------|-------|------------------|
| | 2013 | 2014 | 2015 | 2016 | Total | 2013 | 2014 | 2015 | 2016 | Total | |
| BFO | 50 | 237 | 368 | 281 | 936 | 1.6 | 7.5 | 13.0 | 5.6 | 27.8 | 3711 |
| Software | 7 | 136 | 76 | 6 | 225 | 1.9 | 3.3 | 4.1 | 0.1 | 9.3 | 1158 |
| Engineering | | 56 | 693 | 3017 | 3765 | 0.0 | 0.7 | 1.5 | 1.9 | 4.1 | 4177 |
| Science | | | 6 | | 6 | 0.8 | 2.2 | 3.7 | 0.0 | 6.7 | 679 |

| | | | | | | | | | | | |
|--------------|-----------|------------|-------------|-------------|-------------|------------|-------------|-------------|------------|-------------|--------------|
| Admin | 15 | 5 | | 202 | 222 | 0.0 | 0.3 | 0.6 | 0.2 | 1.1 | 332 |
| Management | 27 | | 52 | | 80 | 0.3 | 1.9 | 1.3 | 0.3 | 3.8 | 457 |
| Total | 99 | 434 | 1195 | 3506 | 5234 | 4.6 | 15.9 | 24.2 | 8.1 | 52.8 | 10513 |

Table 4-6: Cost of the Transition Program

4.9.5 Lessons learned from the Transition Program

The Transition Program Closure report contains detailed descriptions of the lessons learned from the program. Here we briefly summarize these.

- Manage as a program: The TP was originally managed as a collection of individual projects rather than as a program. Program management was put in place by mid-2014 with the change of the TP Executive.
- Manage projects with project managers. Successful program and project execution relies on consistent project management training as well as mentoring for the project managers
- Plan for staff turnover and delay in hiring of temporary staff.
- Realistic estimates of effort. In particular, the software effort was underestimated for the Transition Program.
- Definition and use of risk ledger: Ensure risks are actionable.
- Definition of program dashboard to provide a balanced view of the program progress.
- Tracking of project cost, enabling feedback on resource use as well as cost benefit assessments of projects.
- Tracking of realized benefits, in addition to non-labor savings.
- Communication to the staff: include updates on major programs in the all staff meetings

4.10 Administration and Facilities

The Administration and Facilities Group (AFG) in the North and South provided a wide range of administrative and facility support services to employees and telescope users. The Observatory achieved significant advances in diverse AFG categories in 2016. We cite two examples: electrical energy systems in Gemini North and facilities infrastructure in Gemini South.

Energy projects at the Hilo Base Facility (HBF)

Gemini continues to review energy trends and seek opportunities for energy saving initiatives. Energy savings and reduced electricity costs were achieved through a number of work packages in 2016. Photovoltaic panels installed on the HBF started to deliver savings of 20% in GN electricity usage. Newly installed LED lighting and energy efficient transformers are starting to deliver a savings of \$9k annually. Seven air conditioning units will be replaced with energy efficient units in 2017.

Facilities improvements at the Southern Base Facility (SBF)

In 2016 two new Stulz air conditioning units were installed in the SBF data center, reducing energy consumption by 58.19%. One unit acts as a backup system that runs on UPS power should the primary unit fail. The SBF's building HVAC system was replaced with a more efficient system, saving Gemini 50% in running costs.

In November 2016, construction finished on 17 new offices and a refurbished instrument lab, created in the space formerly occupied by the SBF's original instrument lab. This

enabled vacating the staff from offices in Casa 8 and Casa Verde. Accommodation of 17 extra staff in the SBF main building as well as compliance with health and safety regulations mandated extensions to the current male and female restrooms. These were also completed in 2016.

Under construction is a 25m² extension to the existing instrument lab into the service yard to meet the future needs of the GS Engineering and Development teams. Gemini worked with AURA CAS procurement in establishing an open and transparent bidding process to award this work to an external contractor. The payback period for this project is estimated at 4.4 years.

Overall, this SBF facilities improvement project will provide savings of \$40,000 per year, improve staff interaction, and reduce security risks.

4.11 Safety

The Safety team worked closely with staff and supervisors during 2016, achieving a further reduction of the number safety incidents and resulting lost time (see Figure 4-8). Completion of Level A Safety Awareness training increased to 70% in 2016; this was the result of improved training materials and managerial actions promoting compliance.

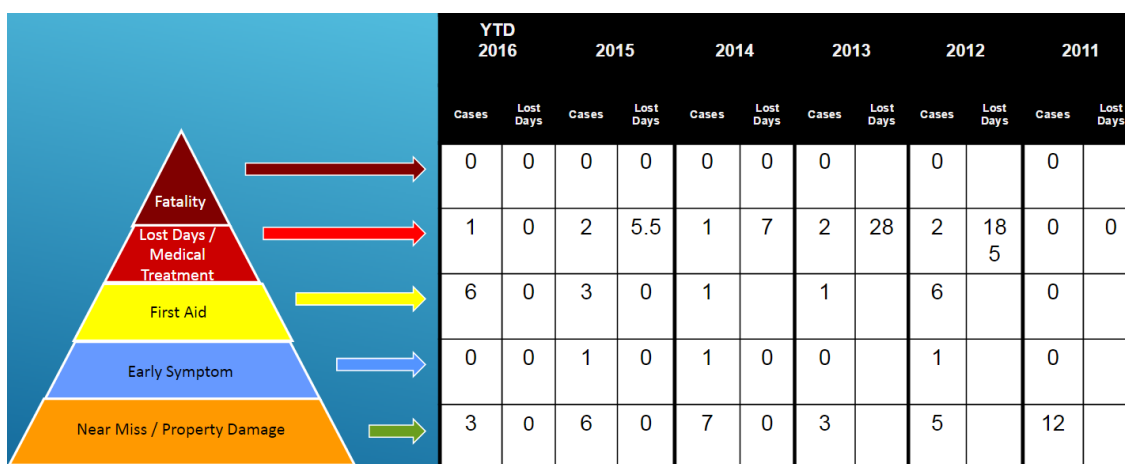


Figure 4-8: Safety incidents for the period 2012 – 2016

As part of the security plan and to support the core value of keeping the safety of people and assets, the Observatory continued training staff on emergency response procedures. Gemini introduced electronic access control at the Gemini South base facility to increase security.

4.12 External Relations

Gemini Observatory remains committed to supporting our users with the broader goal of producing scientific results. We assist them throughout the scientific lifecycle, from initial idea through analysis, leading to final publication of results. As Gemini provides public access to diverse scientific users around the world, we aim to make the Observatory and its capabilities as accessible as possible. We describe here high-level activities between Gemini and our users and Partner countries. We describe the regular communications program for users and the general public in Section 6.3.

In mid-2016, the Gemini Director made an extended “Road Trip” within the US, to provide general information about Gemini’s new operations, instruments, and opportunities. A total of six Institutions were visited:

- University of Maryland
- Penn State University
- University of Texas at Austin
- Texas A&M University
- University of California, Berkeley
- University of California, Santa Cruz

Members of the Gemini Directorate make an effort to attend the national astronomy meetings of our Partners. In 2016, several members attended the American Astronomical Society meeting in Florida. The Deputy Director attended SOCHIAS (Chile) in March and SAB (Brazil) in August; the Director attended CASCA (Canada) in June; and the Associate Director for Operations attended the AAA meeting (Argentina) in September.

We also supported the Gemini Board in their development of Gemini’s strategic vision, setting up a web survey and providing the results for analysis; this information was reported upon at the November 2016 Gemini Board meeting.

Finally, the Director led the work which culminated in signed limited-term partnership Agreements with KASI (for 2017 & 2018) and AAL (for 2017).

5 Instrumentation and Facility Development

The Development Division has three parts: the Project Support Department, the Technology Development Department, and the Program Support Group. Together, they deliver new and upgraded scientific capabilities to the Observatory through our instrumentation, adaptive optics, and technology development programs and support Observatory operations and improvements through our telescope scientist, project managers, and systems engineers.

The Instrumentation program is able to support two large and two smaller efforts each year. In 2016, we saw the Gemini High-resolution Optical SpecTrograph (GHOST) enter the build phase and completed the proposal selection process for our next new instrument, Gen 4#3. We are completing work on the GMOS-N CCD upgrade, with plans to install in early 2017, and we continue our instrument upgrade program with our second annual call in September.

Our adaptive optics (AO) program works to stabilize and improve our adaptive optics facilities, making them accessible to all users, and also maintains telescope performance to deliver seeing improvements for all Gemini observations. This year, we focused on procuring new lasers for both the Gemini South and Gemini North AO systems, along with a new natural guide star sensor (NGS2) for the Gemini Multi-Conjugate Adaptive Optics System (GeMS).

5.1 GHOST

GHOST will provide simultaneous, high-resolution spectroscopy with $R > 50,000$ (for two objects plus sky) and $R > 75,000$ (for one object plus sky only) from 363 to 950 nm. We currently estimate GHOST will be available to the Gemini community for science use in 2018. The GHOST team — Australian Astronomical Observatory (AAO), National Research Council of Canada Herzberg (NRC-H), and the Australian National University (ANU) — completed its final design review in May, 2016 and the instrument is now in the build phase.

5.2 Gen4#3

Following on the successful completion of the Gemini Instrument Feasibility Studies process, we launched the request for proposals for the design and build of Gen 4#3, our next new facility instrument, in December 2015 and received four proposals by the end of August 2016. After an extensive review process, we started final contract negotiations in November. Gen 4#3 will be a broad-wavelength-coverage (at least 0.4 to 1.6 microns), moderate resolution ($R \sim 4000$) spectrograph designed to capture its entire spectral range simultaneously. We aim to start the Conceptual Design Stage in early 2017.

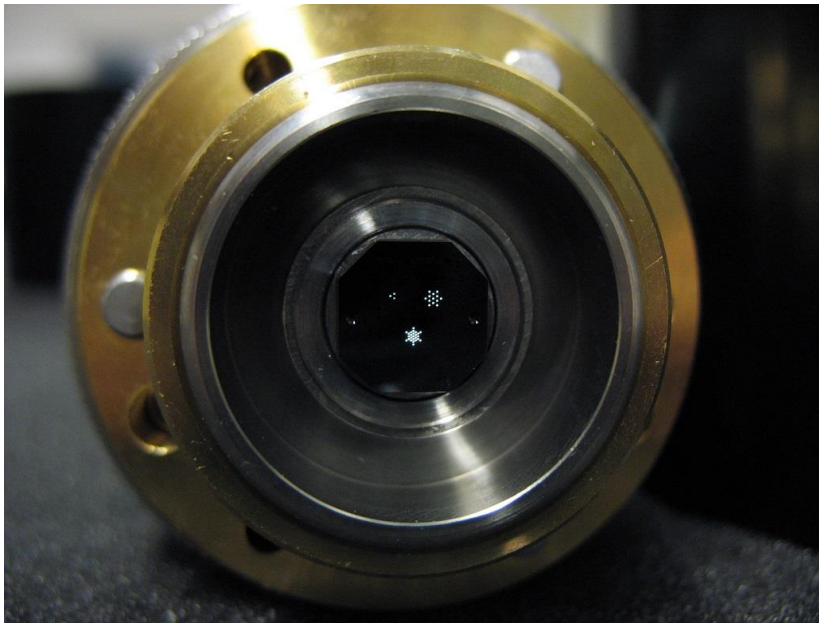


Figure 5-1: The IFU end of one of the GHOST prototype optical cables, completed in 2016. Seen here clockwise from the upper left are the low-resolution sky, low-resolution object, and high-resolution object IFU bundles.

5.3 GeMS

Improved Astrometry

We are pursuing two internal efforts to provide better astrometric calibrations for GeMS. The one closest to implementation is a pinhole mask used prior to each observation to map static astrometric aberrations in the system. We have tested this system in the lab in 2016 and plan to test in GeMS in early 2017.

GeMS Laser

The primary aims for purchasing a new laser for GeMS are to increase its availability and to decrease its support requirements so we can offer GeMS more routinely at Gemini South. In February 2016 we selected Toptica Photonics as our laser supplier. We completed a factory acceptance readiness review in September 2016. Toptica shipped the laser on December 2nd to Cerro Pachón where it arrived safely on December 13th, with none of its 8 shock sensors activating in transport. After installation, configuring, and testing on site, we expect to commission the new laser by August 2017.

NGS2

Throughout 2016, we worked with ANU to deliver a new “Natural Guide Star Next Generation Sensor” (NGS2) to GeMS. The goal for the project is to improve our limiting guide star brightness to enable more sky coverage, while making the system easier to use and support. The ANU team is building a new system to meet these objectives using an EMCCD (electron multiplying charge coupled device) camera built by Nüvü Camēras. We are holding an acceptance review in December 2016 that includes the hardware and some or all of the associated software and documentation. We plan for installation in late 2017, after we commission the new laser, pending a successful review in December.

DM0

GeMS was designed to have three Deformable Mirrors (DMs), conjugated one each at 0, 4.5, and 9.0 km, but during the commissioning, DM0 failed. GeMS now operates with only two DMs (using the former DM4.5 to conjugate at ground level and become the current DM0). A failure of one of the remaining DMs would leave the system either inoperable or delivering very poor performance. In 2016, we completed and oversaw two related contracts to provide a new DM from Xinetics and electronics from Cambridge Innovations. We expect to complete acceptance testing in 2017. Once received, the new components will serve as spares until we have available resources from our other GeMS work (the new laser and NGS2) to install and bring GeMS to its full three-DM configuration, likely in 2018.

RTC

The GeMS real-time computer (RTC) is a complicated piece of hardware that is starting to show some signs of age and operating issues. In 2017, we will study ways we can improve the reliability of the current system or replace it entirely.

Overview

Table 5-1 outlines our expected workplan and major risks for GeMS.

| Milestone | Date | Major Risks |
|-----------------------------------|--------|---|
| Static astrometric mask installed | 2017Q1 | Staff and system availability for testing without rising operational commitments. |
| Laser commissioned | 2017Q3 | Insufficient software effort |
| NGS2 installation | 2017B | Additional work needed after 2016 acceptance testing. |
| DM0 installation | 2018Q2 | Delays in any of the above projects |
| Gems RTC upgrade | TBD | Resource availability |

Table 5-1: Timeline and risks for GeMS development work

5.4 GMOS CCDs

The GMOS CCD upgrade project aims to replace the detectors of both GMOS instruments (there is one at each Gemini telescope) with high-efficiency, fully-depleted CCDs from Hamamatsu. With some additional debugging in 2016 after their 2015 installation, the GMOS-S CCDs continue in standard operation. The GMOS team mounted, tested, and characterized the new CCDs for the GMOS-N instrument for installation planned in February 2017.

5.5 Instrument Upgrades

In order to keep our current instruments competitive as they age and to provide more opportunities for community instrument teams and scientists to work with Gemini, we launched an instrument upgrade program in 2015. In December of that year, we selected a project to add some additional filters to FLAMINGOS-2 that split the K-band. We expect to offer the new filters to the community by 2017B.

In September 2016 we released our second annual call, increasing the funds available to \$600,000 USD to allow for more ambitious projects than did our 2015 call. Going forward, our baseline plan is to have annual calls with funding alternating among \$100,000 USD and \$500,000 USD every year. The 2016 proposal deadline is December 2016 and we hope to start the new projects by the end of 2017Q2.

5.6 Altair

In 2016 we received a replacement dichroic for ALTAIR that extends into the L and M bands, for use with NIRI and GNIRS. This complements the new dichroic we received earlier that allows the use of GMOS-N with ALTAIR. Aided by a new Adaptive Optics Fellow at Gemini North, we plan to commission both dichroics and start a project to replace ALTAIR's aging real-time computer in 2017.

In 2016 we also completed an agreement with Toptica Photonics to purchase an additional laser, like the one ordered for GeMS, for use at Gemini North. The current laser, although sufficient power-wise, is requiring an increasing level of maintenance to keep it running. The new laser will allow us to consolidate spares and expertise, not only among both Gemini sites, but cross-institutionally as well, with ESO, Keck, and Subaru who are also procuring Toptica lasers for their AO systems. We expect delivery of the Altair laser by the end of 2017.

5.7 Additional Development Projects

Our plans to upgrade the Acquisition and Guiding (A&G) units at each site and the detector controller for GNIRS and NIRI remained largely on hold in 2016, due to competing higher priorities. Our planned A&G work includes replacing the obsolete mechanism controllers, and enhancing the performance of our wavefront sensors. We did some testing of our selected A&G mechanism controller to verify that it could control the needed mechanisms. With the GMOS CCD upgrade work completing in early 2017, we will be able to make more progress on both the A&G and GNIRS controller efforts in 2017.

6 Public Outreach and Broadening Participation

Gemini's Public Information (PIO) and Outreach office maintained its strong commitment to core communications efforts and outreach/education programming over the past year, while innovating to meet new and ever-evolving needs. Central to our communications activities is the dissemination of information to our scientific user community through e-publications, social media, conferences, and hard-copy publications. Our non-scientist public audiences benefited from social and electronic media, and local host community outreach programming which grew in breadth and diversity over the past 12 months. An example of the impact from both user and public communications is shown in the high-impact coverage we obtained such as illustrated in Figure 6-1. Also of particular note is a new (Version 2) of Gemini's career brochure which now includes in-depth career profile sheets of selected staff designed to inspire participation in observatory STEM careers by local students and under-represented populations.



Figure 6-1: Gemini North telescope featured on October 15, 2016 as the Astronomy Picture of the Day (APOD). Photo by Joy Pollard of Gemini's PIO staff.

Gemini's PIO efforts highlighted below demonstrate how we effectively engage our users, staff, the public, and our local host communities in the Observatory's scientific and technical work.

6.1 Outreach Programming

6.1.1 Journey Through the Universe & Viaje al Universo

2016 saw the effective delivery of Gemini's flagship local outreach programs to large and diverse audiences in our local host communities in Hawai'i and Chile. The annual *Journey Through the Universe* and *Viaje al Universo* in Hawai'i and Chile respectively, are defining programs for Gemini's local outreach programming and have been ongoing for many years. Both programs continue to grow and diversify (Figure 6-2) in both scope and duration with observatory careers and longer-term impact on students and teachers dominating our efforts during the past year. In 2016 these two programs both expanded well beyond the "traditional" week-long set of activities and now include an extended StarLab portable planetarium programming element for kindergarten and first graders in Hawai'i. In addition, multiple career-awareness events, including very popular Observatory staff Career Panels in both Chile and Hawai'i, are promoting STEM careers

in our host communities (see next section on broader career awareness activities). Work was also initiated in 2016 to establish a formal evaluation of our STEM career awareness programming impact and effectiveness, which will be implemented as a key addition to our program plan in 2017.



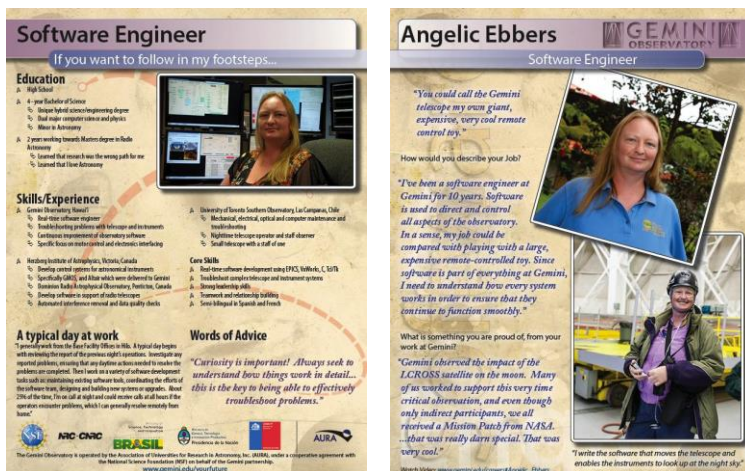
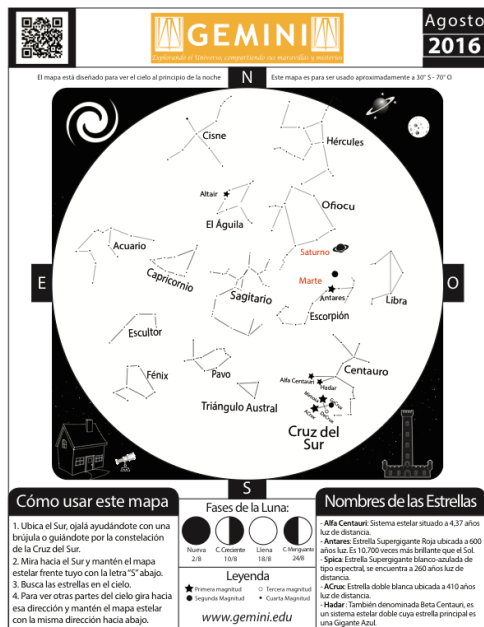
Figure 6-2: Gemini South's Erich Wenderoth speaks to a public audience at the University of La Serena in Chile as part of the extended Viaje al Universo local outreach programming in 2016. During this presentation a video link to the Gemini South control room was established so the audience could experience an insider's perspective on Gemini's facilities and operations.

6.1.2 StarLab Portable Planetarium

Over the prior 12 month period Gemini's StarLab portable planetarium programming has operated completely in a "train and loan" mode where Gemini staff train educators on using the equipment, as well as relevant content in modern astronomy. In addition the equipment has been upgraded to the latest projection system and redundancy to assure reliability in the field. As part of this transformation, Gemini South PIO staff are now producing a monthly sky map for central Chile (Figure 6-3).

6.2 Career Resources

A significant (and growing) component of Gemini's outreach programming is our focus on STEM careers and the participation of local students in the future STEM professional workforce. Over the past 12 months new versions of our two "Career Brochures" (English and Spanish) were produced, printed and made available online – along with new video interviews of selected staff. These are available at: www.gemini.edu/careers. To augment the brochures, a new initiative to produce in-depth career profile sheets is nearing completion; a sample is shown in Figure 6-4. By the end of 2016 we expect to have at least four new in-depth career profiles and more profiles are included in our program plan for 2017.



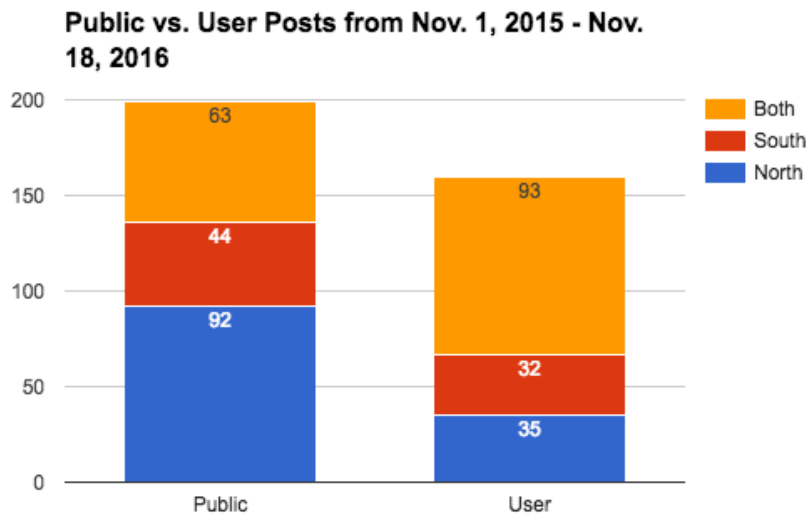


Figure 6-5: Social media posts over past six months showing distribution of posts for the public (left bar) and users (right bar) and how they correspond to either Gemini North, South or both telescopes.

In addition, we continue to track overall impact and engagement levels for our Facebook followers. Engagement continues to grow at a healthy rate (Figure 6-6).

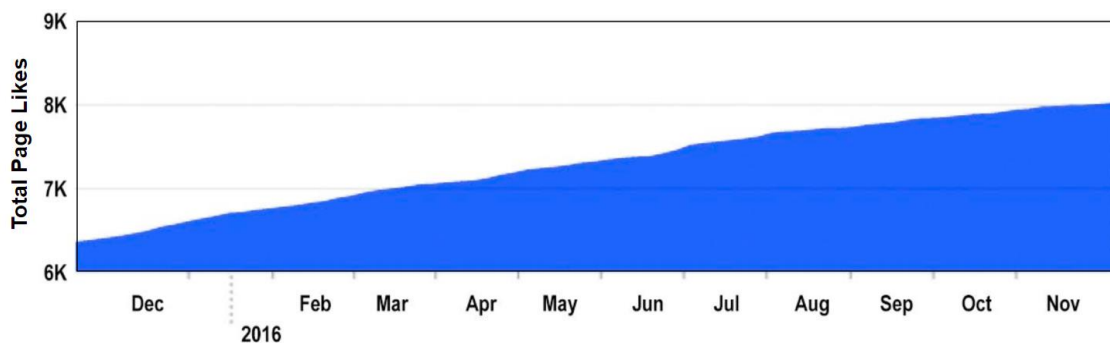


Figure 6-6: Facebook “likes” as an indication of engagement of our followers.

6.3.2 User Communications: Publications and Conferences

Gemini’s PIO group provides critical support in communications with our user community by providing diverse publications that range from our quarterly newsletter (Figure 6-7), to a monthly e-newscasts (<http://www.gemini.edu/enewscast>), e-mail alerts, and printed/online brochures. In addition, we produced a total of 32 “WebFeatures” highlighting recent science results from Gemini (this includes 12 press releases, as described in the next section).

Working closely with Gemini’s Science User Support Department (SUSD), we have extended interactions with our internal scientific staff to better promote timely and accurate information for our user community. Each monthly edition of our e-newscast, as well as the *GeminiFocus* newsletter, now contain regular input from the SUSD.

Support of user communication activities extend beyond publications to include scientific conference support, especially the annual American Astronomical Society (AAS) winter

meetings. The development of conference exhibit materials and publications requires significant resources and over the past year also includes long-term coordination with other optical/infrared NSF/AURA centers for the production of an exhibit “pavilion” projecting a unified look and identity at the AAS meetings. Conference materials and publications are multi-purposed for use at our Partner astronomical conferences as well as other technical meetings such as the annual Society of Women Engineers (SWE).

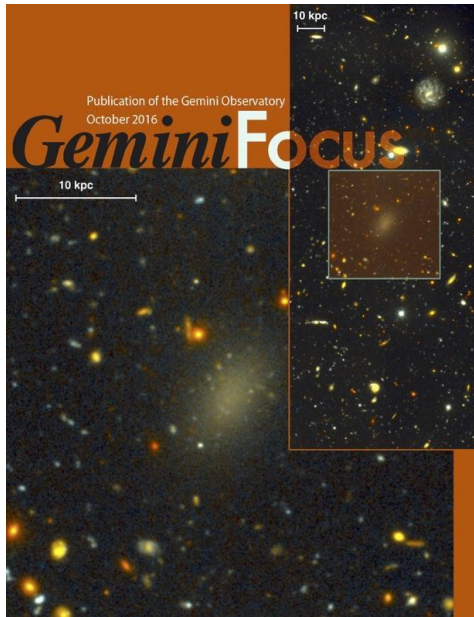


Figure 6-7: Recent issue of the quarterly *GeminiFocus* newsletter, a primary electronic and hard-copy communication product for our user community

6.3.3 Media Relations and Press Releases

We exceeded (150%) our annual goal of eight press releases. The 12 press releases, based on Gemini science, included topics ranging from the coldest known brown dwarf to a galaxy with 99.9% dark matter (shown on the cover of *GeminiFocus* in Figure 6-7). These releases were covered by numerous high-profile publications including blogs and traditional media.

6.4 Additional Opportunities and Milestones

Beyond the accomplishments described in the previous sections, the Gemini PIO efforts, during the period of this report, included other important opportunities and milestones.

Among the opportunities was staff participation in a variety of professional conferences and meetings, including Communicating Astronomy with the Public (CAP), National Science Teachers Association (NSTA), and Society of Women Engineers (SWE) as well as the completion of several staff training opportunities.

Gemini PIO continues to reach milestones set to improve and expand our communication and outreach efforts. In 2016 we restructured our staffing to include a 1.0 FTE PhD astronomer (hired in late 2016). This position provides major user communications, public/media communications, and outreach support from the Gemini South offices. Additionally an entry-level, limited-term position (3 years) was added in

early 2016 to support succession in local outreach and social media content production. Finally, the second year of the PIO internship program is underway at Gemini North.

6.5 PIO Outreach Statistics

| Activity/Event | Attendance |
|---------------------------------------|---------------|
| Journey Through the Universe, Hawai'i | 9,061 |
| Viaje al Universo, Chile | 6,062 |
| StarLab (Hawai'i and Chile) | 3,998 |
| Family Astro, AstroDay etc. | 2,869 |
| Live from Gemini (Hawai'i) | 45 |
| Summit Tours (Hawai'i and Chile) | 789 |
| Total | 22,824 |

Table 6-1: Participation in outreach activities

6.6 Broadening Participation and Workforce Development

Gemini continues to invest in its commitment to Workforce and Diversity. A strategic objective has been drafted for the period 2017-21 – *To be a model organization, inspiring others, in matters of diversity in the STEM workforce.*

Toward this objective, we introduced several new initiatives and actions during 2016:

1. "Diversity and Inclusion" is now a Performance Standard in the annual performance reviews to hold individuals and managers accountable and to promote diversity in the Working Culture.
2. We developed a Workplace Culture and Conduct training program and rolled out part one, focusing on AURA policy and sexual harassment, this year. The second part, focusing on AURA values, diversity and bullying behavior, will roll out in early 2017.
3. Furthermore, AURA continues to monitor developments in the area of unconscious bias (especially during the hiring process) and to educate its management, workforce, and governance. Briefings on the role of unconscious bias to hiring committees and hiring managers are a well-established part of the recruiting process and form the basis for fair treatment of applicant pools.
4. AURA had a booth presence at the October 2016 Society of Women Engineers National Conference held in Philadelphia, a significant presence from Gemini contributed to the ~11,000 attendees that included over 20 female engineers from AURA.

Gemini continued to develop workforce pipeline initiatives during 2016. The programs and internships are highlighted in Table 6-2.

| Department/Specialty | Funding | Educational Skill Level | Intern Program | 2016 Interns |
|-----------------------------------|-------------|--------------------------|----------------------|--------------|
| Engineering / Technical / Science | Akamai | Undergraduate | Akamai | 3 |
| Engineering / Technical | Gemini | Undergraduate / Graduate | Univ. of Victoria | 2 |
| Science | AGUSS | Undergraduate | AGUSS | 5 |
| Science/Engineering | Gemini | Various | Internal | 8 |
| Engineering | INSPIRE | Undergraduate | INSPIRE | 1 |
| Engineering | AURA/Gemini | Undergraduate / Graduate | Chilean Universities | 1 |

Table 6-2: Gemini internships in 2016

7 Finance and Organization

7.1 Finance

The Finance team works with budget managers and directors to develop and analyze budgets, staffing, and compensation plans. The Finance team process all budget input and monthly expenditures. In 2016 the team used the “scratchpad” budget model to carry detailed justifications of all account budgets in a single repository. Gemini Finance implemented a new interface process to transfer payroll and HR data more efficiently and securely, and developed controls to verify that accounting, payroll, and procurement data are transferred accurately between Gemini and CAS business systems. Additionally, in 2016 the Finance and AURA-CAS teams worked actively with NSF-DACS and NSF-CAAR in the negotiation of the 2017-2022 budget for the Gemini renewal Cooperative Agreement (CA) and Cooperative Support Agreement (CSA).

7.2 Organization and Staffing

Gemini currently has 168 employees on a Full Time Equivalent (FTE) basis; 51% of the staff members are based in Hilo and 49% are based in La Serena. Table 7-1 shows the distribution by Division and Core or Extra staff category. Annual staff turnover is currently about 7%. Core positions are the minimum needed to run the Observatory on a daily basis and to enable development capabilities (projects in operations or infrastructure improvements; as well as AO, instrumentation, telescope development, etc.). Extra positions support tactical plans (temporary need for more skills in an area, anticipation of retirement/succession planning, and others). These are enabled by voluntary turnover rate (typically 5-10%) and hiring lag (typically 3-6 months), and do not increase the total Observatory budget beyond the Board-authorized 3% margin.

| FTEs by Division | Core | Extra | Total |
|------------------|------|-------|-------|
| Administration | 16 | 1 | 17 |
| Development | 10 | 8 | 18 |
| Operations | 106 | 7 | 113 |
| Deputy Director | 12 | 1 | 13 |
| Directorate | 7 | | 7 |
| Total | 151 | 17 | 168 |

Table 7-1: 2016 staffing

7.3 Budget

7.3.1 Partner Contributions

The schedule of contributions for the Gemini Observatory budget is governed by a set of Administrative Guidelines; *i.e.* agreements made among the members of the Gemini International Partnership including NSF. These agreements clarify the members' Partnership shares and the timing for the payment of contributions. Table 7-2 shows the distribution of the Gemini cost shares under the current Partnership.

| Partner | Cost Share |
|---|------------|
| United States | 66.46% |
| Canada | 18.01% |
| Australia is I-t after 2015 * | 0.24% |
| Argentina | 3.08% |
| Brazil | 6.47% |
| Korea I-t in 2016, partner after 2018 * | 5.74% |
| Total | 100.00% |

* I-t: limited term

Table 7-2: Partners' cost shares

Table 7-3 sets forth 2016 actual contributions by Partner for Operations and Maintenance (O&M), Instrument Development Fund (IDF), and the combination of the O&M and IDF funds, including contributions from limited-term participants. Throughout, values are given in US\$1000, except where noted.

| Contributions | O&M | IDF | Total |
|---------------|---------------|--------------|---------------|
| US | 18,022 | 1,802 | 19,824 |
| Canada | 5,009 | 500 | 5,509 |
| Australia | 0 | 467 | 467 |
| Argentina | 856 | 86 | 942 |
| Brazil | 1,798 | 179 | 1,978 |
| Korea* | | 1,350 | 1,350 |
| Total | 25,684 | 4,385 | 30,070 |

*Limited-term arrangement in 2016.

Table 7-3: 2016 actual contributions by Partner

We use O&M funds to support the day-to-day activities involved in operating the telescopes and facilities. Broadly speaking, these activities are science support, engineering, instrumentation support, administration (including operations costs for base facilities, fleet and mountain infrastructure), software, information systems, research, public information, safety and the directorate. The IDF is dedicated to renewal and improvement of instrumentation and telescope facilities, primarily executed as contracts to teams within the Partnership and spent over multiple years.

7.3.2 Financial Results for the Year of 2016

Table 7-4 gives the summary of estimated O&M actual expenditures by expense category. Values through the end of 2016Q3 are actuals. Table 7-5 shows the summary of estimated IDF actual expenditures by Instrumentation Development project as of Q3.

| in \$ US | 2016 Estimated Expenses | 2016 Budget | Var 2016 Budget vs. Forecast \$ | Var 2016 Budget vs. Forecast % |
|-------------------------------------|-------------------------------|-------------|---------------------------------------|--------------------------------------|
| Total Labor | 17,334,586 | 17,739,633 | (405,047) | -2% |
| Supplies Equipment | 1,291,143 | 1,089,041 | 202,102 | 19% |
| Travel | 909,288 | 952,791 | (43,503) | -5% |
| Recruiting Relocation | 58,000 | 58,000 | - | 0% |
| Professional Fees | 1,535,597 | 1,531,005 | 4,592 | 0% |
| Meeting, Conf., Prof Dev. | 271,955 | 311,833 | (39,878) | -13% |
| Computer SW. and Equip | 607,238 | 539,854 | 67,384 | 12% |
| Total Site Costs | 2,418,719 | 2,496,498 | (77,779) | -3% |
| Spares | 264,750 | 370,810 | (106,060) | -29% |
| Other | 34,808 | 502,036 | (467,228) | -93% |
| Indirect Costs | 634,559 | 986,169 | (351,610) | -36% |
| Subcontracts | 1,045,291 | 614,829 | 430,462 | 70% |
| Total Non-Cap. Exp. | 26,405,934 | 27,192,500 | (786,566) | -3% |
| Capital - General | 1,394,066 | 107,500 | 1,286,566 | 1197% |
| Total Cap. Exp. | 1,394,066 | 107,500 | 1,286,566 | 1197% |
| Total Non-Labor Expenditures | 10,465,414 | 9,560,367 | 905,047 | 9% |
| TOTAL EXPENDITURES | 27,800,000 | 27,300,000 | 500,000 | 2% |

Table 7-4: O&M 2016 Actual vs Budget Expenditures; actual through Q3-2016

| In \$k | 2016 Budget | 2016 Estimated Expenses | 2016 Actuals as of 09/30/16 | 2016 Remaining Balance (\$) | 2016 Remaining Balance % |
|---------------------------|----------------|-------------------------------|--------------------------------------|-----------------------------------|--------------------------------|
| Program Support | \$80 | \$17 | \$16 | \$1 | 4% |
| GEN 4#3 | \$195 | \$62 | \$42 | \$20 | 32% |
| Instrument Upgrades + A&G | \$870 | \$105 | \$77 | \$29 | 27% |
| GHOST | \$3,114 | \$2,167 | \$1,701 | \$466 | 22% |
| AO Upgrades | \$121 | \$26 | \$11 | \$15 | 58% |
| LGSF upgrades | \$879 | \$836 | \$330 | \$506 | 0% |
| GMOS CCDs | \$42 | \$20 | \$18 | \$2 | 10% |
| GeMS-DM0 | \$461 | \$0 | \$0 | \$0 | 0% |
| GPI – HIA | \$273 | \$273 | \$273 | \$0 | 0% |
| Total | \$6,035 | \$3,506 | \$2,468 | \$1,039 | 30% |

Table 7-5: IDF 2016 Actual vs Budget Expenditures; actual through Q3-2016

7.3.3 2016 Budget vs Actual Variance Analysis

Total O&M 2016 estimated expenses and commitments are \$500k or 2% *greater* than the 2016 spending authority approved in the Board resolutions, and this total exceeds Partners' O&M contributions by \$2,115k. As part of the Transition Program (TP), this cash deficit is balanced from carry-forward that has accumulated during past years.

Figure 7-1 presents the 2016 budget versus actuals variance analysis (labor and non-labor) by variance driver. Overall, the key factors driving O&M spending changes outside the TP in 2016 are: (1) unplanned devaluation of Chilean Peso (CLP) of 8.1% (from CLP 620 to CLP 670 per USD 2016 annual average exchange rate), which leads to a \$381k effect in labor); (2) reduction of 2016 labor costs due to unplanned FTEs reduction or unplanned leavers (net decreasing labor costs of \$1,205k); (3) increasing labor for hiring

of Extra FTEs of \$907k; (4) increasing labor for marginal merit increases of \$ 274; increasing non-labor for additional computer equipment of \$73k; (5) budget underrun due to reduced electricity and indirect costs of \$632k; (6) increasing facilities and travel expenses of \$87k; (7) increasing O&M non-labor costs due O&M and TP projects of \$900k; and (8) increasing O&M expenses for the GN Toptica Laser of \$977k.

With net savings in 2016 and the opportunity to implement additional Transition Program and O&M projects for long-term savings, sustainability, and infrastructure improvement, we included a net 2016 non-labor budget increase of \$1,877k. O&M and TP projects of \$900k include GN shutter repairs, new solar energy systems at GS summit facilities and base facilities office expansion. The remaining budget addition of \$977k O&M expenses are due to the GN Toptica Laser.

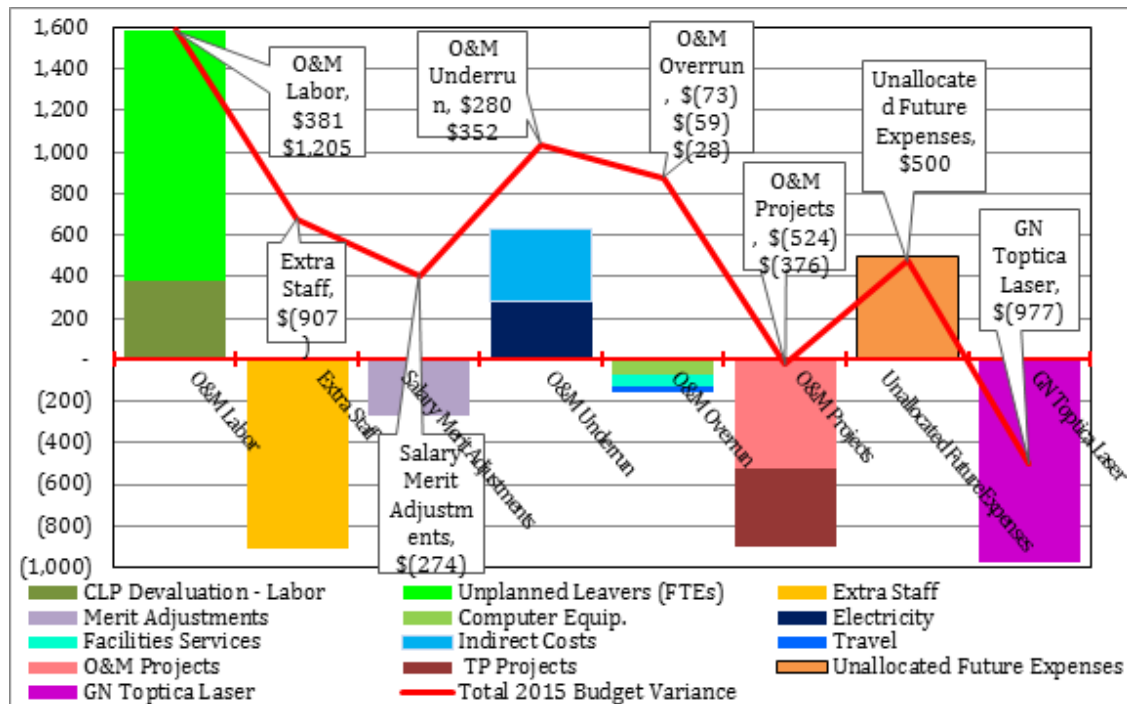


Figure 7-1: 2015 budget vs actuals variance analysis (labor and non-labor)

For IDF the 2016 expenses are \$2,468K below the 2016 approved budget. The summary of IDF 2016 budget vs actual expenditures is shown in Table 7-6.

| In \$k | 2016 Budget | 2016 Estimated Expenses | 2016 Actuals as of 09/30/16 | 2016 Remaining Balance |
|---------------------------|----------------|-------------------------------|-----------------------------------|------------------------------|
| Program Support | \$80 | \$17 | \$16 | \$1 |
| GEN 4#3 | \$195 | \$62 | \$42 | \$20 |
| Instrument Upgrades + A&G | \$870 | \$105 | \$77 | \$29 |
| GHOST | \$3,114 | \$2,167 | \$1,701 | \$466 |
| AO Upgrades | \$121 | \$26 | \$11 | \$15 |
| LGSF upgrades | \$879 | \$836 | \$330 | \$506 |
| GMOS CCDs | \$42 | \$20 | \$18 | \$2 |
| GeMS-DM0 | \$461 | \$0 | \$0 | \$0 |
| GPI – HIA | \$273 | \$273 | \$273 | \$0 |
| Total | \$6,035 | \$3,506 | \$2,468 | \$1,039 |

Table 7-6. IDF 2016 budget vs actual expenditures as of Q4-2016

Table 7-7 presents the drivers of 2016 IDF budget over / under spending.

| IDF PROJECT | OVER / UNDER SPENDING |
|---------------------------|---|
| GEN 4#3 | Gen 4#3 Contract(s) to start in Q1 2017, hence no funds will be spent on contracts in 2016 |
| Instrument Upgrades + A&G | RFP for the 2016 Instrument Upgrade Program released in Q3. IDF Upgrade Contracts to start in Q2-2017. A&G was low priority in 2016 |
| GHOST | Q4 budgeted \$640k payment rescheduled from 2016 to 2017 and rebaseline of GHOST “Contingency Reserve” of \$384k |
| AO Upgrades | NGS2 Factory Acceptance Testing in Q4. ALTAIR RTC upgrade design contract (~\$66k) remains on hold |
| LGSF upgrades | New GS Toptica Laser on track |
| GMOS CCDs | GMOS-S work completed. The GN GMOS upgrade team resolved the GMOS-N anomalies found, GN upgrade rolled over to 2017 |
| GeMS-DM0 | Milestone payments to be paid near completion of the project, which may slip into 2017 |
| GPI - HIA | GPI project complete and closed |

Table 7-7 2016 IDF budget over / under spending analysis

8 Program Plan for 2017

8.1 Overview of 2017

Gemini’s highest priority remains to operate the Observatory to enable our users’ scientific advancement. We continue to develop new capabilities and operate Gemini’s telescopes on a lower budget against the backdrop of regular operations. All milestone dates throughout this section are 2017 unless otherwise noted, and budget figures in this section are in US dollars.

8.1.1 Observatory Budget 2016–2018 Overview

We will continue to manage the Observatory budget in two categories: 1. Operations and Maintenance (O&M – spent on an annual basis); and 2. Instrument Development Fund (IDF - used for long-term, multi-year instrumentation and adaptive optics projects). Table

8-1 presents the 2016 forecast, the approved 2017 budget, and plans for 2018, in the above categories. Note that the IDF budget in a given year is not exactly aligned with the income in these categories given the multi-year nature of the financed projects.

| In \$ US | 2016 Actual ¹ | 2017 Budget | 2018 Budget | 2019 Budget |
|-----------------------------------|--------------------------|-------------------|-------------------|-------------------|
| Operations and Maintenance (O&M) | 27,800,000 | 28,509,441 | 29,585,649 | 30,457,838 |
| Instrument Development Fund (IDF) | 3,506,167 | 4,671,160 | 8,133,523 | 5,435,952 |
| Total | 31,306,167 | 33,180,601 | 37,719,172 | 35,893,790 |

¹Note: 2016 actual through Q3; estimated for Q4

Table 8-1: Budget overview, by fund, 2016–2019

8.1.2 Economic Assumptions through 2018

Table 8-2 sets forth budget assumptions; these economic ratios are based on updated inflation and foreign exchange (FX) rate forecasts from the International Monetary Fund. We include these economic assumptions, including their effect on labor costs, in the planned budgets.

| Year | 2017 | 2018 | 2019 |
|-----------------|-------|-------|-------|
| US inflation | 2% | 2% | 2% |
| Chile inflation | 4.60% | 4.60% | 4.60% |
| FX rate CLP/USD | 640 | 640 | 640 |

Table 8-2 Budget assumptions 2017–2019.

8.1.3 Staffing Plan

Table 8-3 shows the staffing plans for 2016–2018, as FTEs integrated over each calendar year. Most of the planned effort remains within defined Divisions: Operations staff support operations activities and Development staff support instrument development projects. Engineering effort contributes to Development projects at the ~10% level. The staffing plan indicates the number of Core and Extra FTEs for 2016 and 2017-2018. The FTEs' tags "CORE" and "EXTRA" that we introduced in the 2017 staffing plan are defined in section 7.2 above.

| | Core Staff | | | Extra Staff | | | Total Staff | | |
|-----------------------|------------|-----------|-----------|-------------|-----------|-----------|-------------|-----------|-----------|
| | 2016 Plan | 2017 Plan | 2018 Plan | 2016 Plan | 2017 Plan | 2018 Plan | 2016 Plan | 2017 Plan | 2018 Plan |
| Administration | 15.5 | 16.0 | 16.0 | 1.0 | 1.0 | - | 16.5 | 17.0 | 17.0 |
| Accounting | 2.0 | 2.0 | 2.0 | | | | 2.0 | 2.0 | 2.0 |
| Facilities & Admin. | 12.0 | 12.0 | 12.0 | | | | 12.0 | 12.0 | 12.0 |
| Safety | 1.5 | 2.0 | 2.0 | 1.0 | 1.0 | | 2.5 | 3.0 | 3.0 |
| Development | 10.3 | 12.0 | 12.0 | 7.8 | 5.1 | 5.0 | 18.1 | 17.1 | 17.0 |
| Adaptive Optics | 4.0 | 5.0 | 5.0 | 3.1 | 2.8 | 3.0 | 7.1 | 7.8 | 8.0 |
| Sys. Engineering | 3.3 | 4.0 | 4.0 | 1.5 | 1.3 | 1.0 | 4.8 | 5.3 | 5.0 |
| Instrumentation | 3.0 | 3.0 | 3.0 | 3.2 | 1.0 | 1.0 | 6.2 | 4.0 | 4.0 |
| Operations | 106.5 | 110.9 | 111.1 | 6.6 | 7.8 | 8.0 | 113.1 | 118.7 | 119.1 |
| Engineering - S | 8.9 | 10.0 | 10.0 | 1.0 | 0.4 | - | 9.9 | 10.4 | 10.0 |
| Engineering - N | 9.0 | 9.0 | 9.0 | 1.0 | 1.0 | 1.0 | 10.0 | 10.0 | 10.0 |
| ISG | 12.2 | 12.0 | 12.0 | | | | 12.2 | 12.0 | 12.0 |
| SOS - S | 10.0 | 10.0 | 10.0 | | | | 10.0 | 10.0 | 10.0 |
| SOS - N | 10.0 | 10.0 | 10.0 | | | | 10.0 | 10.0 | 10.0 |

| | | | | | | | | | |
|--------------------|--------------|--------------|--------------|-------------|-------------|-------------|--------------|--------------|--------------|
| Astronomers -S | 11.3 | 12.2 | 12.1 | 0.7 | 1.0 | 1.0 | 12.0 | 13.2 | 13.1 |
| Astronomers-N | 9.4 | 10.8 | 11.0 | 0.7 | 1.0 | 1.0 | 10.1 | 11.8 | 12.0 |
| Software | 12.9 | 14.0 | 14.0 | 1.3 | 4.0 | 5.0 | 14.2 | 18.0 | 19.0 |
| Spotters | - | - | - | 1.0 | - | - | 1.0 | - | - |
| Summit Crew - S | 11.8 | 12.0 | 12.0 | 1.0 | 0.3 | - | 12.8 | 12.3 | 12.0 |
| Summit Crew - N | 11.0 | 11.0 | 11.0 | | | | 11.0 | 11.0 | 11.0 |
| Deputy Dir | 11.8 | 12.5 | 12.5 | 1.1 | 1.8 | 2.0 | 12.9 | 14.3 | 14.5 |
| PIO | 5.3 | 6.5 | 6.5 | 0.8 | 1.0 | 1.0 | 6.1 | 7.5 | 7.5 |
| SCI User Support | 5.9 | 6.0 | 6.0 | 0.3 | 0.8 | 1.0 | 6.2 | 6.8 | 7.0 |
| Directorate | 7.0 | 7.0 | 7.0 | | | | 7.0 | 7.0 | 7.0 |
| GRANDTOTAL | 151.1 | 158.4 | 158.6 | 16.5 | 15.6 | 15.0 | 167.6 | 174.0 | 173.6 |

Table 8-3: 2017 Staffing Plan – Core and Extra FTEs

8.1.4 Operations & Maintenance (O&M) Budget 2016 to 2018

The Operations and Maintenance (O&M) budget supports the day-to-day activities involved in operating the telescopes and facilities. Table 8-4 shows the O&M budget by cost categories for the period 2017 - 2018.

8.1.5 Instrument Development Fund (IDF) Budget 2014-2017

The IDF is used for instrumentation and adaptive optics systems. Table 8-5 presents the IDF 2017-2018 spending plan by instrument project, defined in §5 above. The additional account line for program support contains the budget for contracted work to support future instrument upgrades.

| In \$ US | 2015 Actual | 2016 Forecast | 2017 Budget | 2018 Budget |
|---|----------------|------------------|----------------|----------------|
| Salaries | 12,652,582 | 12,530,002 | 13,148,873 | 13,301,512 |
| Benefits | 4,450,309 | 4,348,235 | 4,428,775 | 4,562,708 |
| Temporary | 370,214 | 593,386 | 608,468 | - |
| EXTRA Staff | | | 458,213 | 630,954 |
| Finiquitos accrued liability Amortization | | | 366,667 | 366,667 |
| Accrued Finiquitos | | | 146,000 | 146,000 |
| Total Labor | 17,473,105 | 17,471,624 | 19,156,996 | 19,007,841 |
| Supplies Equipment | 1,165,824 | 1,183,345 | 1,235,064 | 1,155,364 |
| Travel | 975,550 | 967,582 | 952,032 | 1,010,817 |
| Recruiting Relocation | 238,716 | 58,000 | 69,000 | 60,010 |
| Professional Fees | 1,560,846 | 1,546,005 | 1,639,367 | 1,493,408 |
| Meeting, Conferences, Prof Dev. | 328,970 | 309,053 | 347,528 | 330,823 |
| Computer Software and Equipment | 523,898 | 602,708 | 664,096 | 572,731 |
| Facilities | 653,946 | 638,836 | 615,775 | 688,666 |
| Maintenance | 150,673 | 182,069 | 174,408 | 193,157 |
| Utilities | 1,262,051 | 1,298,802 | 1,064,652 | 1,377,687 |
| Meals and Lodging | 525,545 | 391,694 | 277,148 | 389,025 |
| Total Site Costs | 2,592,215 | 2,511,400 | 2,131,983 | 2,648,535 |
| Spares | 412,714 | 247,539 | 138,300 | 393,392 |
| Other | 70,165 | 280,628 | 71,470 | 556,106 |
| Indirect Costs | 722,587 | 836,169 | 672,766 | 828,025 |
| Subcontracts | 1,514,747 | 1,063,151 | 543,660 | 652,272 |
| Total Non-Capital Expenditures | 27,579,337 | 27,077,204 | 27,622,262 | 28,709,323 |
| Capital Computer | 475,371 | - | 115,600 | 379,946 |

| | | | | |
|----------------------------|-------------------|-------------------|-------------------|-------------------|
| Capital General | 438,878 | 222,796 | 771,579 | 496,380 |
| Total Capital Expenditures | 914,250 | 222,796 | 887,179 | 876,326 |
| TOTAL EXPENDITURES | 28,493,586 | 27,300,000 | 28,509,441 | 29,585,649 |

Table 8-4: O&M budget by cost categories for the period 2017–2018, values in USD

| In \$ US | Budget 2017 | Budget 2018 | Budget 2019 |
|---|--------------------|--------------------|--------------------|
| Instr. Program. Support and Maintenance | 69,610 | 71,490 | 73,420 |
| GHOST | 810,000 | 2,211,000 | 549,000 |
| GHOST - Internal | 80,000 | 82,160 | 84,378 |
| Instrument Upgrades | 250,500 | 666,917 | 324,401 |
| AO upgrades (Altair, Canopus) | 246,000 | 131,432 | 0 |
| A&G System Development | 752,000 | 648,932 | 120,375 |
| LGSF upgrades | 895,750 | 0 | 0 |
| GMOS CCD Replacement | 22,900 | 0 | 0 |
| Gen4#3 | 1,000,000 | 4,200,000 | 4,200,000 |
| Gen4#3 - Internal | 80,000 | 82,160 | 84,378 |
| GEN5#1 | 0 | 0 | 0 |
| GEN5#1 Internal | 0 | 0 | 0 |
| GeMS-DM0 | 390,900 | 39,432 | 0 |
| GPI Relocation Study | 73,500 | 0 | 0 |
| GPI - HIA, Opto-Mech Super Structure | 0 | 0 | 0 |
| TOTAL IDF | 4,671,160 | 8,133,523 | 5,435,952 |

Table 8-5: IDF spending for the years 2016 – 2018 by instrumentation project

8.2 Science and Engineering Operations in 2017

8.2.1 Regular Operations

Regular day-to-day and night-to-night Operations is the Observatory's top priority. It includes the following items.

- Maintain the instruments and telescopes in working order consistent with the requested science time on sky; monitor performance and take remedial action as needed.
- Run the International Time Allocation Committee (ITAC) process to combine the national TAC results into an executable queue and visitor program consistent with available time, conditions, and instrumentation.
- Support the user community (in conjunction with the NGOs) in preparing their observations for the telescope.
- Provide web-based documentation suitable for PI reference on instrumentation, software and Observatory processes.
- Execute queue observing programs on behalf of the community as required; currently this equates to more than 80% of the observing.
- Support visiting observers in their execution of their own and others' programs on the telescopes.
- Ensure integrity of data (headers & quality control information) entering the Gemini Observatory Archive.
- Support visiting instruments as needed and as possible.
- Propose and execute continual improvements in instrumentation, telescope, and enclosure to maintain performance levels.
- Propose and execute continual improvements in operations software on behalf of the community and for internal usability, to maintain performance levels.

- Propose and execute continual improvements in operations processes on behalf of the user community, with guidance and input from the appropriate committees.
- Provide expertise and input to the Development Division in carrying out major enhancements of instrumentation.
- To ensure economical operations and a consistent interface with the user community, maintain approximate symmetry between the processes, equipment, and staffing at the two Gemini sites.
- Staff the “third” and final level of a helpdesk to respond to queries from the user community. (The first two levels are (i) NGOs and (ii) instrument specialists at the NGOs.)

A regular system of preventive and corrective maintenance supports the first requirement of maintaining the instruments, telescopes, and enclosures in working order. Preventive maintenance is carried out at the summits on a regular schedule (daily, weekly, etc. depending on the system) using a system of work orders. More major items requiring additional staffing are handled by planning on a variety of timescales. A central list of major maintenance work is held by the Heads of Engineering Operations and is discussed weekly among engineering managers and at the quarterly planning meetings. Items are tracked according to progress or completion, and new needs are evaluated before adding to the list and determining a possible schedule.

8.2.2 Science and Engineering Operations Core Projects and Goals

In addition to this regular operations support, in 2017 we will undertake a number of projects to improve longer-term sustainability impact or to improve service to users.

- Implement first round of work on Observatory-wide Obsolescence Mitigation Plan
 - The obsolescence plan spans five years. Work will continue throughout 2017.
- Real-time Software Upgrades
 - All of the agreed elements of this project should be complete by Q4. Possible threats to schedule include GMOS-N CCDs (if significant issues surface during commissioning) and NGS2 (for which the scope of the software effort is not yet known).
- Replace the Gemini North dome shutter drive chains and gearboxes
 - To be complete by Q3. The Gemini Board has granted engineering shutdown time to carry out this replacement, spanning the end of Semester 2017A and beginning of Semester 2017B.
- Science Operations Model Upgrade
 - Work has commenced on a subset of high-priority options emerging from the Science Operations Review. Some will be complete within 2016; others will extend into 2017, including the provision of an archived method of communication on all accepted science programs (Q3) and advance work on a more automated queue planning system to streamline the Queue Coordinating process (planning and prototyping may extend throughout 2017, as software effort to implement it will not be available initially).
- Upgrade the Cerro Pachón Network Link
 - We expect to be using the LSST’s fiber optic connection early in 2017. At present the timeline calls for the fiber to be in place by the end of 2016 so we hope to be using it by the end of Q1 2017. We are not in control of the schedule on this work.

8.3 Science User Support in 2017

The Science User Support Department (SUSD) advocates for the users and enables Gemini Principal Investigators to produce world-class scientific results in a timely manner. The SUSD leads post-observing user support. Its staff maintain a data reduction package for the user community that enables astronomers to remove instrumental signatures from data obtained using Gemini facility instruments, and they support users in its use. The SUSD maintains communications between the Observatory and the National Gemini Offices and ensures that NGO staff members receive appropriate training. Specific goals for 2017 follow.

- Complete implementation of a coherent system of user communication throughout the “science lifecycle”, including regular support for the Data Reduction User Forum (DR Forum), post-observing contact with PIs, and improvements to the regular support mechanism that the Helpdesk currently provides.
 - Evaluate DR Forum traffic in the context of the “user contributor” role (Q1)
 - Begin collecting and tracking user satisfaction data and feedback (Q1)
 - Implement new external Helpdesk (Q2)
 - Implement post-observing contact with PIs and revise Contact Scientist duties and responsibilities (Q4)
- Improve data reduction software documentation and cookbooks.
 - GNIRS cookbook (Q2)
 - FLAMINGOS-2 cookbook (Q2)
- Release Python science quality data reduction software for facility imaging modes (GMOS, F2, NIRI, and GSAOI) for users (Q3).
- Update Gemini IRAF for new GMOS-N Hamamatsu CCDs (Q2) and F2 MOS mode (Q1). These will be the last releases of Gemini IRAF.

8.4 Instrumentation and Facility Development in 2017

The 2017 plans for instrumentation and facility development build on several substantial on-going projects. We list principal objectives for the year in Table 8-6. Table 8-7 outlines our goals for lower priority projects we will execute on a best-efforts basis. We expect to make substantial progress on many of these projects while understanding that we may postpone or alter any and/or all of them to accommodate increased needs for any of the higher priority projects.

| Project | 2017 Plan |
|---------------------|---|
| GHOST | Complete Build Phase and Start Test Phase (Q4) |
| Gen4#3 | Start Conceptual Design Stage (aggressively, Q2) |
| Toptica Lasers | Commission GeMS laser (Q3); Receive Altair laser (Q4) |
| GMOS CCDs | Install and commission new CCDs into GMOS-N (Q1) |
| Instrument Upgrades | Complete 2015 project (Q2); Start 2016 project(s) (Q2); Issue 2017 RfP (Q4) |
| NGS2 | Tentatively receive, test, and install (TBD) |

Table 8-6: High-priority projects for Development and their planned activities in 2017

| Project | 2017 Goals |
|------------------------|---|
| A&G | Demonstrate new mechanism controller functionality with all required mechanisms Test and characterize new wavefront detector system for higher sensitivity |
| Altair | Start contracted work to replace the real-time computer to allow for higher performing algorithm implementation Commission the new Altair dichroics |
| DM0 | Receive DM0 and electronics to serve as spare for existing DMs. |
| IR Detector Controller | Demonstrate ARC controller with Multiplexer Develop requirements for software needed to support new controller |

Table 8-7: Goals for non-priority Development projects in 2017

8.5 Administration & Facilities and Safety in 2017

8.5.1 Finance and Administration

Administration provides cost-effective administrative support and delivers timely and accurate information to management and governance. The Administration and Facilities Group (AFG) supports in the areas of facilities, infrastructure, fleet, administration, visitor services, and travel services to staff and visitors. The Finance team delivers financial guidance and information to management to support executing budgets within the Board authorized limits. Specific goals for 2017 follow, to be completed by the end of the year.

- Budgetary responsibility
 - Finish 2017 within [-2%; +3%] of the requested O&M Budget (\$28.0M). Continue to promote the accountability among the budget account managers.
- Business Services
 - Assist Gemini employees in receiving training and developing any knowledge required to operate uniformed business systems that will serve Gemini and other NSF-funded programs.
- Business Systems Review (BSR)
 - Gather and prepare the documentation required for the anticipated 2017 BSR, respond to NSF inquiries on the 2017 BSR and if needed, implement changes or actions in response to the results of the BSR process that NSF will complete in 2017.
- Facilities Services
 - Continue modernizing HBF & SBF facilities. Review and adjust AFG's operations standards to achieve 2017 cost efficiency and energy sustainability goals. Provide superior workplaces for Gemini's employees in 2017.

8.5.2 Safety

The Safety program assures a safe and healthy environment for employees and visitors. Gemini's working culture explicitly emphasizes safety of people and equipment. The Safety group will fully integrate safety into the Observatory operational activities, and will work jointly with partner telescopes on Cerro Pachón and Maunakea to establish shared Safety programs that combine best practices and resources. Specific goals to complete by the end of 2017 follow.

- Staff Safety
 - Continue providing and assuring a safe and healthy environment for employees and visitors. Provide modern Safety tools and systems and achieve a 90% completion of the mandatory Level A and 80% of the mandatory Level B safety training.
- Safety Operations
 - Produce a consolidated Gemini Safety group that delivers Safety services consistent with those of co-located NSF-funded centers in Chile. Continue establishing synergies and sharing Safety resources with other Maunakea telescopes.
- Managers' Safety responsibilities
 - Assist Gemini Managers to perform the Managers' Safety tasks required in OSHA's standards and Gemini's Safety program. Suggest actions that help to mitigate Safety hazards identified in the regular Safety walkthroughs.

8.6 Public Information and Outreach in 2017

We will continue our regular local outreach programming, including: Journey Through the Universe (JTTU), Viaje al Universo (Viaje), AstroDay, StarLab, media relations, publications, electronic and social media, user communications support, PIO internship, library services, and science result tracking. We also identify the following specific expanded activities to complete in 2017.

- Integrate, as a pilot program at Gemini North, the utilization of an outside contracted evaluation firm to assess student and teacher understanding of Observatory STEM career opportunities. (Q3)
- Expand Observatory STEM career resources outreach programming to include a minimum of 3 new events/activities at each site over the full year, to increase awareness and understanding of Observatory career opportunities available in host communities (Hilo and La Serena). (Q4)
- Extend JTTU/Viaje beyond concentrated "Journey Week" into multiple (minimum of 6) year-long activities, interactions, and programming in local host communities. (Q4)
- Expand capacity of new Gemini-South PIO hire to user communications support and media relations.
 - Implement formal training and work-plan for new position (Q1)
 - Independent completion of at least 2 press releases and 2 web features (Q3).
 - Expand (and update) international media contact list to 50 relevant journalists (Q4).
- Integrate new press release template within Gemini website (Q1)
- Complete audience segregation of social media content into separate (distinct) accounts (Q2)

8.7 Gemini External Relations in 2017

- Directorate attendance at Partner national astronomy meetings
- Develop Strategic Plan to implement Gemini Board's Strategic Vision
- Lead the development of memoranda of understanding with any new limited-term Partners

8.8 HR Plans in 2017

The following goals focus on increasing staff satisfaction and engagement, as measured by achieving a staff voluntary turnover rate below 6%.

- Implement changes to researcher roles, including changing time allocations for research, specialized roles for Fellow vs. Scientist vs. Astronomer, collaborative activity of Instrument Scientists, and support of Science User Support Department activities
- Issue completed Manager's Handbook
- Identify and implement actions following the Staff Engagement Survey

Appendix A. Acronyms and Abbreviations

| | |
|----------|--|
| A&G | Acquisition and Guiding units |
| AAO | Australian Astronomical Observatory |
| AFG | Administration and Facilities Group |
| Altair | Altitude Conjugated Adaptive Optics for Infrared |
| ANU | Australian National University |
| AO | Adaptive Optics |
| ApJ | Astrophysical Journal |
| ARC | Astronomical Research Cameras |
| AURA | Association of Universities for Research in Astronomy, Inc. |
| BFO | Base Facility Operations |
| CADC | Canadian Archive and Data Centre |
| CAS | (AURA) Central Administrative Services |
| CCD | Charge-Coupled Device |
| CP | Cerro Pachón |
| DR | Data Reduction |
| DSSI | Differential Speckle Survey Instrument |
| ESPaDOnS | Echelle Spectro-Polarimetric Device for the Observation of Stars |
| FDF | Facilities Development Fund |
| FITS | Flexible Image Transport System |
| FTE | Full-Time Equivalent |
| FTZ | Foreign Trade Zone |
| GeMS | Gemini Multi-conjugate Adaptive Optics System |
| Gen4#3 | Gemini's next facility class instrument (Generation 4, #3) |
| GHOST | Gemini High-resolution Optical SpecTrograph |
| GIFS | Gemini Instrument Feasibility Studies (for Gen4#3) |
| GMOS | Gemini Multi-Object Spectrograph |
| GMOS-N | Gemini Multi-Object Spectrograph-North |
| GMOS-S | Gemini Multi-Object Spectrograph-South |
| GN | Gemini North |
| GNIRS | Gemini Near-Infrared Spectrograph |
| GOA | Gemini Observatory Archive |
| GPI | Gemini Planet Imager |
| GRACES | Gemini Remote Access to Canada-France-Hawaii ESPaDOnS Spectrograph |
| GS | Gemini South (or Gemini Staff, only in time allocation listing) |
| GSA | Gemini Science Archive |
| GSAOI | Gemini South Adaptive Optics Imager |
| HBF | Hilo Base Facility |
| ICD | Interface Control Document |

| | |
|--------|--|
| IDF | Instrument Development Fund |
| IFU | Integral Field Unit |
| IMF | International Monetary Fund |
| IR | Infrared |
| ITAC | International Time Allocation Committee |
| IYoL | International Year of Light |
| KASI | Korea Astronomy and Space Science Institute |
| LCH | Laser Clearing House |
| LGSF | Laser Guide Star Facility |
| LLP | Large and Long Program |
| MCAO | Multi-Conjugate Adaptive Optics |
| MK | Mauna Kea |
| MNRAS | Monthly Notices of the Royal Astronomical Society |
| NGO | National Gemini Office |
| NGS2 | Natural Guide Star Wavefront Sensor upgrade project (for Gemini South) |
| NIFS | Near-Infrared Integral Field Spectrometer |
| NIR | Near-infrared |
| NIRI | Near Infrared Imager and Spectrometer |
| NOAO-S | National Optical Astronomy Observatory-South |
| NRC-H | National Research Council of Canada, Hertzberg Institute for Astronomy |
| NSF | National Science Foundation |
| NTAC | National Time Allocation Committee |
| O&M | Operations and Maintenance (budget fund) |
| OCS | Observatory Control Systems |
| OIWFS | On-Instrument Wavefront Sensor |
| PI | Principal Investigator |
| PIO | Public Information and Outreach |
| PNAS | Publications of the National Academy of Sciences |
| Q1 | Quarter 1 |
| QAP | Quality Assessment Pipeline |
| SOS | Science Operations Specialist |
| SPEC | Special grants and awards fund |
| STAC | Science and Technology Advisory Committee |
| STEM | Science, Technology, Engineering, and Mathematics |
| SUSD | Science User Support Department |
| TAC | Time Allocation Committee |
| TP | Transition Program |
| UCG | Users' Committee for Gemini |
| US | United States |
| z | Redshift |

Appendix B. Publications by Staff

B.1 Staff Refereed Publications

Levenson, N. A.[8]. The nuclear and extended mid-infrared emission of Seyfert galaxies. *Monthly Notices of the Royal Astronomical Society*, 463:3531-3555. 12/2016.

Levenson, N. A.[10]. The complex evolutionary paths of local infrared bright galaxies: a high-angular resolution mid-infrared view. *Monthly Notices of the Royal Astronomical Society*, 463:2405-2424. 12/2016.

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Miller, Bryan W.[12]. Faint dwarf galaxies in Hickson Compact Group 90*. *Monthly Notices of the Royal Astronomical Society*, 463:1284-1290. 12/2016.

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Margheim, Steve[5]. Observational Confirmation of a Link Between Common Envelope Binary Interaction and Planetary Nebula Shaping. *The Astrophysical Journal*, 832:125. 12/2016.

Trujillo, Chadwick.[2]. New Extreme Trans-Neptunian Objects: Toward a Super-Earth in the Outer Solar System. *The Astronomical Journal*, 152:221. 12/2016.

Schwamb, Megan[29]. OSSOS. IV. Discovery of a Dwarf Planet Candidate in the 9:2 Resonance with Neptune. *The Astronomical Journal*, 152:212. 12/2016.

Rantakyro, Fredrik[38]. Dynamical Mass Measurement of the Young Spectroscopic Binary V343 Normae AaAb Resolved With the Gemini Planet Imager. *The Astronomical Journal*, 152:175, 12/2016.

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Appendix C. Publications by Users

See notes for methodology.⁷⁸

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⁸ Gemini’s qualifying criterion is the same as that used by Hubble Space Telescope and European Southern Observatory. To qualify, papers based on their output, must employ in an original way an image, spectrum or data set produced by Gemini to derive new scientific results. No attempt is made to fractionate papers per telescope used in the case of papers based on the use of two or more other facilities. Hence, the same paper may be counted several times, for example by Gemini, Keck and Subaru, if it includes data from any of these telescopes.

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Appendix D. Science Programs 2015B and 2016A

D.1 2015B Science Programs – Gemini North

| Band | Gemini ID | PI Name | Partners | Instrument | Title | Time |
|------|---------------|------------|----------|---------------|--|------------|
| C | GN-2015B-C-1 | Tominaga | Subaru | GMOS-N | Spectroscopy of host galaxies of supernovae with shock breakout detection | 10.00 HR |
| C | GN-2015B-C-2 | Morokuma | Subaru | GMOS-N | High-z Low-Mass Active BHs Selected via Hour-Scale Variability | 10.00 HR |
| C | GN-2015B-C-3 | Griffith | US | GMOS-N | Elementary Abundances of Planetary Systems | 1.00 NIGHT |
| 1 | GN-2015B-LP-1 | Fraser | LP | GMOS-N, NIRI | COL-OSSOS: COLOURS for the Outer Solar System Object Survey (North) | 50.00 HR |
| 1 | GN-2015B-LP-2 | Kasliwal | LP | GMOS-N | Rapid Spectroscopy of Elusive Transients and Young Supernovae | 23.10 HR |
| 1 | GN-2015B-LP-3 | Huitson | LP | GMOS-N | The First Survey Dedicated to the Detection and Characterization of Clouds in Exoplanet Atmospheres | 6.30 HR |
| 1 | GN-2015B-LP-4 | Balogh | LP | GMOS-N | The GOGREEN Survey of dense galaxy environments at $1 < z < 1.5$ (North) | 19.50 HR |
| 2 | GN-2015B-LP-5 | Crossfield | LP | DSSI, GNIRS | Validating K2's Habitable and Rocky Planets with AO Imaging | 20.00 HR |
| 2 | GN-2015B-LP-6 | Huitson | LP | GMOS-N | The First Survey Dedicated to the Detection and Characterization of Clouds in Exoplanet Atmospheres | 17.50 HR |
| 2 | GN-2015B-LP-7 | Shen | LP | GNIRS | A GNIRS Near-IR Spectroscopic Survey of $z > 5.7$ Quasars | 35.30 HR |
| 1 | GN-2015B-Q-1 | A. Almeida | BR | GRACES | Radial velocity confirmation of an exoplanet candidate around an evolved compact binary | 4.80 HR |
| 1 | GN-2015B-Q-2 | Arias | AR | GNIRS | Disentangling the population of evolved massive stars in the galaxy M33 | 8.60 HR |
| 1 | GN-2015B-Q-3 | Bresolin | UH | GMOS-N | The extreme environments of thermonuclear supernovae | 6.00 HR |
| 1 | GN-2015B-Q-4 | Chambers | UH | GMOS-N, GNIRS | Pan-STARRS Survey for Counterparts to aLIGO GW events | 4.00 HR |
| 1 | GN-2015B-Q-5 | Chapman | CA | GNIRS | The origin of multiplicity in sub-mm sources: physical associations or line-of-sight projections? | 22.00 HR |
| 1 | GN-2015B-Q-6 | Dawson | US | GMOS-N | Two Ideal Dark Matter Colliders | 9.00 HR |
| 1 | GN-2015B-Q-7 | Diamond | US | GNIRS | Late-time near-infrared spectroscopy of SN 2014J | 2.10 HR |
| 1 | GN-2015B-Q-8 | Fesen | US | NIRI | Near-infrared identification of the young pulsar J0205+6449 | 1.50 HR |
| 1 | GN-2015B-Q-9 | Foley | US | GMOS-N | Understanding the Progenitor Systems, Explosion Mechanisms, and Cosmological Utility of Type Ia Supernovae (North) | 4.00 HR |
| 1 | GN-2015B-Q-10 | Foley | BR/US/AU | GMOS-N | DES Supernova Cosmology (North) | 21.60 HR |
| 1 | GN-2015B-Q-11 | Freeman | AU | GRACES | Were globular clusters born in dark matter sub-halos? | 5.40 HR |
| 1 | GN-2015B-Q-12 | Gladders | US | GMOS-N | Time delays for the sextuply-lensed quasar SDSS J2222+2745 from GMOS | 3.20 HR |
| 1 | GN-2015B-Q-13 | González | AR | GMOS-N | Exploring the faint galaxy content and the globular cluster systems of the Pegasus I Group. | 3.30 HR |
| 1 | GN-2015B-Q-14 | Im | KR | NIRI | NIR Imaging of A Quasar Candidate at $z \sim 7$ | 2.40 HR |
| 1 | GN-2015B-Q-15 | Lee | KR | GMOS-N | An Integral Field Unit Spectroscopy of Local E+A Galaxies | 14.10 HR |
| 1 | GN-2015B-Q-16 | Lee | KR | NIRI | Precise Mass Measurement of Circumstellar Shells around Luminous Blue Variables | 2.50 HR |

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| 1 | GN-2015B-Q-17 | Mackey | AU | GMOS-N | A Search for Intergalactic Globular Clusters in the Local Group | 2.60 HR |
| 1 | GN-2015B-Q-18 | Meech | UH | GMOS-N | Onset of 238P/Read Activity | 2.60 HR |
| 1 | GN-2015B-Q-19 | Meech | UH | GMOS-N | Testing Solar System Formation Models | 4.80 HR |
| 1 | GN-2015B-Q-20 | Montet | US | DSSI | Fundamental Parameters of Pre-Main Sequence M Dwarfs | 14.10 HR |
| 1 | GN-2015B-Q-21 | Moskovitz | US | GMOS-N, GNIRS | Mission Accessible Near-Earth Objects Survey (MANOS) (North) | 22.04 HR |
| 1 | GN-2015B-Q-22 | Reipurth | UH | GNIRS,NIRI | Properties of Newborn Triple and Quadruple Systems | 11.50 HR |
| 1 | GN-2015B-Q-23 | Rodney | US | GMOS-N | The Next Frontier: High-Redshift Supernovae in the HST Frontier Fields (North) | 5.00 HR |
| 1 | GN-2015B-Q-24 | Saha | US | GMOS-N | Establishing a Network of DA White Dwarf SED Standards | 8.50 HR |
| 1 | GN-2015B-Q-25 | Scholz | CA | NIRI | Rapid Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (North) | 5.00 HR |
| 1 | GN-2015B-Q-26 | Sharon | US | GMOS-N | "Resolving the Star Formation in Distant Galaxies" - supporting ground-based spectroscopy for a large HST program | 13.50 HR |
| 1 | GN-2015B-Q-27 | Sharon | US | GMOS-N | Multi Object Spectroscopy of the field of SDSSJ2222+2745: A cluster-lensed sextuple quasar | 4.50 HR |
| 1 | GN-2015B-Q-28 | Skemer | US | GNIRS | The First Spectrum of the Coolest Brown Dwarf | 29.00 HR |
| 1 | GN-2015B-Q-29 | Storchi-Bergmann | BR | NIFS | NIFS survey of feeding and feedback processes in nearby Active Galaxies | 15.00 HR |
| 1 | GN-2015B-Q-30 | Tominaga | Subaru | GMOS-N | Detection and Follow-up Observations of Type II Plateau Supernovae | 7.50 HR |
| 1 | GN-2015B-Q-31 | van Kerkwijk | CA | GMOS-N, GNIRS,NIRI | SN2014J at very late phases | 16.00 HR |
| 1 | GN-2015B-Q-32 | van Velzen | US | GMOS-N | Reverberation mapping of stellar tidal disruption flares | 5.00 HR |
| 1 | GN-2015B-Q-33 | Williams | UH | GRACES | Accretion & wind properties of protoplanetary disks in σ Orionis | 14.00 HR |
| 1 | GN-2015B-Q-34 | Wolf | AU | NIRI | Measuring star formation in red spirals with Paschen-Alpha | 5.00 HR |
| 2 | GN-2015B-Q-35 | Alexandroff | US | NIFS | Quasar feedback at the peak of the galaxy formation epoch | 9.00 HR |
| 2 | GN-2015B-Q-36 | Baldwin | AU | GNIRS | A comparison of spectroscopic and dynamical IMF determinations in the near-infrared | 12.50 HR |
| 2 | GN-2015B-Q-37 | Beck | J:AU/US | NIFS | Characterizing "Typical" Jets from Young Stars | 15.00 HR |
| 2 | GN-2015B-Q-38 | Cenko | US | GMOS-N | The Demographics of Tidal Disruption Flares (North) | 3.00 HR |
| 2 | GN-2015B-Q-39 | Cidale | AR | NIRI | On the structure and kinematics of the nebulae around the star MWC 137 | 2.80 HR |
| 2 | GN-2015B-Q-40 | Daemgen | CA | NIFS | Protoplanetary disk lifetimes revisited - The frequency of disks around single stars and binary components as a function of age | 6.00 HR |
| 2 | GN-2015B-Q-41 | Drahus | US | GMOS-N | Disruption of the Active Asteroid P/2012 F5 (Gibbs) | 7.00 HR |
| 2 | GN-2015B-Q-42 | Ebeling | UH | GMOS-N | GMOS imaging of massive galaxy clusters at $z>0.5$ | 13.00 HR |
| 2 | GN-2015B-Q-43 | Esplin | US | GNIRS | Searching for the Bottom of the IMF | 10.30 HR |
| 2 | GN-2015B-Q-44 | Esteves | CA | NIRI | Cloudy with a chance of storms: breaking the degeneracy between reflection and heat circulation in Kepler exoplanet atmospheres with NIRI | 6.00 HR |
| 2 | GN-2015B-Q-45 | Ferrero | AR | NIRI | Searching for the HH 31 counter-jet in the IRAS 04248+2612 | 4.00 HR |
| 2 | GN-2015B-Q-46 | Fu | US | GNIRS | How Massive Galaxies Get Cold Gas: The Circum-Galactic Medium of Dusty Starburst Galaxies at $z \sim 2$ | 12.60 HR |

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| 2 | GN-2015B-Q-47 | Glazebrook | AU/CA | GMOS-N | Observational testing of a new model of angular momentum as a fundamental driver of clumpy disks. (North) | 4.38 HR |
| 2 | GN-2015B-Q-48 | Grunhut | CA | GNIRS | Characterizing the magnetosphere of the new magnetic Bp star HD 23478 via IR emission features | 3.40 HR |
| 2 | GN-2015B-Q-49 | Horch | US | DSSI | Understanding the Metal-Poor Mass-Luminosity Relation | 4.20 HR |
| 2 | GN-2015B-Q-50 | Jayawardhana | CA | GMOS-N | Measuring the Alkali line-profiles in the transmission spectrum of XO-2b | 7.00 HR |
| 2 | GN-2015B-Q-51 | Kim | KR | GNIRS | Black hole mass versus stellar velocity dispersion relation of red AGNs | 11.00 HR |
| 2 | GN-2015B-Q-52 | Kim | KR | GNIRS | Near-Infrared Spectroscopy of Massive Young Stellar Objects with Remarkable Bipolar Outflow in Infrared Dark Cloud G53.2 | 4.00 HR |
| 2 | GN-2015B-Q-53 | Knight | US | GNIRS,NIRI | Multi-scale investigation of the coma of comet 67P/Churyumov-Gerasimenko: Combined Gemini and Rosetta study of activity | 13.40 HR |
| 2 | GN-2015B-Q-54 | Liu | AU | GRACES | Chemical signatures of planet formation in small planets hosts | 5.60 HR |
| 2 | GN-2015B-Q-55 | Liu | US | GMOS-N | IFU mapping of the most energetic BAL quasar outflows: C IV absorption | 10.80 HR |
| 2 | GN-2015B-Q-56 | Luna | AR | GMOS-N | Hydrogen burning symbiotics in M31 | 5.10 HR |
| 2 | GN-2015B-Q-57 | Magnier | UH | GNIRS | Rare Brown Dwarfs in the Solar Neighborhood from Pan-STARRS 1 | 17.00 HR |
| 2 | GN-2015B-Q-58 | Mast | BR/US | NIFS | Integral Field Spectroscopy of Arcs - Unveiling the hundreds pc of star-forming galaxies at redshift ~ 0.7 to 2.5 | 8.00 HR |
| 2 | GN-2015B-Q-59 | McConnell | CA | GMOS-N | The Black Hole Population of the Most Massive Nearby Galaxies | 11.80 HR |
| 2 | GN-2015B-Q-60 | Milisavljevic | US | GMOS-N, GNIRS | The Unprecedented Supernova Metamorphosis of SN 2014C | 3.30 HR |
| 2 | GN-2015B-Q-61 | Moskovitz | US | GMOS-N, GNIRS | Mission Accessible Near-Earth Objects Survey (MANOS) (North) | 22.04 HR |
| 2 | GN-2015B-Q-62 | Méndez | UH/AR | GRACES | A search for spectroscopic binary central stars of planetary nebulae | 15.00 HR |
| 2 | GN-2015B-Q-63 | Reggiani | BR | GRACES | Distinct halo populations revealed through precise chemical abundances of very metal-poor stars | 12.00 HR |
| 2 | GN-2015B-Q-64 | Scholz | CA | NIRI | Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (North) | 13.13 HR |
| 2 | GN-2015B-Q-65 | Sengupta | KR | GMOS-N | Dark matter content of TDGs -- kinematics of Arp 202 TDG. | 3.60 HR |
| 2 | GN-2015B-Q-66 | Spina | BR | GRACES | Unveiling chemical signatures of triggered star formation in the Orion B NGC2068/71 clusters | 10.50 HR |
| 2 | GN-2015B-Q-67 | Szkody | US | GMOS-N | Do white dwarfs in cataclysmic variables grow in mass? | 13.50 HR |
| 2 | GN-2015B-Q-68 | Tucker | AU/US | GMOS-N | Catching Supernovae in the Act with KEGS (Kepler Extra-Galactic Survey) (North) | 3.10 HR |
| 2 | GN-2015B-Q-69 | Valenti | US | GMOS-N | Nebular observations of SNe type II (North) | 6.00 HR |
| 3 | GN-2015B-Q-70 | Berg | US | GMOS-N | Ultra Extreme Star-Forming Galaxies: A Window On Low Metallicity Star Formation at High Redshift (North) | 5.20 HR |
| 3 | GN-2015B-Q-71 | Bowler | US | GRACES | Reconnaissance of Young M Dwarfs: Confirming the Elusive Majority of Nearby Moving Groups | 18.00 HR |
| 3 | GN-2015B-Q-72 | Burningham | US | GNIRS | Characterising a very fast moving T dwarf in the Galactic Plane | 2.10 HR |
| 3 | GN-2015B-Q-73 | Chiang | US/BR/CL | GMOS-N | Mapping out the Densest Structures in the COSMOS Field at $z \sim 2-3$ (North) | 20.36 HR |

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|---|---------------|-------------------|-------|--------|---|----------|
| 3 | GN-2015B-Q-74 | Cidale | AR | GNIRS | Search for molecular emission bands in the circumstellar environment of symbiotic stars | 1.60 HR |
| 3 | GN-2015B-Q-75 | Gagné | CA | GRACES | Identification and characterization of very low-mass and brown dwarf candidate members of nearby young associations | 17.30 HR |
| 3 | GN-2015B-Q-76 | Henry | US | DSSI | Searching for Companions to Nearby Stars on Solar System Scales | 40.00 HR |
| 3 | GN-2015B-Q-77 | Jeon | KR | GNIRS | Near Infrared Spectroscopic Observation of Quasars at $5 < z < 6$ | 5.20 HR |
| 3 | GN-2015B-Q-78 | Jha | US | GMOS-N | Spectroscopy of Type Ia Supernovae (North) | 5.00 HR |
| 3 | GN-2015B-Q-79 | Kellogg | CA | GNIRS | Determining the Cause of Dustiness in New Peculiar L and T Dwarfs | 14.90 HR |
| 3 | GN-2015B-Q-80 | LaMassa | US | GNIRS | Unveiling Obscured Supermassive Black Hole Growth with Infrared Spectroscopy | 12.00 HR |
| 3 | GN-2015B-Q-81 | Marino | AU | GRACES | The Milky Way globular cluster NGC6934: a remnant of a cannibalized dwarf like Omega Centauri? | 6.70 HR |
| 3 | GN-2015B-Q-82 | Meyer | CA | NIFS | A NIFS near-infrared probe of the variable IMF in nearby spiral bulges | 9.60 HR |
| 3 | GN-2015B-Q-83 | Méndez | UH | GRACES | A search for spectroscopic binary central stars of planetary nebulae | 25.00 HR |
| 3 | GN-2015B-Q-84 | Orosz | US | GMOS-N | An optical search for black holes and neutron stars in the Kepler field | 17.50 HR |
| 3 | GN-2015B-Q-85 | Rich | US/AU | GMOS-N | Quenching Caught in the Act: Measuring Feedback in Shocked Post-Starburst Galaxies | 5.30 HR |
| 3 | GN-2015B-Q-86 | Schlaufman | US | GMOS-N | An All-Sky Search for the Brightest Metal-Poor Stars (North) | 26.50 HR |
| 3 | GN-2015B-Q-87 | Stockton | UH | GMOS-N | Large Scale LINER Emission around Low-Excitation Radio Galaxies | 8.00 HR |
| 3 | GN-2015B-Q-88 | Taak | KR | NIRI | K Band Imaging of a Strong Gravitational Lens System Candidate at $z = 1.60$ | 2.20 HR |
| 3 | GN-2015B-Q-89 | Torres | AR | GNIRS | Evolutionary state and circumstellar environment of evolved massive stars | 2.10 HR |
| 3 | GN-2015B-Q-90 | Venn | CA | GRACES | The chemistry of accreted stars in the Milky Way outer halo | 4.50 HR |
| 3 | GN-2015B-Q-91 | Winkler | US | GMOS-N | Supernova Remnants in the Most Fertile Galaxy: NGC 6946 | 8.00 HR |
| 3 | GN-2015B-Q-92 | Woo | KR | GMOS-N | AGN feedback in action: extreme gas outflows in type 2 AGNs | 5.00 HR |
| 4 | GN-2015B-Q-93 | Asplund | AU | GRACES | The most metal-rich stars: probing exoplanets, stellar nucleosynthesis, Galactic archaeology, and galaxy evolution | 15.00 HR |
| 4 | GN-2015B-Q-94 | Kedziora-Chudczer | AU | GNIRS | Mapping and Characterization of Jovian Aurora in near infrared bands | 5.00 HR |
| 4 | GN-2015B-Q-95 | Manset | CA | GRACES | Unveiling the Nature of FS CMa Type Stars | 12.50 HR |
| 4 | GN-2015B-Q-96 | Yong | AU | GRACES | Unveiling the origin of the young $[\alpha/\text{Fe}]$ -rich stars | 6.70 HR |

D.2 2015B Science Programs – Gemini South

| Band | Gemini ID | PI Name | Partners | Instrument | Title | Time |
|------|---------------|--------------|----------|---------------------|---|----------|
| C | GS-2015B-C-1 | Kannappan | US | GMOS-S | Resolved Spectroscopy of a Local Volume: The RESOLVE Survey in Stripe 82 | 40.00 HR |
| 1 | GS-2015B-LP-1 | Balogh | LP | GMOS-S | The GOGREEN Survey of dense galaxy environments at $1 < z < 1.5$ (South) | 68.90 HR |
| 1 | GS-2015B-LP-2 | Fritz | LP | GMOS-S, GSAOI | Probing the dark halo of the Milky Way with GeMS/GSAOI | 27.40 HR |
| 1 | GS-2015B-LP-7 | Kasliwal | LP | GMOS-S | Rapid Spectroscopy of Elusive Transients and Young Supernovae | 4.00 HR |
| 1 | GS-2015B-LP-3 | Masiero | LP | GMOS-S | Follow up of newly discovered Near-Earth objects from the NEOWISE survey | 8.00 HR |
| 2 | GS-2015B-LP-5 | Buckley-Geer | LP | GMOS-S | Spectroscopic Confirmation and AO imaging Follow-Up of Dark Energy Survey Strong Lensing Systems and Spectra for Photometric Redshift Calibration | 80.00 HR |
| 2 | GS-2015B-LP-6 | Chen | LP | GPI | Characterizing Dusty Debris in Exoplanetary Systems | 28.00 HR |
| 1 | GS-2015B-Q-2 | Burgasser | US | GMOS-S | Mass Measurements Across the Hydrogen Burning Limit: Astrometric Orbits for Spectral Binaries | 9.20 HR |
| 1 | GS-2015B-Q-3 | Cieza | CL | GPI | GPI Characterization of the Benchmark Brown Dwarf HD4747 B | 2.60 HR |
| 1 | GS-2015B-Q-4 | Diaz/Diaz | AR | Flamingos 2, GMOS-S | The Origin of CIV absorptions at $z > 3$ | 5.00 HR |
| 1 | GS-2015B-Q-5 | Folatelli | AR | Flamingos 2 | Late-Time Near-Infrared Spectroscopy of Type Ia SN 2015F | 3.00 HR |
| 1 | GS-2015B-Q-6 | Foley | US | GMOS-S | Understanding the Progenitor Systems, Explosion Mechanisms, and Cosmological Utility of Type Ia Supernovae (South) | 4.00 HR |
| 1 | GS-2015B-Q-7 | Foley | AU/BR/US | GMOS-S | DES Supernova Cosmology (South) | 21.60 HR |
| 1 | GS-2015B-Q-8 | Galbany | CL | GMOS-S | New Approaches to Supernova Standardization for Cosmology | 10.00 HR |
| 1 | GS-2015B-Q-9 | Graham | CA/US | GPI | Astrometry of beta Pic b with GPI | 3.70 HR |
| 1 | GS-2015B-Q-10 | Gromadzki | CL | GSAOI | Trigonometric parallax of ancient T dwarf WISE0833+0052 | 2.50 HR |
| 1 | GS-2015B-Q-11 | Jensen-Clem | US | GPI | The first detection of polarized radiation from exoplanets | 9.50 HR |
| 1 | GS-2015B-Q-12 | Kavelaars | CA | GMOS-S | Tracking observations in preparation for New Horizons fly-by. | 30.00 HR |
| 1 | GS-2015B-Q-13 | Kim | KR | GMOS-S | Spectroscopy of Early and Peculiar Supernovae from the KMTNet | 3.00 HR |
| 1 | GS-2015B-Q-14 | Lacy | US | GSAOI | The highest resolution view of distant massive galaxies | 16.20 HR |
| 1 | GS-2015B-Q-15 | Lira | CL | Flamingos 2 | Reverberation Mapping of high- z QSOs: the final stages | 1.50 HR |
| 1 | GS-2015B-Q-16 | Luhman | US | GMOS-S | Testing Model Atmospheres with the Coldest Known Brown Dwarf | 6.80 HR |
| 1 | GS-2015B-Q-17 | Marois | CA/US | GPI | GPI detailed spectroscopic and astrometric characterization of HR 8799cde. | 10.00 HR |
| 1 | GS-2015B-Q-18 | Martini | US | GMOS-S | Search for $z > 6$ QSOs with the Dark Energy Survey | 10.00 HR |
| 1 | GS-2015B-Q-19 | Moskovitz | US | GMOS-S | Mission Accessible Near-Earth Objects Survey (MANOS) (South) | 6.50 HR |
| 1 | GS-2015B-Q-20 | Prieto | CL | GMOS-S | Near Explosion GMOS-South Spectroscopy of ASAS-SN and CHASE Nearby Supernovae | 3.50 HR |

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| 1 | GS-2015B-Q-21 | Richtler | CL | GMOS-S | Isolated ellipticals - key objects for the dark matter problem? | 7.00 HR |
| 1 | GS-2015B-Q-22 | Rodney | US | GMOS-S | The Next Frontier: High-Redshift Supernovae in the HST Frontier Fields (South) | 2.50 HR |
| 1 | GS-2015B-Q-23 | Rojo | CL | GPI | Investigating seasonal changes in Titan's meteorology through cloud monitoring with GPI | 1.80 HR |
| 1 | GS-2015B-Q-24 | Shannon | AU | GMOS-S | The optical counterpart to a Fast Radio Burst | 1.50 HR |
| 1 | GS-2015B-Q-25 | Steiner | BR | GMOS-S | LLP - The Gemini Survey of Galactic Nuclei - GSGN | 21.60 HR |
| 1 | GS-2015B-Q-26 | Tucker | US/AU | GMOS-S | Catching Supernovae in the Act with KEGS (Kepler Extra-Galactic Survey) (South) | 6.80 HR |
| 1 | GS-2015B-Q-27 | Webster | AU | Flamingos 2 | Understanding the Physics of the Broad Line Region | 4.00 HR |
| 1 | GS-2015B-Q-28 | Yang | KR | GMOS-S | First Systematic Survey of the Environment of Submillimeter Galaxies (South) | 7.50 HR |
| 2 | GS-2015B-Q-29 | Ammons | US | GSAOI | A GEMS Probe for Superearths Orbiting Luhman 16AB | 1.50 HR |
| 2 | GS-2015B-Q-30 | Bleem | US | GMOS-S | SPT-CL J0329?2330: Spectroscopic Observations of an Exceptional High-Redshift Galaxy Cluster | 7.30 HR |
| 2 | GS-2015B-Q-31 | Cenko | US | GMOS-S | The Demographics of Tidal Disruption Flares (South) | 3.00 HR |
| 2 | GS-2015B-Q-32 | Cerqueira | BR | GMOS-S | Spatially distributed line ratios for the HH1/2 system | 5.60 HR |
| 2 | GS-2015B-Q-33 | Chiang | US/CL/BR | GMOS-S | Mapping out the Densest Structures in the COSMOS Field at $z \sim 2.3$ (South) | 8.14 HR |
| 2 | GS-2015B-Q-34 | Christiaens | CL | GPI | Planets in the transition disk of Herbig AeBe stars | 2.80 HR |
| 2 | GS-2015B-Q-35 | Crossfield | US | GPI | Exometeorology: Searching for Weather on Beta Pictoris b | 17.00 HR |
| 2 | GS-2015B-Q-36 | Cúneo | AR | GMOS-S | Chemical abundances of polluted visible component atmosphere in black hole binary systems (South) | 4.00 HR |
| 2 | GS-2015B-Q-37 | Duchene | US | GPI | Assessing the fundamental limits of multiple star formation | 18.50 HR |
| 2 | GS-2015B-Q-38 | Faifer | AR/BR | Flamingos 2, GMOS-S | The Brazil-Argentina Gemini Group of globular Cluster systems (BAGGS): FLAMINGOS-2 and GMOS data for NGC1395 | 10.80 HR |
| 2 | GS-2015B-Q-39 | Glazebrook | AU/CA | GMOS-S | Observational testing of a new model of angular momentum as a fundamental driver of clumpy disks. (South) | 2.92 HR |
| 2 | GS-2015B-Q-40 | Kim | KR | GMOS-S | GMOS-S observation for distant galaxy clusters with extraordinarily high star formation rates | 8.00 HR |
| 2 | GS-2015B-Q-41 | Lee | KR | GMOS-S | The Evolution-Free Dark Energy Test: Spectroscopy of Early-Type Host Galaxies of Type Ia Supernovae | 6.00 HR |
| 2 | GS-2015B-Q-42 | Liu | US | GMOS-S | Spectroscopic Follow-Up of Variability Selected Binary Supermassive Black Hole Candidates | 4.80 HR |
| 2 | GS-2015B-Q-43 | Mendes de Oliveira | BR | GSAOI | Mapping the dust extinction and censusing the star cluster population in the merging system HCG 31 | 4.20 HR |
| 2 | GS-2015B-Q-44 | Millar-Blanchaer | US/CA | GPI | Constraining the Dust grain population of Beta Pic's Inner Disk | 8.50 HR |
| 2 | GS-2015B-Q-45 | Moskovitz | US | GMOS-S | Mission Accessible Near-Earth Objects Survey (MANOS) (South) | 6.50 HR |
| 2 | GS-2015B-Q-46 | Muzic | CL | Flamingos 2 | Probing the effects of environment on brown dwarf formation | 10.00 HR |
| 2 | GS-2015B-Q-47 | Opitz | US | GSAOI | Astrometry and Binarity of WISE Y dwarfs with MCAO | 4.30 HR |

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| 2 | GS-2015B-Q-48 | Piatti | AR | GMOS-S | Unveiling the origin of the LMC clusters NGC1928 and NGC1939 | 3.40 HR |
| 2 | GS-2015B-Q-49 | Pritchard | AU | Flamingos 2, GMOS-S | Spectroscopic Observations of High-z Superluminous Supernovae | 4.80 HR |
| 2 | GS-2015B-Q-50 | Rameau | CA | GPI | A GPI search for planets around remarkable young and nearby stars | 9.60 HR |
| 2 | GS-2015B-Q-51 | Rotermund | CA/US | GMOS-S | Detecting Dark Matter Subhalos with ALMA and GMOS-S/IFU using Strongly Lensed Submm Galaxies | 31.00 HR |
| 2 | GS-2015B-Q-52 | Scholz | CA | Flamingos 2 | Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (South) | 13.88 HR |
| 2 | GS-2015B-Q-53 | Shim | KR | GMOS-S | cd MgII absorbers in cluster environment | 8.00 HR |
| 2 | GS-2015B-Q-54 | Tappert | CL | GMOS-S | Recovery of old novae | 10.50 HR |
| 2 | GS-2015B-Q-55 | Telles | BR | GMOS-S | Towards precision cosmology with HII galaxies | 5.50 HR |
| 2 | GS-2015B-Q-56 | Yanny | US/BR | GMOS-S | Imaging Follow-up of Milky-Way companions revealed by the Dark Energy Survey | 14.89 HR |
| 3 | GS-2015B-Q-57 | Baron | CA | Flamingos 2, GMOS-S | Wide Imaging Search for Benchmark Planets | 23.00 HR |
| 3 | GS-2015B-Q-58 | Bessiere | CL | GMOS-S | The evolution of quasar host galaxies | 11.00 HR |
| 3 | GS-2015B-Q-59 | Chanchaiworawit | US | GMOS-S | Spectroscopic Confirmation of LAE Candidates at $z=6.5$: Evidence of the highest redshift protocluster | 46.00 HR |
| 3 | GS-2015B-Q-60 | Diaz/Diaz | AR | Flamingos 2 | Galactic feedback and the circum-galactic medium of $z \sim 2.3$ galaxies | 3.50 HR |
| 3 | GS-2015B-Q-61 | Faundez-Abans | BR | GMOS-S | Kinematics of the peculiar galaxy AM 0405-371 | 1.90 HR |
| 3 | GS-2015B-Q-62 | Graham | US | GMOS-S | Understanding the Power Source in Type Ia Supernovae with Nebular Phase Spectroscopy | 18.30 HR |
| 3 | GS-2015B-Q-63 | Hernandez-Jimenez | BR | GMOS-S | Constraining two physically distinct inflow mechanisms in the minor merger AM0737-764 | 3.90 HR |
| 3 | GS-2015B-Q-64 | Hsieh | US | GMOS-S | Observations of the Activity of Main-Belt Comets P/2010 R2, 233P, and 313P (South) | 6.00 HR |
| 3 | GS-2015B-Q-65 | Jeong | KR | GMOS-S | Nature of high-z SMGs in ADF-S | 0.00 HR |
| 3 | GS-2015B-Q-66 | Kim | KR | GMOS-S | Search for Activity in Lixiaohua Family Asteroids (South) | 0.07 HR |
| 3 | GS-2015B-Q-67 | Lee | KR | Flamingos 2 | NIR spectroscopy of strong line emitting SNRs in nearby galaxies (South) | 0.00 HR |
| 3 | GS-2015B-Q-68 | M.Manseau | CA | GMOS-S | A Spectroscopic Analysis of Chemically Stratified DAO White Dwarfs in the Sloan Digital Sky Survey (South) | 4.00 HR |
| 3 | GS-2015B-Q-69 | Mendes de Oliveira | BR | GMOS-S | The Cosmic Penguin cooking an AGN and a Tidal Dwarf Galaxy. | 6.10 HR |
| 3 | GS-2015B-Q-70 | Parisi | AR/CL | GMOS-S | Accurate Ages, Metal Abundances and Kinematics of a Large Sample of Small Magellanic Cloud Star Clusters and Surrounding Fields | 11.00 HR |
| 3 | GS-2015B-Q-71 | Schlaufman | US | GMOS-S | An All-Sky Search for the Brightest Metal-Poor Stars (South) | 42.90 HR |
| 3 | GS-2015B-Q-72 | Sung | KR | GMOS-S | Precision measurements of the central dark matter distribution in low mass dwarf galaxies: GMOS-N long-slit absorption line spectroscopy of DDO 210 | 10.00 HR |
| 3 | GS-2015B-Q-73 | Wade | CA | GMOS-S | Monitoring the magnetospheric structure of the first candidate extra-Galactic magnetic massive stars | 12.30 HR |
| 3 | GS-2015B-Q-74 | Webster | AU | Flamingos 2 | Understanding the Physics of the Broad Line Region | 12.71 HR |

D.3 2016A Science Programs – Gemini North

| Band | Gemini ID | PI Name | Partners | Instrument | Title | Time |
|------|---------------|------------------|----------|---------------------|--|------------|
| C | GN-2016A-C-1 | Baldwin | AU | GNIRS | A comparison of spectroscopic and dynamical IMF determinations in the near-infrared | 10.00 HR |
| C | GN-2016A-C-2 | Levesque | US | GMOS-N | Red Supergiants in the Extremely Metal Poor Sextans Galaxies: Local Analogs of Massive Stars in the Early Universe | 2.00 NIGHT |
| 1 | GN-2016A-LP-1 | Kasliwal | LP | GMOS-N | Rapid Spectroscopy of Elusive Transients and Young Supernovae | 23.10 HR |
| 1 | GN-2016A-LP-2 | Huitson | LP | GMOS-N | The First Survey Dedicated to the Detection and Characterization of Clouds in Exoplanet Atmospheres | 12.00 HR |
| 1 | GN-2016A-LP-3 | Fraser | LP | GMOS-N,NIRI | COL-OSSOS: COLOURS for the Outer Solar System Object Survey | 50.00 HR |
| 1 | GN-2016A-LP-4 | Balogh | LP | GMOS-N | The GOGREEN Survey of dense galaxy environments at $1 < z < 1.5$ | 45.50 HR |
| 2 | GN-2016A-LP-5 | Shen | LP | GNIRS | A GNIRS Near-IR Spectroscopic Survey of $z > 5.7$ Quasars | 35.90 HR |
| 2 | GN-2016A-LP-6 | Huitson | LP | GMOS-N | The First Survey Dedicated to the Detection and Characterization of Clouds in Exoplanet Atmospheres | 11.80 HR |
| 2 | GN-2016A-LP-7 | Crossfield | LP | NIRI | Validating K2's Habitable and Rocky Planets with AO Imaging | 10.00 HR |
| 1 | GN-2016A-Q-1 | Yoon | KR | GRACES | Search for signatures of pair-instability supernova nucleosynthesis in alpha-enhanced metal poor stars | 10.60 HR |
| 1 | GN-2016A-Q-2 | Yang | KR | GMOS-N | Star Formation Efficiency and AGN Feedback in Gas-Rich Post-Starburst Galaxies | 12.00 HR |
| 1 | GN-2016A-Q-3 | van Velzen | US | GMOS-N | Reverberation mapping of stellar tidal disruption flares | 6.20 HR |
| 1 | GN-2016A-Q-4 | van Kerkwijk | CA | GMOS-N, GNIRS, NIRI | SN2014J at very late phases | 18.00 HR |
| 1 | GN-2016A-Q-5 | Tominaga | Subaru | GMOS-N | Detection and Follow-up Observations of Type II Plateau Supernovae | 7.50 HR |
| 1 | GN-2016A-Q-6 | Storchi-Bergmann | BR | NIFS | NIFS survey of feeding and feedback processes in nearby Active Galaxies | 15.00 HR |
| 1 | GN-2016A-Q-7 | Stern | US | GNIRS | Spatially Resolving the Kinematics of the ~100 uas Quasar Broad Line Region Using Spectroastrometry | 15.00 HR |
| 1 | GN-2016A-Q-8 | Shim | KR | GRACES | Probing Circumgalactic Medium of Cluster Galaxies using Background QSO | 7.50 HR |
| 1 | GN-2016A-Q-9 | Seth | US | GNIRS | A GNIRS Survey of the Nearest Nuclear Star Clusters | 13.10 HR |
| 1 | GN-2016A-Q-10 | Scholz | CA | NIRI | Rapid Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (North) | 5.00 HR |
| 1 | GN-2016A-Q-11 | Sardane | US | GMOS-N | Mapping cool, metal-rich gas around $z < 0.08$ galaxies | 4.00 HR |
| 1 | GN-2016A-Q-12 | Ross | CA | NIFS | Testing the triggering mechanisms of SMGs with spatially-resolved dynamics | 14.00 HR |
| 1 | GN-2016A-Q-13 | Moskovitz | US | GMOS-N,GNIRS | Mission Accessible Near-Earth Objects Survey (MANOS) (North) | 22.50 HR |
| 1 | GN-2016A-Q-14 | Mendez | UH | GRACES | A search for spectroscopic binary central stars of planetary nebulae | 12.00 HR |
| 1 | GN-2016A-Q-15 | Meech | UH | GMOS-N | The Manx Comets – Testing Solar System Formation Models | 6.90 HR |
| 1 | GN-2016A-Q-16 | Liss | US | GMOS-N | Star Clusters in Interacting Dwarf Galaxies | 11.60 HR |

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| 1 | GN-2016A-Q-17 | Lee | KR | GRACES | High-Resolution Spectroscopy of Candidate Ultra Metal-Poor Stars from Sloan Digital Sky Survey | 9.80 HR |
| 1 | GN-2016A-Q-18 | Kim | KR | NIFS | New, unveiled epoch of IRC+10216 | 3.00 HR |
| 1 | GN-2016A-Q-19 | Karouzos | KR | GMOS-N | Unveiling circumnuclear star formation: AGN feeding vs feedback? | 12.40 HR |
| 1 | GN-2016A-Q-20 | Jofr  | AR | GRACES | The chemical signature of giant stars with planets | 3.10 HR |
| 1 | GN-2016A-Q-21 | Im | KR | GMOS-N | Spectroscopic Identification of A Faint Quasar at $z \sim 7$ | 9.00 HR |
| 1 | GN-2016A-Q-22 | Hosek | UH | NIFS | The Arches Cluster: Initial Mass Function and Stellar Evolution | 13.00 HR |
| 1 | GN-2016A-Q-23 | Ho | UH | NIRI | Gemini/NIRI narrowband imaging of H2 1-0 S(1): probing the structure of galactic winds in the local Universe | 6.00 HR |
| 1 | GN-2016A-Q-24 | Hinkle | US | NIRI | Imaging the Expanding Debris Cloud around Sakurai's Object | 1.00 HR |
| 1 | GN-2016A-Q-25 | Hamann | US | GNIRS | Feedback vs Cold Mode Accretion: Accurate Velocities for Complex Infall/Outflow Gas Near Redshift 3 Quasars | 1.70 HR |
| 1 | GN-2016A-Q-26 | Hagelberg | UH | GMOS-N | Probing the atmosphere of an evaporating Jupiter candidate | 4.40 HR |
| 1 | GN-2016A-Q-27 | Gonzalez | US | GMOS-N | A Metallicity Determination for the Diffuse Emission around NGC 4874 in the Coma Cluster | 11.70 HR |
| 1 | GN-2016A-Q-28 | Gladders | US | GMOS-N | Time delays for the sextuply-lensed quasar SDSS J2222+2745 from GMOS | 3.80 HR |
| 1 | GN-2016A-Q-29 | Garnavich | US | GMOS-N | Return to the Heart of Darkness: An Unbiased Survey of Void Galaxies | 13.00 HR |
| 1 | GN-2016A-Q-30 | Frye | US | GMOS-N | Spectroscopic redshifts of the lenses producing the brightest high- z sources in the Planck all-sky survey | 10.00 HR |
| 1 | GN-2016A-Q-31 | Esplin | US | GNIRS | Searching for the bottom of the IMF in Ophiuchus and RCrA (North) | 3.20 HR |
| 1 | GN-2016A-Q-32 | Ebeling | UH | GMOS-N | Jellyfish: the dynamics of extreme ram-pressure stripping | 4.00 HR |
| 1 | GN-2016A-Q-33 | de Kleer | US | NIRI | Observing Io's volcanoes in eclipse: Eruption temperatures and SO gas content | 3.00 HR |
| 1 | GN-2016A-Q-34 | de Kleer | US | NIRI | Linking Io's Volcanic Activity to Plasma Torus Variability | 20.00 HR |
| 1 | GN-2016A-Q-35 | Cote | CA | NIFS | Black Holes in the Smallest Galaxies | 10.90 HR |
| 1 | GN-2016A-Q-36 | Chambers | UH | GMOS-N,GNIRS | Pan-STARRS Survey for Counterparts to ALIGO GW events | 4.00 HR |
| 1 | GN-2016A-Q-37 | Bowler | US | NIFS | Spectroscopic Confirmation of the Planetary Companion ROXs12 b | 7.00 HR |
| 1 | GN-2016A-Q-38 | Beck | US | NIFS | Understanding protostellar jet launching in Herbig Stars | 2.50 HR |
| 1 | GN-2016A-Q-39 | Almeida | BR | Graces | Radial velocity confirmation of an exoplanet candidate around an evolved compact binary | 3.75 HR |
| 2 | GN-2016A-Q-40 | Tremblay | US | NIFS | A Multi-Wavelength Approach to AGN Feedback and Star- Formation | 20.00 HR |
| 2 | GN-2016A-Q-41 | Stockton | UH | GNIRS | The Cold-Gas Environments of Sub-Millimeter Galaxies | 10.00 HR |
| 2 | GN-2016A-Q-42 | Simpson | US | GNIRS | Characterizing the Candidate Planetary Nebulae in the Galactic Center with Gemini GNIRS | 1.00 HR |
| 2 | GN-2016A-Q-43 | Scholz | CA | NIRI | Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (North) | 13.13 HR |

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| 2 | GN-2016A-Q-44 | Sand | US | GMOS-N | The first high quality Mg II reverberation lag measurement (South) | 11.00 HR |
| 2 | GN-2016A-Q-45 | Salim | US | GMOS-N | Metallicity calibrations for high-redshift galaxies | 31.00 HR |
| 2 | GN-2016A-Q-46 | Pavlov | US | GMOS-N, NIRI | Confirming the nature of the knot near pulsar B1951+32 | 2.20 HR |
| 2 | GN-2016A-Q-47 | Moskovitz | US | GMOS-N, GNIRS | Mission Accessible Near-Earth Objects Survey (MANOS) (North) | 22.50 HR |
| 2 | GN-2016A-Q-48 | Liu | US | GMOS-N | Signposts of Quasar Feedback in the High-Redshift Universe | 15.00 HR |
| 2 | GN-2016A-Q-49 | Liu | US | GMOS-N | IFU mapping of the most energetic BALQSO outflows | 7.20 HR |
| 2 | GN-2016A-Q-50 | Leggett | US | NIRI | Near-Infrared Imaging of our 250K Neighbor | 24.00 HR |
| 2 | GN-2016A-Q-51 | Lee | KR | GMOS-N | The Evolution-Free Dark Energy Test: Spectroscopy of Early-Type Host Galaxies of Type Ia Supernovae | 8.00 HR |
| 2 | GN-2016A-Q-52 | Lee | KR | GMOS-N | Stellar population gradients of local early-type galaxies in different NUV and MIR classes | 5.00 HR |
| 2 | GN-2016A-Q-53 | Knight | US | GMOS-N NIRI | Multi-scale investigation of the coma of comet 67P/Churyumov-Gerasimenko: Combined Gemini and Rosetta study of activity | 10.20 HR |
| 2 | GN-2016A-Q-54 | Kilic | US | GMOS-N | The Shortest Period Binary White Dwarfs in SDSS DR10 (North) | 22.45 HR |
| 2 | GN-2016A-Q-55 | Harrison | US | GNIRS | Quantification of the ^{13}C Abundance in the Secondary Stars of Cataclysmic Variables | 1.80 HR |
| 2 | GN-2016A-Q-56 | Hamann | US | GRACES | A Remarkable New Transient Outflow in the Quasar PG1411+442 | 1.30 HR |
| 2 | GN-2016A-Q-57 | Hall | CA | GMOS-N | Contemporaneous spectra of Chandra-observed BAL quasars | 8.50 HR |
| 2 | GN-2016A-Q-58 | Greene | US | GMOS-N | Black Holes in Dwarf Galaxies at Intermediate Redshift | 11.00 HR |
| 2 | GN-2016A-Q-59 | Gianninas | CA/US | GMOS-N | Searching for Pulsations in Mixed Atmosphere Extremely Low-mass White Dwarfs | 4.10 HR |
| 2 | GN-2016A-Q-60 | Ebeling | UH | GMOS-N | GMOS imaging of massive galaxy clusters at $z > 0.5$ | 10.00 HR |
| 2 | GN-2016A-Q-61 | den Brok | US | NIFS | Massive black holes in small galaxies | 9.50 HR |
| 2 | GN-2016A-Q-62 | Cortesi | AR/BR | GMOS-N | The Brazil-Argentina Gemini Group of globular Cluster systems (BAGGS): Tracing the star formation history of the S0 NGC4382 with GMOS spectroscopy of globular clusters and field stars | 9.10 HR |
| 2 | GN-2016A-Q-63 | Chu | UH | GNIRS | Investigating the Structure and Composition of Molecular Clouds in Preparation for the James Webb Space Telescope | 9.75 HR |
| 2 | GN-2016A-Q-64 | Chies Santos | BR | GMOS-N | Probing Assembly Histories of MASSIVE Survey Galaxies from their Globular Cluster Colors | 7.00 HR |
| 2 | GN-2016A-Q-65 | Charbonnier | BR | NIRI | In the outskirts of compact massive quiescent galaxies at intermediate redshifts | 3.20 HR |
| 2 | GN-2016A-Q-66 | Chapman | CA | GMOS-N, GNIRS | Redshifts for the SCUBA-2 CLS brightest SMGs. | 16.00 HR |
| 2 | GN-2016A-Q-67 | Carlin | US | GRACES | Recreating the chemical evolution of the Sagittarius dwarf spheroidal from its tidal debris | 37.60 HR |
| 2 | GN-2016A-Q-68 | Bresolin | UH | GMOS-N | Metallicity gradients in the smallest spiral galaxies | 6.40 HR |
| 2 | GN-2016A-Q-69 | Bassino | AR | GMOS-N | Reconstructing the history of highly disturbed galaxies by means of their globular cluster systems | 3.30 HR |

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|---|---------------|----------------|-------|------------|---|----------|
| 2 | GN-2016A-Q-70 | Antoniadis | CA | GMOS-N | Revisiting the mass of the millisecond pulsar J1012+5307 | 12.00 HR |
| 2 | GN-2016A-Q-71 | Almeida | BR | GRACES | Radial velocity confirmation of an exoplanet candidate around an evolved compact binary | 3.75 HR |
| 3 | GN-2016A-Q-72 | Wade | CA | GRACES | The nature of photometric variability in white dwarf stars | 8.40 HR |
| 3 | GN-2016A-Q-73 | Valenti | US | GMOS-N | Nebular observations of SNe type II (North) | 4.60 HR |
| 3 | GN-2016A-Q-74 | Stockton | UH | GNIRS | The Cold-Gas Environments of Sub-Millimeter Galaxies | 10.00 HR |
| 3 | GN-2016A-Q-75 | Schlaufman | US | GMOS-N | An All-Sky Search for the Brightest Metal-poor Stars (North) | 79.40 HR |
| 3 | GN-2016A-Q-76 | Saker | AR | GMOS-N | Gaseous disks in white dwarfs | 2.80 HR |
| 3 | GN-2016A-Q-77 | S. Couto | BR | GMOS-N | Probing the relation between the radio jet and the circumnuclear gas in the radio galaxy 4C+29.30 | 4.50 HR |
| 3 | GN-2016A-Q-78 | Richardson | CA | GRACES | The Unusual Pulsations from the Interacting Binary MWC 314 | 8.10 HR |
| 3 | GN-2016A-Q-79 | Mirabel | US/AR | GNIRS,NIRI | STAR FORMATION TRIGGERED BY MICROQUASAR JETS | 1.90 HR |
| 3 | GN-2016A-Q-80 | Miller | CA | GMOS-N | Globular Cluster Spectroscopy of dEs in the Local Volume (North) | 12.61 HR |
| 3 | GN-2016A-Q-81 | Mendez | UH | GRACES | A search for spectroscopic binary central stars of planetary nebulae | 20.00 HR |
| 3 | GN-2016A-Q-82 | Malo | CA | GRACES | Identification and characterization of very low-mass and brown dwarf candidate members of nearby young associations | 13.80 HR |
| 3 | GN-2016A-Q-83 | Liu | US | GMOS-N | Hunting for Massive Binary Black Holes with Continued Quasar Spectroscopic Monitoring | 12.00 HR |
| 3 | GN-2016A-Q-84 | Lee | KR | GNIRS | GNIRS spectroscopy of strong line emitting SNRs in nearby galaxies | 2.32 HR |
| 3 | GN-2016A-Q-85 | Krafton | US | GMOS-N | Late-Time Dust Formation in Core-Collapse Supernovae | 10.40 HR |
| 3 | GN-2016A-Q-86 | Kim | KR | GNIRS | Black hole mass versus stellar velocity dispersion relation of red AGNs | 7.00 HR |
| 3 | GN-2016A-Q-87 | Jha | US | GMOS-N | Spectroscopy of Type Ia Supernovae (North) | 1.99 HR |
| 3 | GN-2016A-Q-88 | Hsieh | US | GMOS-N | Main-Belt Comet Activity and Nucleus Characterization (North) | 2.00 HR |
| 3 | GN-2016A-Q-89 | Herczeg | US | GMOS-N | Accretion and the formation of very low mass objects (North) | 17.80 HR |
| 3 | GN-2016A-Q-90 | Hamann | US | GNIRS | Outflows & Feedback in Extremely Red Quasars | 30.00 HR |
| 3 | GN-2016A-Q-91 | Ferrero | AR | GNIRS | A High Resolution Near Infrared Spectral Atlas of O stars with Gemini | 2.00 HR |
| 3 | GN-2016A-Q-92 | Cidale | AR | GNIRS | Unveiling the nature of the unclassified B[e] star MWC 819 | 1.90 HR |
| 3 | GN-2016A-Q-93 | Chies Santos | CA/US | GMOS-N | Probing Assembly Histories of MASSIVE Survey Galaxies from their Globular Cluster Colors | 17.00 HR |
| 4 | GN-2016A-Q-94 | Smith Castelli | AR | GMOS-N | Stellar population and Initial Mass Function analysis of the bright elliptical galaxies NGC 7619 and NGC 7626 (North) | 5.60 HR |
| 4 | GN-2016A-Q-95 | Manset | CA | GRACES | Unveiling the Nature of FS CMa Type Stars | 7.60 HR |
| 4 | GN-2016A-Q-96 | Cochetti | AR | GNIRS | Be stars spectroscopic variability in the infrared | 4.30 HR |

D.4 2016A Science Programs – Gemini South

| Band | Gemini ID | PI Name | Partners | Instrument | Title | Time |
|------|---------------|--------------|----------|---------------------|---|------------|
| C | GS-2016A-C-1 | Kool | AU | GSAOI | Project SUNBIRD: Supernovae UNmasked By Infra-Red Detection | 2.00 NIGHT |
| 1 | GS-2016A-LP-1 | Balogh | LP | GMOS-S | The GOGREEN Survey of dense galaxy environments at $1 < z < 1.5$ | 24.70 HR |
| 1 | GS-2016A-LP-2 | Fritz | LP | GMOS-S, GSAOI | Probing the dark halo of the Milky Way with GeMS/GSAOI | 20.40 HR |
| 1 | GS-2016A-LP-3 | Masiero | LP | GMOS-S | Follow up of newly discovered Near-Earth objects from the NEOWISE survey | 8.00 HR |
| 1 | GS-2016A-LP-7 | Kasliwal | LP | GMOS-S | Rapid Spectroscopy of Elusive Transients and Young Supernovae | 4.00 HR |
| 2 | GS-2016A-LP-5 | Buckley-Geer | LP | GMOS-S | Spectroscopic Confirmation and AO imaging Follow-Up of Dark Energy Survey Strong Lensing Systems and Spectra for Photometric Redshift Calibration | 12.00 HR |
| 2 | GS-2016A-LP-6 | Chen | LP | GPI | Characterizing Dusty Debris in Exoplanetary Systems | 19.80 HR |
| 2 | GS-2016A-LP-8 | Hynes | LP | GMOS-S | Dynamical Masses of Black Holes and Neutron Stars from the Galactic Bulge Survey | 30.00 HR |
| 2 | GS-2016A-LP-9 | Crossfield | LP | DSSI | Validating K2's Habitable and Rocky Planets with AO Imaging | 10.00 HR |
| 1 | GS-2016A-Q-1 | Tappert | CL | GMOS-S | Recovery of old novae | 3.30 HR |
| 1 | GS-2016A-Q-2 | Strauss | US | GMOS-S | Spectroscopic identification of new low-luminosity quasars at $z > 6$ | 27.90 HR |
| 1 | GS-2016A-Q-3 | Steiner | BR | GMOS-S | LLP - The Gemini Survey of Galactic Nuclei - GSGN | 17.00 HR |
| 1 | GS-2016A-Q-4 | Smith | US | Phoenix | [Fe II] Kinematics of Galactic Luminous Blue Variables | 20.00 HR |
| 1 | GS-2016A-Q-5 | Sheppard | US | GMOS-S | The Inner Oort Cloud Population | 2.00 HR |
| 1 | GS-2016A-Q-6 | Scholz | CA | Flamingos 2 | Rapid Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (South) | 5.20 HR |
| 1 | GS-2016A-Q-7 | Sawicki | CA | GMOS-S | Spectroscopy of very luminous $z \sim 6$ galaxies from the HSC-Wide survey | 15.60 HR |
| 1 | GS-2016A-Q-8 | Rettura | US | Flamingos 2, GMOS-S | The Gemini survey of the most distant galaxy clusters in the $\sim 100\text{deg}^2$ Spitzer-SPT Deep Field | 13.00 HR |
| 1 | GS-2016A-Q-9 | Rest | US | GMOS-S | Spectrophotometric Time Series of Eta Carinae's Great Eruption | 22.60 HR |
| 1 | GS-2016A-Q-10 | Principe | CL | GPI | Searching for Evidence of Planet Formation in a Nearby Solar Nebula Analog | 4.10 HR |
| 1 | GS-2016A-Q-11 | Piskorz | US | DSSI | Imaging Friends of Hot Jupiters | 0.80 HR |
| 1 | GS-2016A-Q-12 | Patience | CA | GPI | Characterizing the atmosphere of the imaged planet HD 95086 b with GPI | 4.00 HR |
| 1 | GS-2016A-Q-13 | Moskovitz | US | GMOS-S | Mission Accessible Near-Earth Objects Survey (MANOS) (South) | 6.50 HR |
| 1 | GS-2016A-Q-14 | Montet | US | DSSI | Fundamental Parameters of Young M Dwarfs | 7.30 HR |
| 1 | GS-2016A-Q-15 | Meshkat | US | GPI | Occurrence of giant planets in the dustiest new WISE debris disk systems | 6.00 HR |
| 1 | GS-2016A-Q-16 | Mennickent | CL | Flamingos 2 | Infrared detection of circumbinary planets around NN Ser | 3.00 HR |
| 1 | GS-2016A-Q-17 | McCollum | US | Flamingos 2, GSAOI | Identifying the Progenitor Star for a New Red Transient | 3.40 HR |
| 1 | GS-2016A-Q-18 | Mauro | AR/CL | GSAOI | Searching for past multiple bursts of star formation in Terzan 5, a fossil relic of the Galactic bulge | 9.30 HR |

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|---|---------------|------------------|----------|---------------------|--|----------|
| 1 | GS-2016A-Q-19 | Lira | CL | Flamingos 2 | Reverberation Mapping of high-z QSOs: the final stages | 1.50 HR |
| 1 | GS-2016A-Q-20 | Lin | US | GMOS-S | Measuring the Distance to an Intermediate-mass Black Hole with Gemini | 3.90 HR |
| 1 | GS-2016A-Q-21 | Levesque | US | Flamingos 2, GMOS-S | Rapid Spectroscopy of SN 2010da: Observing a Neutron Star + Luminous Blue Variable Binary in Outburst | 3.00 HR |
| 1 | GS-2016A-Q-22 | Lee | KR | GSAOI | Understanding Clumping in circumstellar shells around Luminous Blue Variables | 2.00 HR |
| 1 | GS-2016A-Q-23 | Kraus | US/CL/CA | GPI | The Planetary Systems of Young Massive Stars | 30.40 HR |
| 1 | GS-2016A-Q-24 | KIM | KR | GMOS-S | Optical spectroscopy of Early and Peculiar Supernovae found by the KMTNet | 3.00 HR |
| 1 | GS-2016A-Q-25 | Kilic | US/AR | GMOS-S | A Benchmark Pulsating White Dwarf Companion to PSR J1738+0333 | 11.00 HR |
| 1 | GS-2016A-Q-26 | Horch | US | DSSI | Towards a True Population II Mass-Luminosity Relation | 5.30 HR |
| 1 | GS-2016A-Q-27 | Hirano | Subaru | GSAOI | ESPRINT II: A New Search for Transiting Planets Unveiled by K2 | 3.75 HR |
| 1 | GS-2016A-Q-28 | Henry | US | DSSI | Searching for Companions to Nearby Stars on Solar System Scales | 25.00 HR |
| 1 | GS-2016A-Q-29 | Gromadzki | CL | GSAOI | Trigonometric parallax of ancient T dwarf WISE0833+0052 (2016A) | 2.50 HR |
| 1 | GS-2016A-Q-30 | Greenbaum | US | GPI | Probing the inner regions of HD142527 and HD100546 | 4.50 HR |
| 1 | GS-2016A-Q-31 | Ghez | US | GSAOI | Using MCAO to Enable Unique Test of General Relativity at the Galactic Center | 7.20 HR |
| 1 | GS-2016A-Q-32 | Esplin | US | Flamingos 2 | Searching for the bottom of the IMF in Ophiuchus and RCrA (South) | 2.70 HR |
| 1 | GS-2016A-Q-33 | Dong | US | GPI | Planet-drive spiral arms in SAO 206462: a direct test of dynamical model predictions | 2.50 HR |
| 1 | GS-2016A-Q-34 | Do | US | GSAOI | Measuring the orbital history of the ultra-faint dwarf galaxy Hercules with GSAOI | 10.50 HR |
| 1 | GS-2016A-Q-35 | Currie | CA | GPI | Confirming and Characterizing Young Planetary Companions with GPI | 5.50 HR |
| 1 | GS-2016A-Q-36 | Currie | Subaru | GPI | A Sensitive GPI Search for Super-Jovian Planets Orbiting WISE-Detected Early-Type Stars | 15.00 HR |
| 1 | GS-2016A-Q-37 | Cuadra | CL | GSAOI | Star formation and dynamics at < 2.5 pc from Sgr A* | 6.80 HR |
| 1 | GS-2016A-Q-38 | Cohen | CL | GSAOI | Unveiling the Heart of the Milky Way: Ages and Structural Parameters of Inner Galactic Bulge Globular Clusters | 2.40 HR |
| 1 | GS-2016A-Q-39 | Cassata | CL | Flamingos 2 | Compactness-dependent quenching in z~2 galaxies? | 10.00 HR |
| 1 | GS-2016A-Q-40 | Brandt | US | GPI | Imaging the Only Known White Dwarf with a >8 Msun Progenitor | 1.50 HR |
| 1 | GS-2016A-Q-41 | Bary | US | Phoenix | High-Resolution Spectroscopy of Orbitally-Modulated Accretion Activity in Pre-Main Sequence Binaries | 11.00 HR |
| 1 | GS-2016A-Q-42 | Arroway | CA | GMOS-S | Tracking Observations in Preparation for New Horizons FlyBy Kuiper Belt Objects | 11.90 HR |
| 1 | GS-2016A-Q-43 | Arias | AR | Phoenix | Molecular emission bands in the symbiotic star BI Crucis | 1.40 HR |
| 2 | GS-2016A-Q-44 | Urrutia-Viscarra | CL | GMOS-S | On the 2D structure and kinematics of the tidal dwarf candidate in the gas-rich interacting galaxy pair NGC 3166/9 | 8.00 HR |
| 2 | GS-2016A-Q-45 | Treister | CL | GMOS-S | Resolving the peak of the Cosmic X-ray Background: Optical Follow-up for the NuSTAR Serendipitous Survey | 10.20 HR |

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| 2 | GS-2016A-Q-46 | Torres-Flores | CL | GMOS-S | Understanding the origin of a peculiar star-forming region in NGC 6845A | 3.20 HR |
| 2 | GS-2016A-Q-47 | Scholz | CA | Flamingos 2 | Target of Opportunity Gemini Infrared Observations of Magnetars in Outburst (South) | 13.88 HR |
| 2 | GS-2016A-Q-48 | Sawicki | CA | GMOS-S | Spectroscopy of very luminous $z \sim 6$ galaxies from the HSC-Wide survey | 10.40 HR |
| 2 | GS-2016A-Q-49 | Sand | US | GMOS-S | The first high quality Mg II reverberation lag measurement (South) | 22.20 HR |
| 2 | GS-2016A-Q-50 | Rodriguez | CL | GMOS-S | Testing Type II supernova as cosmological probes at near-infrared wavelength | 6.50 HR |
| 2 | GS-2016A-Q-51 | Parsons | CL | GMOS-S | Accurate component masses and radii of three white dwarf-M dwarf binaries observed by Kepler | 16.00 HR |
| 2 | GS-2016A-Q-52 | Nogueira-Cavalcante | BR | GMOS-S | Star Formation Quenching in Different Morphological Types of Green Valley Galaxies at Intermediate Redshifts | 10.80 HR |
| 2 | GS-2016A-Q-53 | Moskovitz | US | GMOS-S | Mission Accessible Near-Earth Objects Survey (MANOS) (South) | 6.50 HR |
| 2 | GS-2016A-Q-54 | McConnachie | CA | Flamingos 2 | Precision photometry of globular clusters with GeMS/GSAOI: the need for FLAMINGOS-2 | 1.40 HR |
| 2 | GS-2016A-Q-55 | Mauerhan | US | GSAOI | GeMS narrowband imaging of the massive star factory Sagittarius B | 4.20 HR |
| 2 | GS-2016A-Q-56 | Liu | US | GMOS-S | Spectroscopic Follow-Up of Variability Selected Binary Supermassive Black Hole Candidates | 1.60 HR |
| 2 | GS-2016A-Q-57 | Kong | US | GMOS-S | A Spectroscopic Study of the Black Hole Binary MAXI J1659-152 in Quiescence | 2.30 HR |
| 2 | GS-2016A-Q-58 | Kilic | US | GMOS-S | The Shortest Period Binary White Dwarfs in SDSS DR10 (South) | 3.20 HR |
| 2 | GS-2016A-Q-59 | Im | KR | GMOS-S | Spectroscopic Confirmation of Faint Quasars at $z \sim 6$ | 6.70 HR |
| 2 | GS-2016A-Q-60 | Hargreaves | US | Phoenix | Hydrocarbon emission in the polar regions of Jupiter | 16.70 HR |
| 2 | GS-2016A-Q-61 | Graham | US | GMOS-S | Understanding the Power Source in Type Ia Supernovae with Nebular Phase Spectroscopy | 23.90 HR |
| 2 | GS-2016A-Q-62 | Ferrero | AR | GSAOI | GSAOI+GeMS high resolution images of protostellar jets: The case of HH 137/138 | 2.80 HR |
| 2 | GS-2016A-Q-63 | Crnojevic | US | Flamingos 2 | Environment and the evolution at low-mass galactic scales: clues from the Cen A group | 22.10 HR |
| 2 | GS-2016A-Q-64 | Cody | US | DSSI | The effect of binarity on protoplanetary disk dissipation and accretion | 30.00 HR |
| 2 | GS-2016A-Q-65 | Cerqueira | BR | GMOS-S | Investigating different excitation conditions in the HH 228 jet | 8.80 HR |
| 2 | GS-2016A-Q-66 | Caso | AR | GMOS-S | Globular clusters in Antlia: a wider photometric look | 1.80 HR |
| 2 | GS-2016A-Q-67 | Carrasco | US | GSAOI | Investigating galaxy structure and mass assembly at $z > 3$ with GeMS/GSAOI | 20.00 HR |
| 2 | GS-2016A-Q-68 | Baume | AR | Flamingos 2 | Embedded clusters in the Carina Galactic Arm | 1.60 HR |
| 2 | GS-2016A-Q-69 | Baron | CA | Flamingos 2, GMOS-S | WEIRD : Wide orbit Exoplanet search with InfraRed Direct imaging | 15.70 HR |
| 2 | GS-2016A-Q-70 | Bahramian | US | Flamingos 2 | Determining the nature of donor stars in sub-luminous transient X-ray binaries | 7.60 HR |
| 2 | GS-2016A-Q-71 | Angeloni | US | GMOS-S | Spectroscopic follow-up of puzzling OGLE sources: a new class of variable stars? | 5.00 HR |
| 2 | GS-2016A-Q-72 | Ammons | US | GSAOI | The Best Parallaxes and a Limit on the Giant Planet Occurrence Rate for Nearby T Dwarf Hosts | 4.20 HR |

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|---|---------------|------------------------|----|---------------------|---|----------|
| 2 | GS-2016A-Q-73 | Abia | US | Phoenix | Extragalactic AGB carbon stars as probes to test the origin of fluorine in the universe. | 20.30 HR |
| 3 | GS-2016A-Q-74 | Weidmann | AR | GMOS-S | Determining the actual nature of weak emission line stars in the nucleus of Planetary Nebulae | 2.10 HR |
| 3 | GS-2016A-Q-75 | Valenti | US | GMOS-S | Nebular observations of SNe type II (South) | 6.90 HR |
| 3 | GS-2016A-Q-76 | Schlaufman | US | GMOS-S | An All-Sky Search for the Brightest Metal-poor Stars (South) | 74.00 HR |
| 3 | GS-2016A-Q-77 | Origlia | US | Phoenix | MIRA STARS AS PROBES OF THE STAR FORMATION HISTORY OF THE COMPLEX STELLAR SYSTEM TERZAN 5 | 6.00 HR |
| 3 | GS-2016A-Q-78 | Oka | US | Phoenix | Probing the Galactic Center's Molecular Gas using H3+ and CO | 42.60 HR |
| 3 | GS-2016A-Q-79 | Miller | CA | GMOS-S | Globular Cluster Spectroscopy of dEs in the Local Volume (South) | 9.89 HR |
| 3 | GS-2016A-Q-80 | Howell | US | DSSI | Validation, Radius Determination, and Host Star Binarity of K2 Exoplanets | 50.00 HR |
| 3 | GS-2016A-Q-81 | Herczeg | US | GMOS-S | Accretion and the formation of very low mass objects (South) | 5.20 HR |
| 3 | GS-2016A-Q-82 | Fernandes Lopes Soares | BR | GMOS-S | Deciphering the star-formation scenario of the Sh2-296 nebula | 3.60 HR |
| 3 | GS-2016A-Q-83 | de Souza Angelo | BR | GMOS-S | Investigation of Galactic open cluster remnants: the cases of Ruprecht 31 and ESO 570-SC12 | 3.30 HR |
| 3 | GS-2016A-Q-84 | Davidge | CA | Flamingos 2, GMOS-S | Carbon Star Signatures in Integrated Galaxy Light | 34.60 HR |
| 3 | GS-2016A-Q-85 | Cúneo | AR | GMOS-S | Chemical abundances of polluted visible component atmosphere in black hole binary systems (South) | 2.70 HR |
| 3 | GS-2016A-Q-86 | Carrasco | US | Flamingos 2, GMOS-S | Studying the matter distribution in the bimodal lensing group SA78-SA790 | 9.00 HR |
| 3 | GS-2016A-Q-87 | Britt | US | GMOS-S | Spectroscopic evolution of a Black Hole X-ray Transient | 41.00 HR |
| 3 | GS-2016A-Q-88 | Bessiere | CL | GMOS-S | Are all high luminosity AGN triggered at the peaks of major, gas-rich mergers? | 30.00 HR |
| 4 | GS-2016A-Q-89 | Fang | US | GMOS-S | Spectroscopic observations of faint young stars in Orion nebula cluster | 19.60 HR |

Appendix E. Research Staff Effort

Tables E-1 and E-2 list the distribution of effort of staff who have research time allocated. Values are listed by fraction of effort, in the following ten categories. Tables are based on timecard data from 1-Dec-2015 to 30-Nov-2016.

- **night** - nighttime support of regular science operations
- **day** - daytime operations
including queue coordination, routine Head of Science Operations duties, Gemini Science Archive operations, and unscheduled daytime work
- **inst** - instrument support
including instrument maintenance, troubleshooting faults and responding to instrument quality issues, defining calibrations and checkouts, performing nonroutine instrument tests, and instrument documentation
- **user** - user support
including direct program support, visiting observer support, response to Helpdesk, and regular semester activities to allocate programs
- **ops imp** - improvement work and small operations projects
- **res** – research
- **dev** - major development projects
including instrument commissioning
- **trans** - transition projects
- **mgt** – management
- **other ops** - other operations, including scheduled non-project staff meetings, career development and training, and outreach activities

| Staff Member | day | user | res | night | inst | other ops | dev | trans | ops imp | mgt |
|---------------------|------------|-----------|------------|-----------|------------|------------|-----------|-----------|-----------|-----------|
| Andersen, Morten | 8% | 5% | 62% | 8% | 1% | 15% | 1% | 0% | 0% | 0% |
| Angeloni, Rodolfo | 9% | 0% | 74% | 14% | 1% | 3% | 0% | 0% | 0% | 0% |
| Carrasco, Rodrigo | 1% | 4% | 35% | 12% | 19% | 29% | 0% | 0% | 0% | 0% |
| Conn, Blair* | 28% | 16% | 27% | 0% | 0% | 28% | 0% | 0% | 1% | 0% |
| Diaz, Ruben | 2% | 12% | 14% | 10% | 14% | 15% | 34% | 0% | 0% | 0% |
| Garrel, Vincent | 16% | 0% | 8% | 0% | 59% | 6% | 11% | 0% | 0% | 0% |
| Gimeno, German | 17% | 6% | 13% | 9% | 37% | 14% | 4% | 0% | 0% | 0% |
| Hayward, Thomas | 10% | 0% | 19% | 6% | 0% | 6% | 0% | 0% | 58% | 0% |
| Margheim, Steven | 0% | 5% | 10% | 5% | 2% | 56% | 22% | 0% | 0% | 0% |
| Miller, Bryan | 19% | 2% | 38% | 3% | 0% | 16% | 0% | 1% | 21% | 0% |
| Rantakyro, Fredrik | 8% | 5% | 7% | 20% | 21% | 32% | 7% | 0% | 1% | 0% |
| Rutten, Rene | 9% | 4% | 2% | 6% | 0% | 29% | 4% | 2% | 4% | 42% |
| Salinas, Ricardo | 12% | 2% | 70% | 10% | 0% | 5% | 0% | 0% | 0% | 0% |
| Sanmartim, David* | 13% | 0% | 46% | 15% | 0% | 26% | 0% | 0% | 0% | 0% |
| Schirmer, Michael | 14% | 14% | 25% | 7% | 2% | 3% | 0% | 0% | 35% | 0% |
| Sivo, Gaetano | 26% | 1% | 11% | 0% | 23% | 20% | 10% | 0% | 8% | 0% |
| Thomas-Osip, Joanna | 0% | 3% | 4% | 0% | 0% | 19% | 0% | 0% | 3% | 71% |
| Turner, James | 0% | 42% | 19% | 0% | 0% | 21% | 0% | 2% | 14% | 0% |
| Average | 11% | 7% | 27% | 7% | 10% | 19% | 5% | 0% | 8% | 6% |

Table E-1: Gemini South

| Staff Member | day | user | res | night | inst | otherops | dev | trans | opsim | mgmt |
|---------------------------|-----------|------------|------------|-----------|-----------|------------|-----------|-----------|-----------|------------|
| Adamson, Andy | 0% | 0% | 3% | 0% | 0% | 1% | 0% | 4% | 0% | 92% |
| Chene, Andre-Nicolas | 0% | 8% | 16% | 0% | 6% | 16% | 7% | 0% | 47% | 0% |
| Chiboucas, Kristin | 35% | 9% | 21% | 4% | 11% | 9% | 2% | 0% | 8% | 0% |
| Geballe, Thomas | 4% | 7% | 51% | 22% | 4% | 10% | 0% | 0% | 2% | 0% |
| Gomez, Percy* | 0% | 23% | 30% | 13% | 17% | 12% | 5% | 0% | 0% | 0% |
| Guyon, Katherine* | 9% | 27% | 5% | 24% | 9% | 12% | 8% | 0% | 7% | 0% |
| Jorgensen, Inger | 0% | 0% | 50% | 0% | 0% | 1% | 0% | 12% | 0% | 37% |
| Kleinman, Atsuko N. | 23% | 2% | 17% | 4% | 0% | 12% | 0% | 43% | 0% | 0% |
| Kleinman, Scot | 0% | 0% | 3% | 0% | 0% | 1% | 0% | 0% | 0% | 96% |
| Labrie, Kathleen | 1% | 30% | 14% | 0% | 0% | 21% | 5% | 3% | 26% | 0% |
| Lai, Olivier* | 0% | 0% | 41% | 0% | 5% | 8% | 46% | 0% | 0% | 0% |
| Leggett, Sandra | 20% | 0% | 40% | 0% | 0% | 21% | 0% | 0% | 0% | 18% |
| Lemoine-Busserolle, Marie | 30% | 13% | 25% | 0% | 13% | 16% | 0% | 1% | 1% | 0% |
| Lundquist, Michael | 8% | 11% | 47% | 7% | 2% | 4% | 21% | 0% | 1% | 0% |
| Mason, Rachel* | 1% | 5% | 27% | 0% | 0% | 33% | 3% | 0% | 31% | 0% |
| Peck, Alison* | 0% | 31% | 13% | 0% | 0% | 3% | 53% | 0% | 0% | 0% |
| Petric, Andreea* | 18% | 17% | 38% | 2% | 4% | 2% | 0% | 0% | 18% | 0% |
| Scharwaechter, Julia* | 6% | 13% | 22% | 12% | 14% | 29% | 2% | 0% | 3% | 0% |
| Schwamb, Megan* | 13% | 15% | 14% | 20% | 19% | 20% | 0% | 0% | 0% | 0% |
| Shih, Hsin-Yi* | 14% | 7% | 54% | 7% | 14% | 2% | 0% | 0% | 2% | 0% |
| Simunovic, Mirko* | 0% | 0% | 27% | 0% | 0% | 73% | 0% | 0% | 0% | 0% |
| Stephens, Andrew | 15% | 14% | 7% | 11% | 11% | 9% | 0% | 1% | 32% | 0% |
| Trujillo, Chadwick* | 0% | 0% | 30% | 0% | 1% | 0% | 0% | 0% | 0% | 68% |
| Average | 9% | 10% | 26% | 5% | 6% | 14% | 7% | 3% | 8% | 14% |

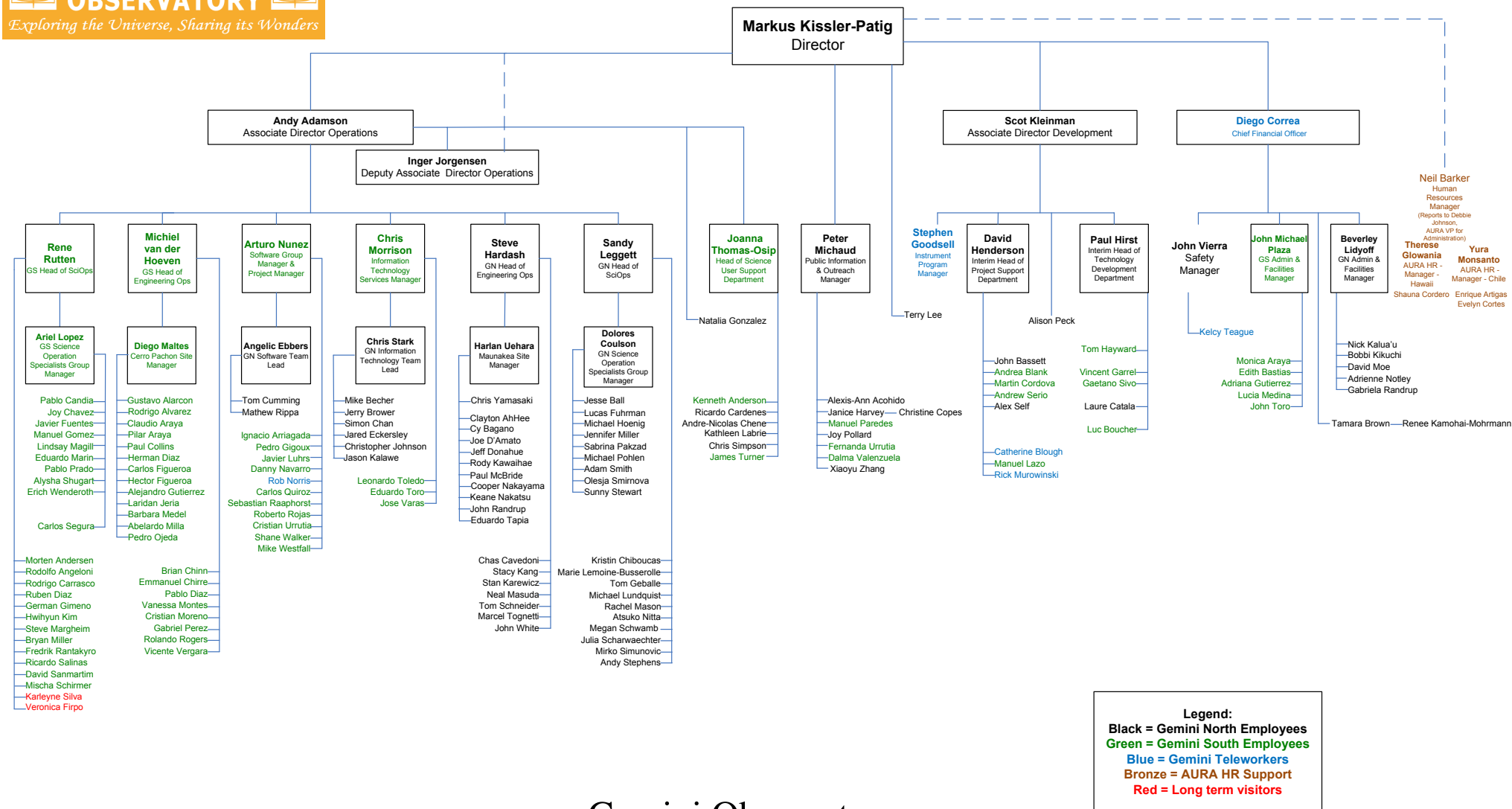
Table E-2: Gemini North

Appendix F. Organizational Chart

Significant organizational changes since the 2015 Annual report include (i) the creation of the Project Support Department, consolidating the previous Systems Engineering Department and project managers who previously reported directly to the AD for Development; (ii) the replacement of the AO and Telescope Development Department with Technology Development Department. Two temporary arrangements are in place pending recruitment of the Chief Scientist/Deputy Director: the Science User Support Department reports within Operations and PIO reports to the Director.

See the following page for the organizational chart.

Exploring the Universe, Sharing its Wonders



Gemini Observatory
Effective as of December 1, 2016

Appendix G. Report Requirements

Guide to Locating Material According to Cooperative Agreement Terms and Conditions

| Requirement | Description | Fulfillment |
|-------------|--|-----------------------|
| a.i | Summary of achievements, including a comparison of actual accomplishments versus goals | 4,5; Appendix I |
| a.ii | Identification of problems faced, their solutions and impact on observatory operations | 4,5 |
| a.iii | List of observing programs, with their investigators, site visitors, observers, and hours devoted to each | Appendix D |
| a.iv | Report on the education and public outreach activities including non-scientific visitor statistics, press releases, etc. | 6, Table 6-1 |
| b. | Scientific accomplishments of the scientific staff, as well as their activities and expectations for the calendar year | Appendices B, E |
| c. | Technical accomplishments of each technical department, as well as the departments' expectations for the calendar year | 4, 5, 8 |
| d. | Listing of publications and reports produced by staff and, as far as possible, the users of the observatory | Appendices B, C |
| e. | Table showing the division of effort, adding up to 100 percent, for all scientific staff and/or Key Personnel among major activities, such as administration, visiting observer support, scientific research, etc. | Appendix E, Table E-1 |
| f. | Chart or other description of Gemini's organization during the new program year together with an explanation for any changes from the previous year's organization | Appendix F |
| g. | Status report and plan for the new program year on the state of the Observatory | 8 |
| g.i | Scientific plans | 8.2, 8.3, 8.4 |
| g.ii | Detailed technical plans and the potential impact on the observatory of these plans | 8.2, 8.3, 8.4 |
| g.iii | Plans with schedules, milestones, and principal activities, for all major projects | 8 |
| g.iv | Estimated budgets for all major activities | 8.1 |

Appendix H. Metrics

| No. | Metric | Fulfillment | Target | Result | | | |
|-----|---|---|-------------------------------------|---|-------------|-------------|-------------|
| | | | | GN 2015B | GN 2016A | GS 2015B | GS 2016A |
| 1 | Telescope time use (per telescope, per semester), identifying fraction of time for science, weather loss, commissioning, telescope system faults, and instrument faults | 4.7.1 Tables 4-4, 4-5 * See text for explanation of major downtime. | <4% fault loss | 3% | 3% | 2.80% | 1.50% |
| 2 | Over-subscription rate (per telescope, per semester) | 4.7.2, Fig. 4-2 | | See Fig. 4-2 | | | |
| 3 | Queue program completion fraction, by band (per telescope, per semester) | 4.7, Fig. 4-1 | Band 1 Complete: 75% | 70% | 77% | 30% | 22% |
| | | | Band 2 Complete: 60% | 61% | 74% | 11% | 26% |
| | | | Band 3 (started) at 75% complete | 74% | 68% | 64% | 50% |
| 4 | List of staff research and technical achievements, including prizes, awards, and invited talks | Appendix E | | | | | |
| 5 | List of staff technical achievements, including prizes, awards, and invited talks | Appendix K | | | | | |
| 6 | Number of peer-reviewed publications based on Gemini data | 3.1, Appendix C | 190/year | 188 (72 based on GN, 92 based on GS; 24 based on combined observations from both) | | | |
| 7 | Number of peer-reviewed publications with Gemini Observatory staff as co-authors | Appendix B.1 | 20/year | 28 | | | |

Appendix I. 2016Q3 Financial Report

See attached.

Appendix J. Completion of 2016 Program Plan

The original 2016 Program Plan is included here, excerpted for brevity, and annotated with specific comments on progress in [blue](#). We do not list optional projects undertaken on a best-effort basis, nor do we list additional activities and accomplishments during 2016. The main report body provides more explanation of these activities.

J.1 Science and Engineering Operations in 2016

J.1.1 Regular Operations

Here we list regular operations items with specific developments within the year.

- Maintain the instruments and telescopes in working order consistent with the requested science time on sky; monitor performance and take remedial action as needed.
 - Remedial action on NIRI (cold heads, mechanisms) was completed. Remedial action on F-2 cold heads was completed. Remedial action on GMOS-S optics oil bubbles now planned for 2017.
- Run the International Time Allocation Committee (ITAC) process to combine the national TAC results into an executable queue and visitor program consistent with available time, conditions, and instrumentation.
 - ITAC chair duties passed to Rodrigo Carrasco (Gemini South).
- Provide web-based documentation suitable for PI reference on instrumentation, software and Observatory processes.
 - See Science User Support
- Support visiting instruments as needed and as possible.
 - Gemini South saw both DSSI and Phoenix as visiting instruments in 2016A.
 - GRACES saw significant use at Gemini North. See the run report web page for details⁹.
 - Worked with the DSSI team to develop a more robust version of the instrument which may be possible to permanently mount on Gemini North. At the time of writing the instrument is in the build phase and we are discussing options for Gemini South.
- Propose and execute continual improvements in instrumentation, telescope, and enclosure to maintain performance levels.
 - Vibration monitoring equipment similar to that in the North is working well and being readied for installation on Gemini South. This requires a 2017 project in operations to complete. Once done, we will retrofit the computer and software to the North.

⁹ <http://www.gemini.edu/sciops/instruments/graces/status-and-availability/end-graces-schedule-block-report>

- Provide expertise and input to the Development Division in carrying out major enhancements of instrumentation.
 - Done; various engineering and science input provided to multiple projects.
- Staff the “third” and final level of a helpdesk to respond to queries from the user community. (The first two levels are (i) NGOs and (ii) instrument specialists at the NGOs.)
 - Done with some changes in the detail, e.g. US support of cookbooks

J.1.2 Science and Engineering Operations Core Projects and Goals

Develop an Observatory-wide Obsolescence Mitigation Plan

- Initial plan was completed and submitted to NSF by August 2016. Details and timing are being worked on and will be complete by end of 2016.

Operations Software Upgrades and Obsolescence Management

- A prototype of a web-based user interface for the new sequence executor was completed in May. Progress on the execution engine was slower than planned due to resource unavailability during Q3. A first operational version of the system is now expected by Q2/2017.
- Due to staff availability, the TCC rewrite will begin once the Seqexec is complete, namely in 2017.
- RT Upgrades: We completed the consolidation of common software libraries for the real-time systems, with a final review in October. With this in place, the upgrade stage has started. We plan to complete the upgrade of the simpler systems (GWS, A&G Sequencer, GCAL) by end of 2016, as originally planned.

Science Operations Model Upgrade

- Prioritization document was in fact completed within Q3. A set of priorities were agreed with the Director and we are now working to fit them into the overall program.

Upgrade the Cerro Pachón Network Link

- Microwave connection was fully implemented and has been in reliable operation, adequate for base facility operations, for many months. We now expect the LSST fiber to become available in early 2017 so this item will recur in the 2017 program plan.

Commission the FLAMINGOS-2 MOS mode

- This remains to be completed. Work was completed on the On-Instrument Wavefront Sensor (OIWFS) on schedule, but subsequent weather issues and a new fault have so far prevented any progress on the MOS commissioning.

J.1.3 Science User Support Department in 2016

- Implement new post-observing communication strategy, including regular support for the Data Reduction User Forum, post-observing contact with PIs, and improvements to the regular support mechanism that the Helpdesk currently provides.
 - Technical and structural improvements (Q1)
 - completed

- Utilize dedicated non-staff contributors to Forum (Q2)
 - completed
 - Implement post-observing contact with PIs (Q3)
 - delayed to 2017 in favor of starting project to develop new public web pages with user-centric navigation and structure
 - Define requirements for external Helpdesk (Q4)
 - completed
- Improve data reduction software documentation and cookbooks.
 - Evaluate inventory and cookbook template (Q1)
 - Cookbook template delayed to prioritize a successful advertisement campaign for the Data Reduction Users' Forum. This work to be completed in Q4.
 - Define specific plans for cookbooks and documentation (Q2)
 - completed
- Complete final imaging mode of QAP (GSAOI; Q1) and release to public (Q2).
 - completed except public release delayed to 2017 in order to provide software useful for scientific quality reduction
- Complete quick-look tool for quality assessment of spectroscopic observations (Q1).
 - completed
- With STScI, repackage the STScI- and Gemini-developed software within Ureka using Conda (Q2). Use of Conda will allow the user to stay more up-to-date with third-party Python packages such as numpy and matplotlib (among hundreds of others).
 - behind schedule; will be completed in Q4.
- Update Gemini IRAF for new GMOS-N Hamamatsu CCDs (Q3) and F2 MOS mode (Q4). This will be one of the last releases of Gemini IRAF.
 - The Gemini IRAF updates for GMOS-N and F2 MOS will be delayed into 2017 due to delays in commissioning of both modes.

J.2 Transition Program in 2016

Base Facility Operations at Gemini South

- Base facility operations started on scheduled in mid-November at Gemini South

GN Energy Savings

Execute large contracts: (A) related to the Gemini North Energy Audit; and (B) installation of HBF photovoltaic system.

(A) The following energy savings measures and replacements of cooling equipment are included in the energy-audit-related contract.

1. Lighting and energy management systems improvements, including replacement of HBF and MK lighting with LEDs
2. Installation of fluid cooler (Maunakea)
3. Replacement of transformers with premium efficiency models
4. Installation of variable frequency drive pump (Maunakea)
5. Replacement of Maunakea chiller
6. HBF HVAC upgrades and replacements

- (A) The contract for implementation of the content of the Gemini North Energy Audit was about 4 months behind schedule in October, due to delays on the contractor's side involving late start of engineering design, delays in hiring contractors, and delays in providing work schedule for the major components. We have taken action to pull out of the current contract as agreed with the contractor. We will complete the remaining work with smaller direct contracts. This construction work will go into 2017. We expect to realize \$85k out of the total planned \$193k non-labor savings from the project by the end of 2016. The delay in realizing the rest of the savings has been taken into account in the 2017 budget.

We provide an overview of the work here.

| Energy Conservation Measures | ECM #s | Status | Main reasons for implementation |
|---------------------------------|--------|---|--|
| Lighting at HBF & MK | 1,2,3 | Expected complete by end of Dec 2016 (6 month delay, plus missed incentive funds of \$4.4k) | Energy savings; eliminate mercury in CFL lights. |
| HBF energy management | 4 | Will be descoped. | Energy savings |
| Step down transformers | 8 | HBF + MK install complete by mid-September (2.5 month delay) | Energy savings |
| MK fluid cooler | 5 | 100% engineering design is in place. Requests for county permits submitted. Fluid cooler and Chiller purchased. Construction to be under separate contract. | Energy savings; enable cheaper solution on replacement of aging chiller. |
| Pump – variable frequency drive | 9 | | Energy savings |
| MK chiller replacement | 6 | | Replace aging equipment; meet Federal requirements to phase out HCFC-22 refrigerant. |
| HVAC upgrades (HBF) | 7 | 100% engineering design is in place. Requests for county permits submitted. Construction to be under separate contract. | Replace aging equipment; meet Federal requirements to phase out HCFC-22 refrigerant. |

(B) Our contractor has finalized the design of the 100kW HBF photovoltaic system and has submitted it to the County of Hawaii for a building permit. We expect construction to start in early February and complete by mid-2016, weather and inspections permitting.

- (B) The HBF photovoltaic system was completed in July 2016.

GS Energy Savings

Install 200kW+ photovoltaic system at Cerro Pachón (Q3, weather and inspections permitting).

- The CP photovoltaic system was completed in July 2016.

Reduce Base facility Expenses

Modify SBF to create 15 offices and new lab space. Vacate Casa 8 and reduce usage of Casa Verde. (Q3)

- Construction and furnishing as well as move to the new office space was completed in early December.

J.3 Instrumentation and Facility Development in 2016

Principal projects are listed in the table below.

| Project | 2016 Plan |
|---------------------|---|
| GHOST | <ul style="list-style-type: none">• Start Build Phase<ul style="list-style-type: none">◦ Done May 2016 |
| Gen4#3 | <ul style="list-style-type: none">• Start contract negotiations with selected teams<ul style="list-style-type: none">◦ Released RfP in May, 2016. Proposals received in August; submitted section report to Board subcommittee in October. |
| GeMS Laser | <ul style="list-style-type: none">• Select and complete contract for purchase• Start work for necessary internal modifications for new laser<ul style="list-style-type: none">◦ Done. Vendor contract signed in Q1. Internal work proceeding. Vendor delivery expected by end of year. |
| GMOS CCDs | <ul style="list-style-type: none">• Install and commission new CCDs into GMOS-N<ul style="list-style-type: none">◦ GMOS-N installation delayed due to technical problems with the “duplicated” components from the GMOS-S installation. We resolved all technical issues by May 2016 and are testing the science CCDs in October, for installation in early 2017. |
| Instrument Upgrades | <ul style="list-style-type: none">• Select and begin project(s) from 2015 call• Launch 2016 call with expanded budget<ul style="list-style-type: none">◦ The 2016 project is underway and we released the RfP for the 2017 projects in September 2016. |
| NGS2 | <ul style="list-style-type: none">• Receive, install and commission NGS2<ul style="list-style-type: none">◦ Project delayed. Descoped acceptance testing planned for December 2016 with plans for further work to be evaluated afterward. |

The high priority projects for Development and their planned activities in 2016.

J.4 Administration & Facilities and Safety in 2016

J.4.1 Finance and Administration

Budgetary responsibility

Finish 2016 within [-2%; +3%] of the requested O&M Budget (\$27.30M). Continue to promote the accountability among the budget account managers.

- Training on financial planning and the Gemini budget system CONTROL was delivered to Gemini managers in Q1-2016 and Q3-2016. The 2016 budget and Forecast analysis presented in the table below show a net value in 2016 budget changes of \$0 (Zero USD\$) and the 2016 budget is expected to be spent within the [-1%, 2%] range.

| <i>in \$ 000</i> | 2016 Budget | 2016 Forecast | Var 2016 Budget vs. Forecast\$ | Var 2016 Budget vs. Forecast % |
|-----------------------------|-------------------|-------------------|-----------------------------------|-----------------------------------|
| Salaries | 12,869,409 | 12,429,645 | (439,764) | -3.5% |
| Benefits | 4,441,007 | 4,321,789 | (119,218) | -2.8% |
| Temporary | 429,216 | 569,386 | 140,170 | 24.6% |
| Total Labor | 17,739,633 | 17,320,821 | (418,812) | -2.4% |
| Supplies Equipment | 1,089,041 | 1,322,320 | 233,279 | 17.6% |
| Travel | 952,791 | 978,974 | 26,183 | 2.7% |
| Recruiting Relocation | 58,000 | 58,000 | - | - |
| Professional Fees | 1,531,005 | 1,545,497 | 14,492 | 0.9% |
| Meeting, Conf., Prof Dev. | 311,833 | 314,053 | 2,220 | 0.7% |
| Computer SW. and Equip | 539,854 | 612,314 | 72,460 | 11.8% |
| Facilities | 649,134 | 646,124 | (3,010) | -0.5% |
| Maintenance | 182,069 | 182,069 | - | - |
| Utilities | 1,298,602 | 1,258,802 | (39,800) | -3.2% |
| Meals and Lodging | 366,694 | 410,494 | 43,800 | 10.7% |
| Total Site Costs | 2,496,498 | 2,497,488 | 990 | 0.0% |
| Spares | 370,810 | 247,539 | (123,271) | -49.8% |
| Other | 502,036 | (8,712) | (510,748) | 5862.9% |
| Indirect Costs | 986,169 | 836,169 | (150,000) | -17.9% |
| Subcontracts | 614,829 | 1,091,972 | 477,143 | 43.7% |
| Total Non-Cap. Exp. | 27,192,500 | 26,816,435 | (376,065) | -1.4% |
| General | 107,500 | 483,565 | 376,065 | 77.8% |
| Total Cap. Exp. | 107,500 | 483,565 | 376,065 | 77.8% |
| Total Non-Labor Exp. | 9,560,367 | 9,979,179 | 418,812 | 4.2% |
| TOTAL EXPENDITURES | 27,300,000 | 27,300,000 | 0 | 0.0% |

Integration of Shared Services

Support strong partnership with AURA CAS and other Administration and Facilities services providers (NOAO-S and MKSS) With CAS, identify risks and their mitigation in the domain of accounting and contracting. Review and revise the NOAO-S Service Level Agreements for 2017 in view of BFO and power savings.

- Active participation in NOAO-S and MKSS Oversight Committees (no participation in the CAS/HR oversight)
- Increasing budget of NOAO-S for FY2016 of \$150k to support Chile facilities' infrastructure improvements
- MKSS increasing shared fee of \$32k for FY2016-2017 to offset decreasing lodge & meals fees
- Additional CAS and HR staff to cause increasing FY 2017 fees of \$50K
- Gemini is participating on the NOAO-S FO (NOAO-S Facilities Operations) working groups that will review and update the Service Level Agreements
- Gemini is working jointly with other AURA centers on the BSR Working groups that aim at unifying business systems, practices and procedures.
- Gemini, CAS and HR are working jointly on the plan of the coming Business Systems review of the Gemini Observatory that NSF will conduct in early 2017.

Instrument Development Fund management

Develop a long term Instrumentation reporting system that meets IDF reporting needs of management, governance, and future Partners.

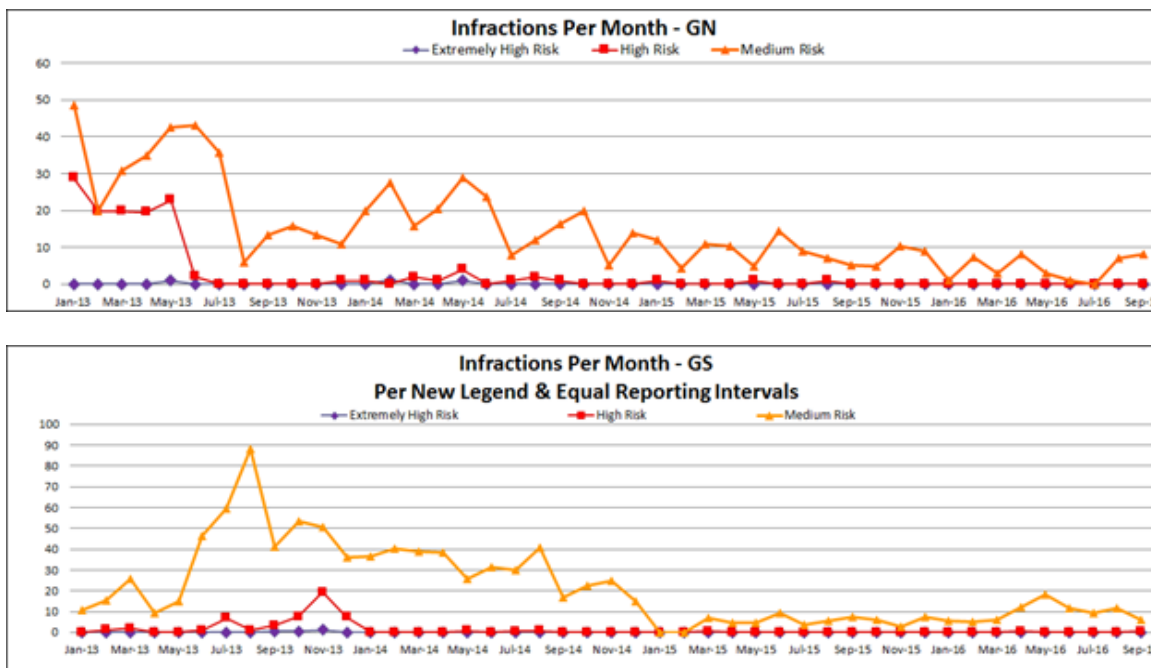
- New reports created in 2016 to track IDF's current and future spending of instrumentation projects. Ongoing efforts to develop financial reports to track IDF contributions by project.

J.4.2 Safety

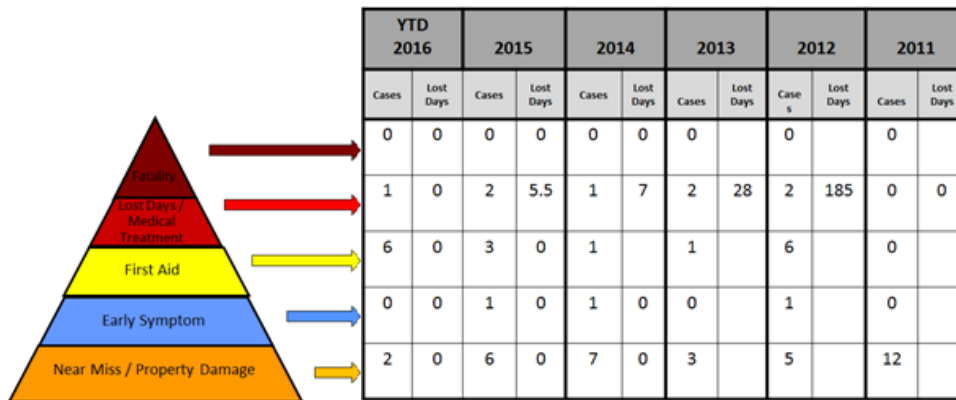
Staff Safety

Continue delivering Safety standards and services at the highest level, and meet the goal of an observatory-wide 90% completion of the mandatory safety training.

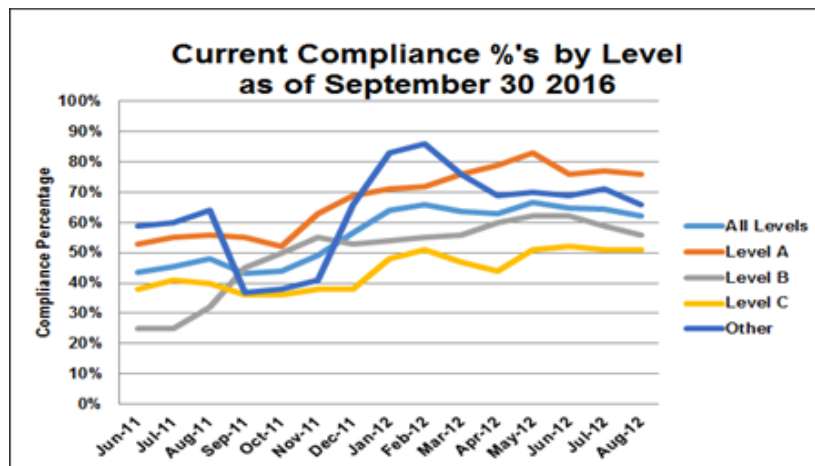
- Goal: *Mitigate Gemini's driving/speeding high risk.*
High risk driving violations are nil and the following charts depict the decreasing trend of speeding violations in Gemini North and Gemini South for the period Jan 2013 to Sep 2016:



- Goal: *Minimize the number of Safety incidents and time lost*
Time lost in 2016 is zero; the chart below shows the decreasing trend of safety incidents and lost days since 2012 to date:



- Goal: Achieve a training ratio of 90% in 2016.
Gemini has achieved an Observatory-wide completion of the mandatory safety training of 78% as of September 30, 2016. The following graph shows the increasing trend of safety training completed by all Gemini staff:



Safety operations and Management structure

Implement long-term Safety operations and management structure around the synergies with Maunakea and Cerro Tololo/Cerro Pachón partner telescopes.

- Throughout 2016 the Safety team has been working on a long-term Safety operations and management structure around the synergies with Cerro Tololo/Cerro Pachón partner telescopes. Safety resources were initially shared with LSST in the South and after the transfer of the Safety Coordinator to LSST in August 2016, the Safety department has been restructured. We are recruiting a Safety Specialist for Gemini South, and retained the staffing services of an Admin Safety officer until December 31, 2017. Ongoing joint training efforts and sharing of Safety resources are ongoing with Maunakea telescopes.

J.5 Public Information and Outreach in 2016

We will continue our regular local outreach programming, including Journey Through the Universe (JTJU), Viaje, AstroDay, StarLab, media relations, publications, user communications support, library services, and science result tracking.

- JTJU was completed, with significant changes to program based on input from

- teachers and Department of Education administrators. (100% completed)
- Viaje was implemented in late October, on schedule. (100% completed)
- Media relations are on-track and 12 press releases were produced exceeding our minimum goal of 8. (150% completed)
- Publications are all either on schedule or published. (100% completed)
- All other items are on target. (100% completed)

We also identify the following specific activities to complete in 2016.

- Unify user communications at both sites by introducing regular processes and balance quantity of content for users and the public.
 - Implemented 50:50 social media posts (50% user, 50% public and of these 50% for each site) and tracking is now part of normal operations. (100% completed)
- Expand media relations resources and public relations capabilities, with annual public information internship (beginning Q3 - hired through March 2017) and limited-term (2-year), entry-level, public information assistant position (beginning Q1 - hired, as 3-year position). Continue limited engagement in Maunakea/TMT and related crisis communications, including ongoing dissemination of information to key stakeholders. (100% completed)
 - Activities related to Maunakea and TMT are reduced due to delayed court proceedings, but supporting “Envision Maunakea” initiative with commitment to assist in communications effort.
- Complete integration of common outreach objectives for core Gemini South outreach programming. Transition StarLab in Chile to “train and loan” model (as done in Hawaii) by June. (100% completed)
 - StarLab training and loan program established at GS, Viaje al Universo formal implementation plan developed in alignment with Journey Through the Universe program at GN
- Complete addition of eight new staff profiles in Career Brochure and post video interviews on companion website (Q2). Produce a minimum of four additional 8.5x11” in-depth profile sheets (Q2). (90% completed as of November 30, 2016, on track for 100% completion by end of 2016)
 - New versions of career brochures (English and Spanish) completed with two 8.5” by 11” in-depth sheets complete; remaining sheets will be completed by the end of the year.
- Complete tablet-based virtual tour production and install in both Gemini lobbies (Q3).
 - In production at GS; GN content almost complete, 80% at end of November, will be 90% completed by end of 2016.
- Develop and implement new publication tracking procedures that include improved data on archival use of Gemini data (Q2). 100% completed
 - Completed by library staff

J.6 Gemini External Relations in 2016

- North American institutional visits: Director Road Trip for general information about Gemini’s new operations, instruments, and opportunities (Q1; completed late Aug/early Sept); Development Road Trip with focus on instrument-building institutions (Q3; cancelled)
- Directorate attendance at Partner national astronomy meetings
 - Several members attended the American Astronomical Society meeting in

- January 2016. The Deputy Director attended SOCHIAS in March, Director attended CASCA in June, Deputy Director attended SAB in August, and Assoc. Director for Operations attended AAA in September
- Support Gemini Board in development of strategic vision exercise
 - Ongoing, report to be delivered at the November 2016 Gemini Board meeting
 - Lead the development of memoranda of understanding with any new limited-term Partners
 - Agreements signed with KASI (for 2017 & 2018) and AAL (for 2017)

Appendix K. Staff Accomplishments

K.1 AURA Awards

- The 2016 AURA Science Award went to Rachel Mason, tenured associate astronomer, and Jared Eckersley, web applications developer, for the creation of Gemini's Fast Turnaround Program. Fast Turnaround (FT) utilizes peer review to enable a community scientist to go from an idea to delivered data in as little as one month. Currently up to 10% of Gemini observing time may be distributed through this mode.
- The 2016 AURA Service Award went to John Michael Plaza, Gemini South Administration and Facilities Manager, for consistently meeting and managing the service expectations of Gemini's staff and following through on commitments accurately and in a timely manner.
- The 2016 AURA Technology/Innovation Award went to Paul Hirst, observatory scientist, for envisioning, designing and implementing the new cloud-based Gemini Observatory Archive. The new archive uses Cloud-based data storage (purchased from Amazon Web Services) and interfaces live with our observing software at the two telescope sites. New data are transferred in real time and made available to the users within minutes of being obtained.

K.2 Invited Talks

André-Nicolas Chené

"Massive infrared clusters in the Milky Way", at the IAU Symposium 329: The Lives and Death-Throes of Massive Stars, Auklan, New-Zealand (November 30, 2016).

Tom Geballe

"Observations of Ices in the Galactic Center: Past, Present, and JWST-Future," at Leiden University / "Ice Age - The Era of the James Webb Space Telescope" at Leiden, The Netherlands, (October 7, 2016)

Atsuko Nitta Kleinman

"From Theoretical Particle Physics to Observational Astrophysics" at the Final Lecture Series on Prof. Akio Sugamoto's Retirement, Ochanomizu University, Tokyo, Japan (March 05, 2016)

Megan Schwamb

"Planet Four and Planet Four: Terrains," Our Red Planet Workshop, Greenbelt, Maryland, USA (September 20, 2016).

"Colours of the Outer Solar System Origins Survey (Col-OSSOS): New Insights into Kuiper belt Surfaces," AGU, San Francisco, USA (December 15, 2016).

Gaetano Sivo

"Wide-field Adaptive Optics for Astronomical Ground-based Telescopes: Science Results and Ongoing Upgrades for GeMS", at the Optical Society of America conference, Heidelberg, Germany, July, 27, 2016.

K.3 Special Accomplishments

Atsuko Nitta Kleinman

Listed in "The World's Most Influential Scientific Minds 2015" by Thomson Reuters.